

# Motto

Hamlet Act. I,  
scene 5

There are more things in heaven and earth,  
Horatio,  
Than are dreamt of in your philosophy

# GRAVI at DESY

A progress report

# GRAVI

An experiment to test NEWTON

in the laboratory

at very small acceleration

with

W. Bartel, A. Glazov, B. Lühr,

C. Niebuhr, S. Schubert, E. Wunsch

DESY

E. Lohrmann, Univ. Hamburg

H. Meyer, Univ. Wuppertal

# OUTLINE

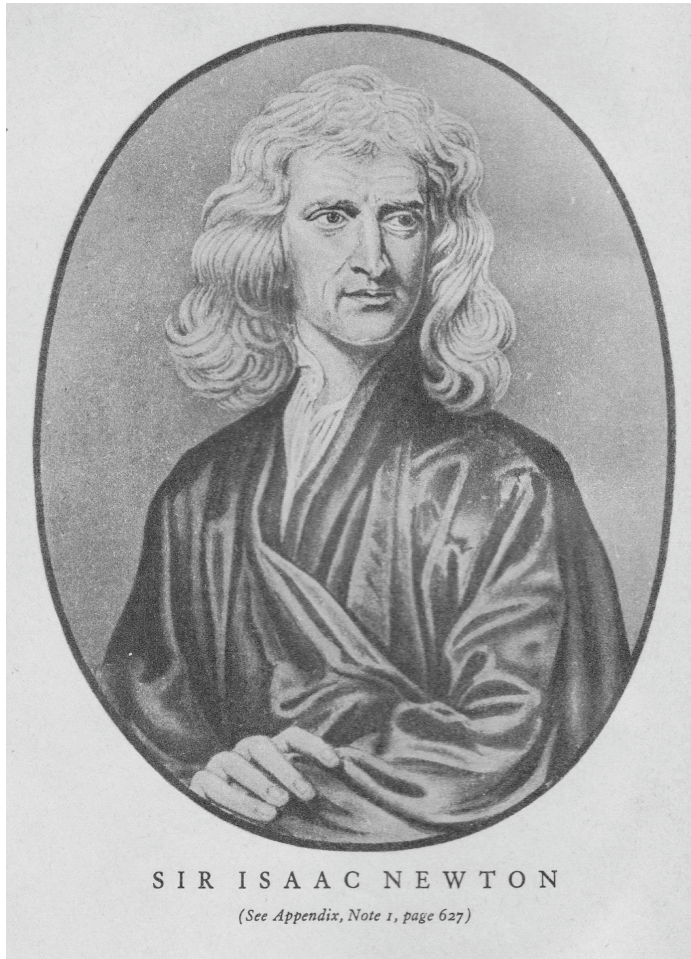
- Motivation
- Big G at Wuppertal University
- Setup at DESY
- Data taking
- Preliminary results

# Motivation in Wuppertal

In the early 1980 there was a discussion on the possible existence of a 5th force, which should be detectable by a deviation from the  $1/r^2$  law of NEWTON

But we decided to go for a precise determination of the gravitational constant  $G$

# Newton

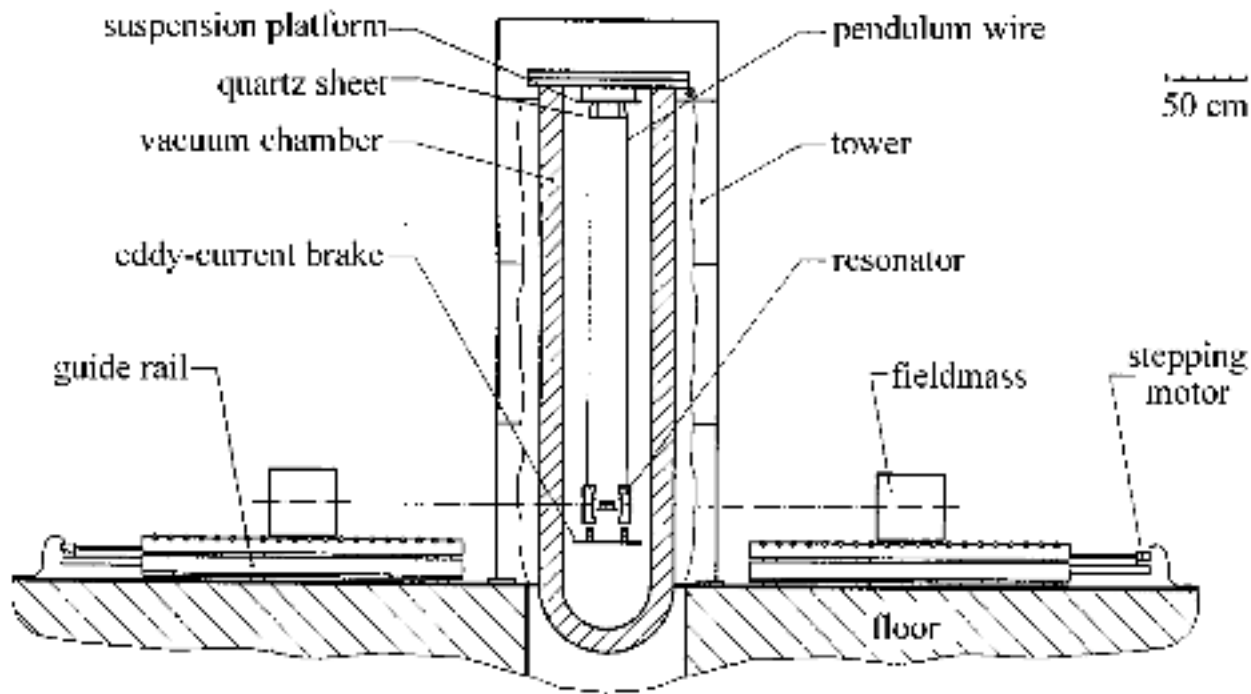


$$F = G M m / r^2$$

Aim:

determine  $F$  by shifting  $m$   
through the movement of  $M$

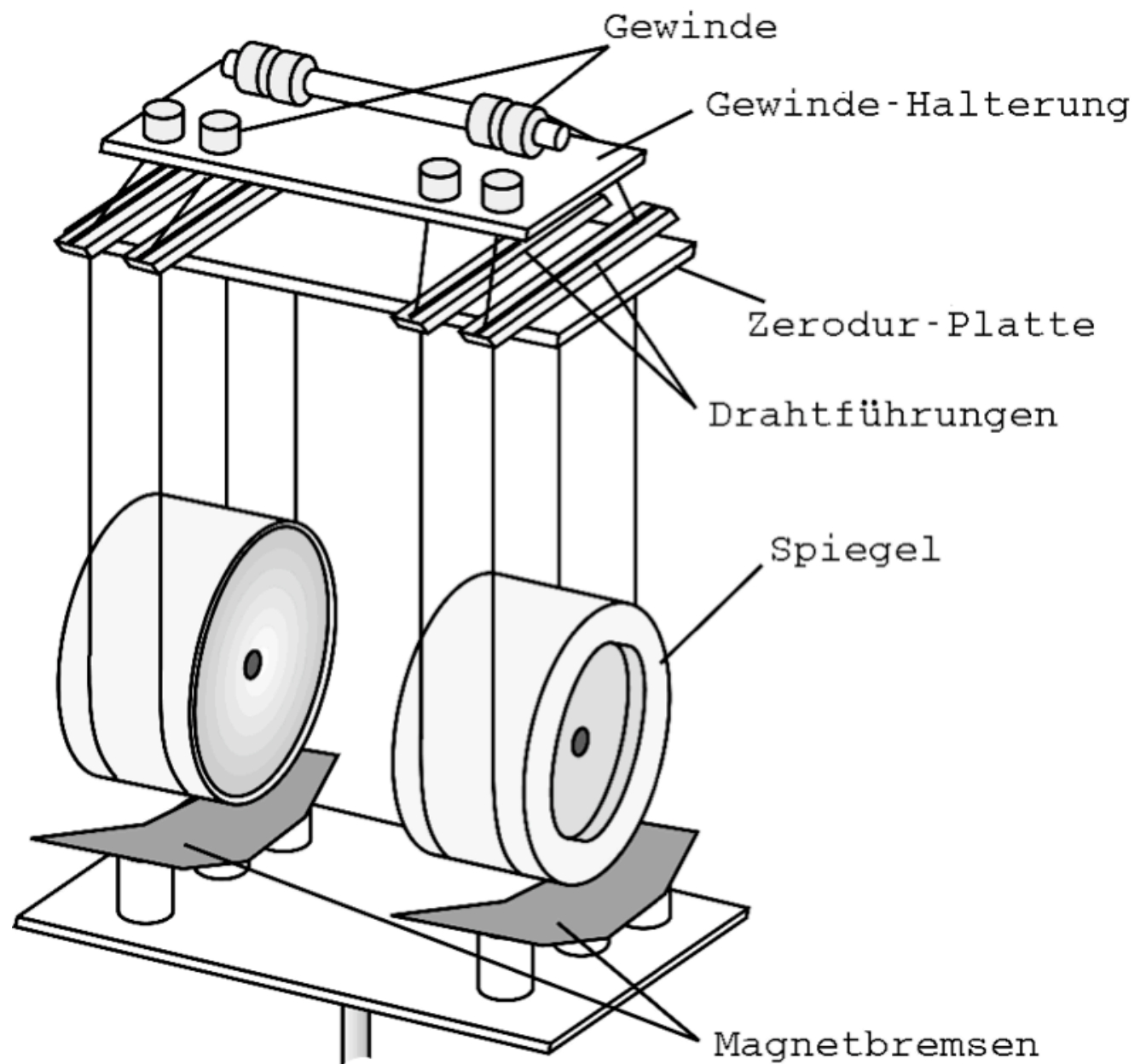
Calibrate  $M, m, r$  very precisely



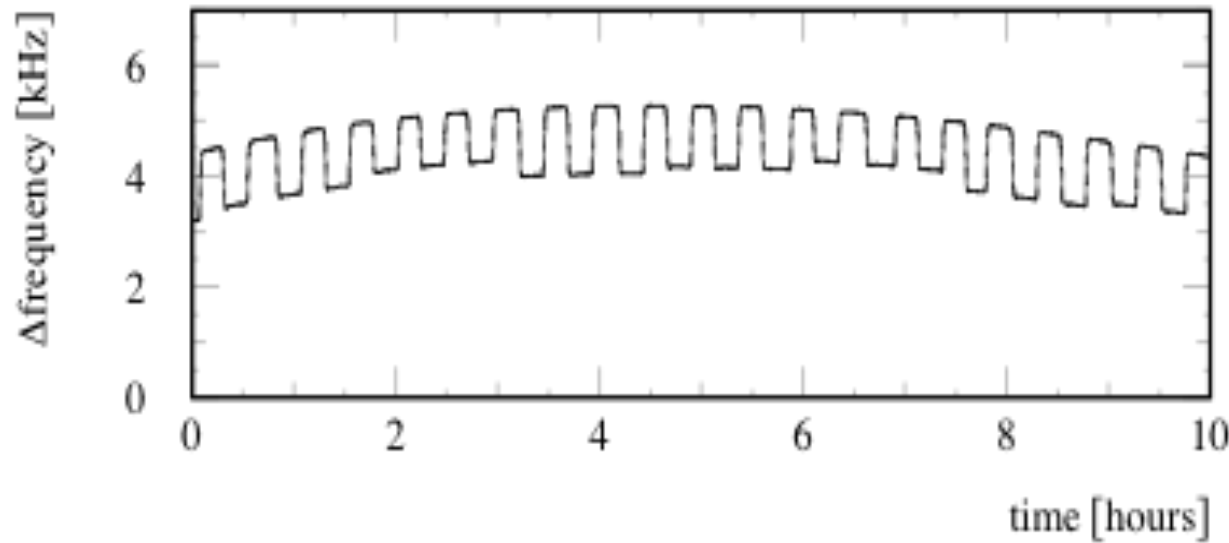
Fieldmass  
 $M = 2$  times  
 560 kg  
 small  $m$   
 a “cavity”

**Figure 1.** Schematic view of the experimental set-up with the Fabry–Pérot resonator and the two fieldmasses.

thesis work by  
 J.Schurr, H.Walesch, A.Schumacher, U.Kleinevoss  
 1988 - 2002  
 with H. Piel  
 Univ. of Wuppertal

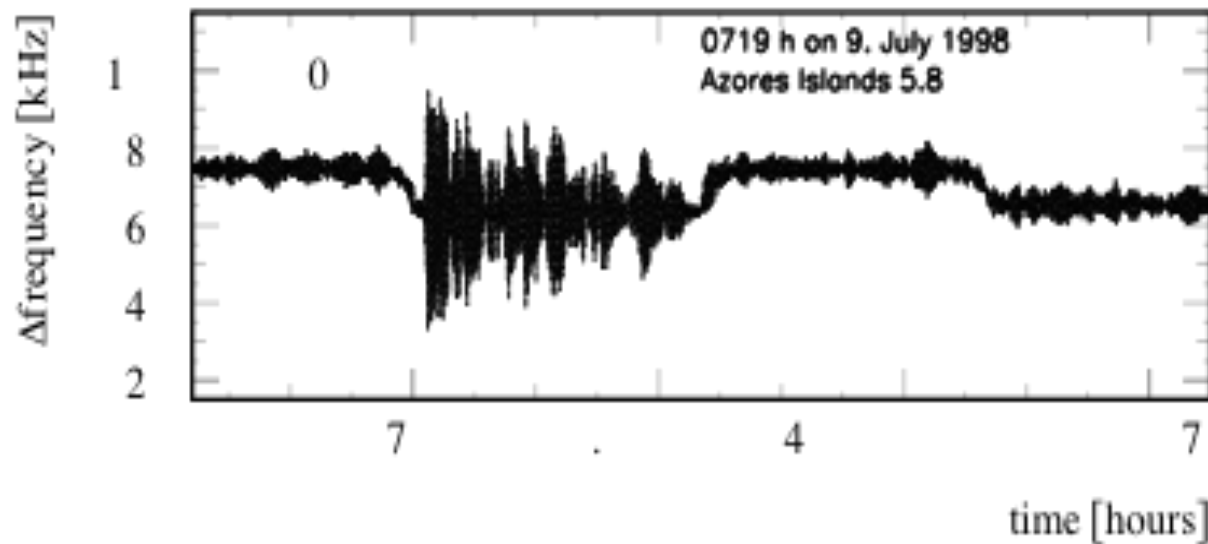






$$db = df \text{ const}$$

$$\text{const} = 10^{-8} \text{ m/kHz}$$



thesis  
H. Walesch

**Figure 2.** Measured resonance frequency change  $\Delta f$  plotted against time. The upper trace shows a 10 h portion of the measurements and lower trace shows a rather large earthquake.

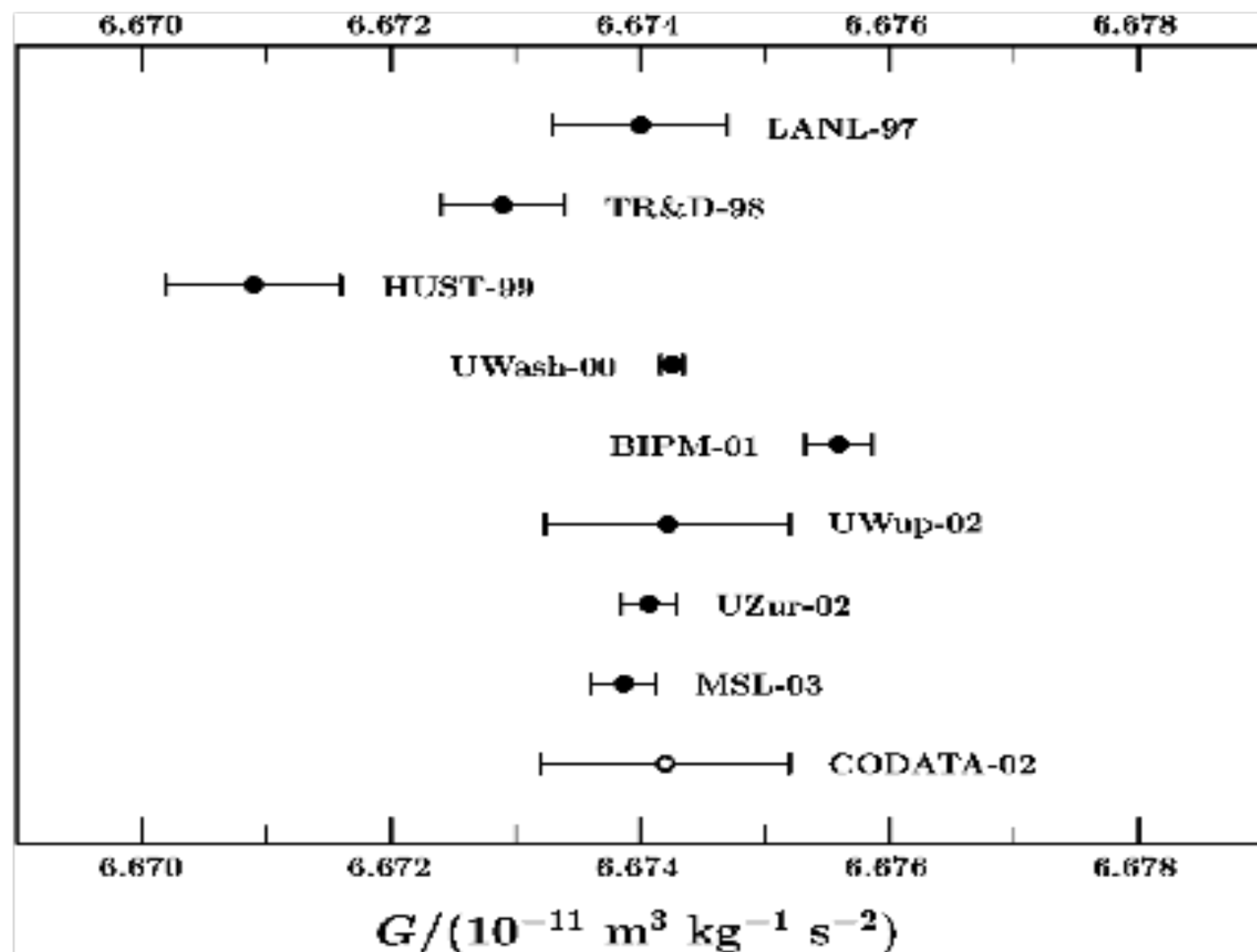


FIG. 1. Values of the Newtonian constant of gravitation  $G$ . See Glossary for the source abbreviation.

## Goals with “GRAVI” at DESY

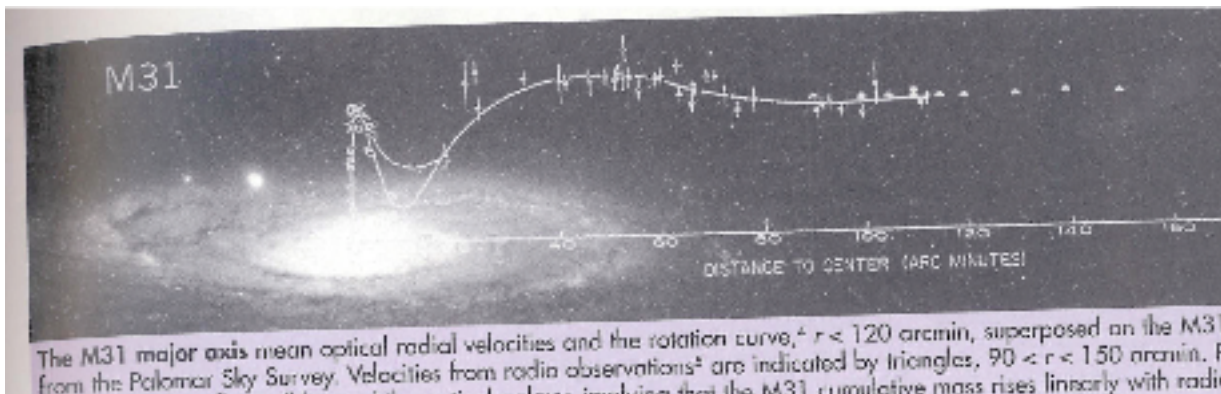
1: measurements in the  
MOND region at  
very small acceleration about  
 $10^{-10}$  m/sec<sup>2</sup>

2: accurate determination  
of big “G”

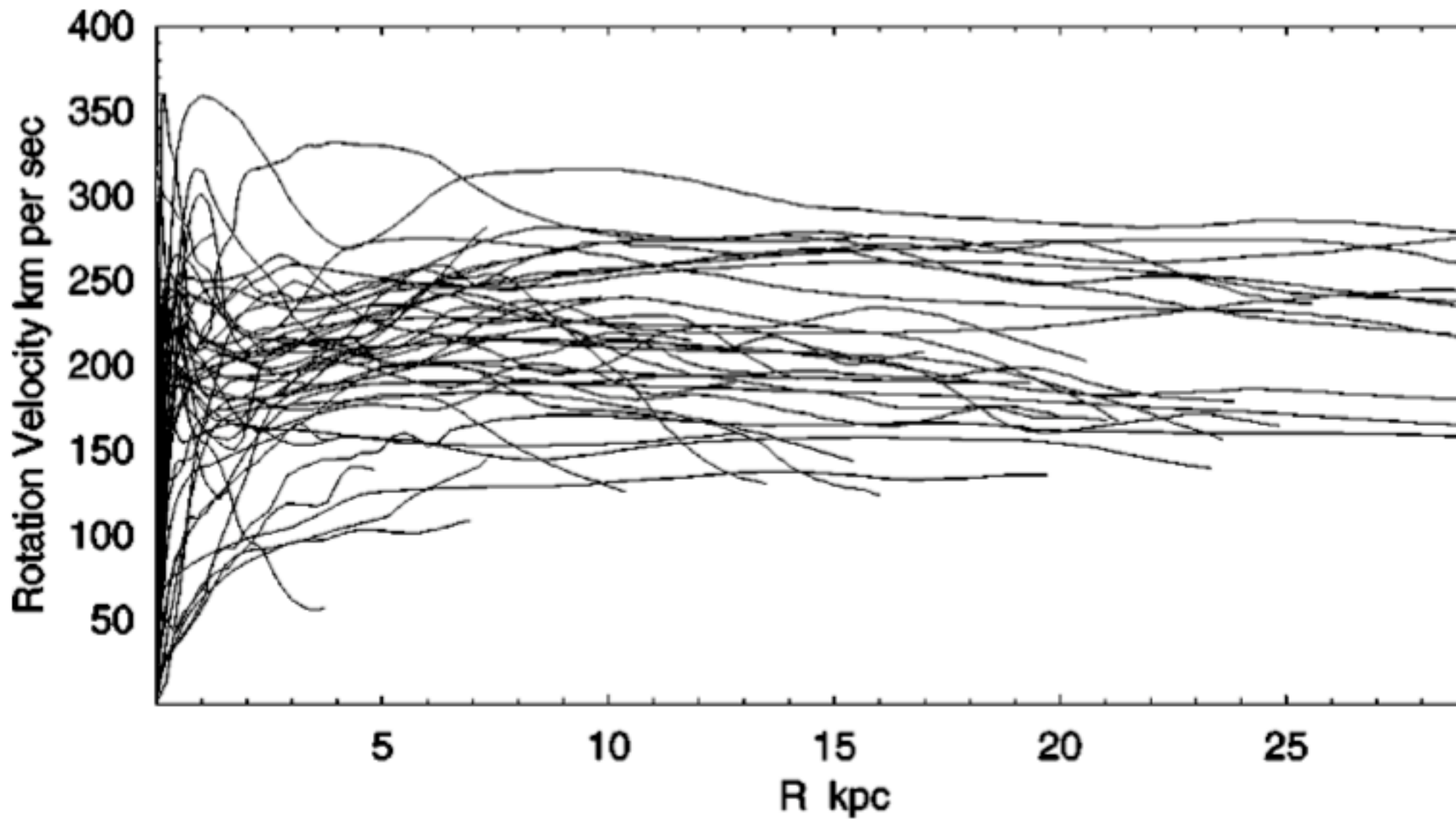


Vera Rubin 1970, with K. Ford

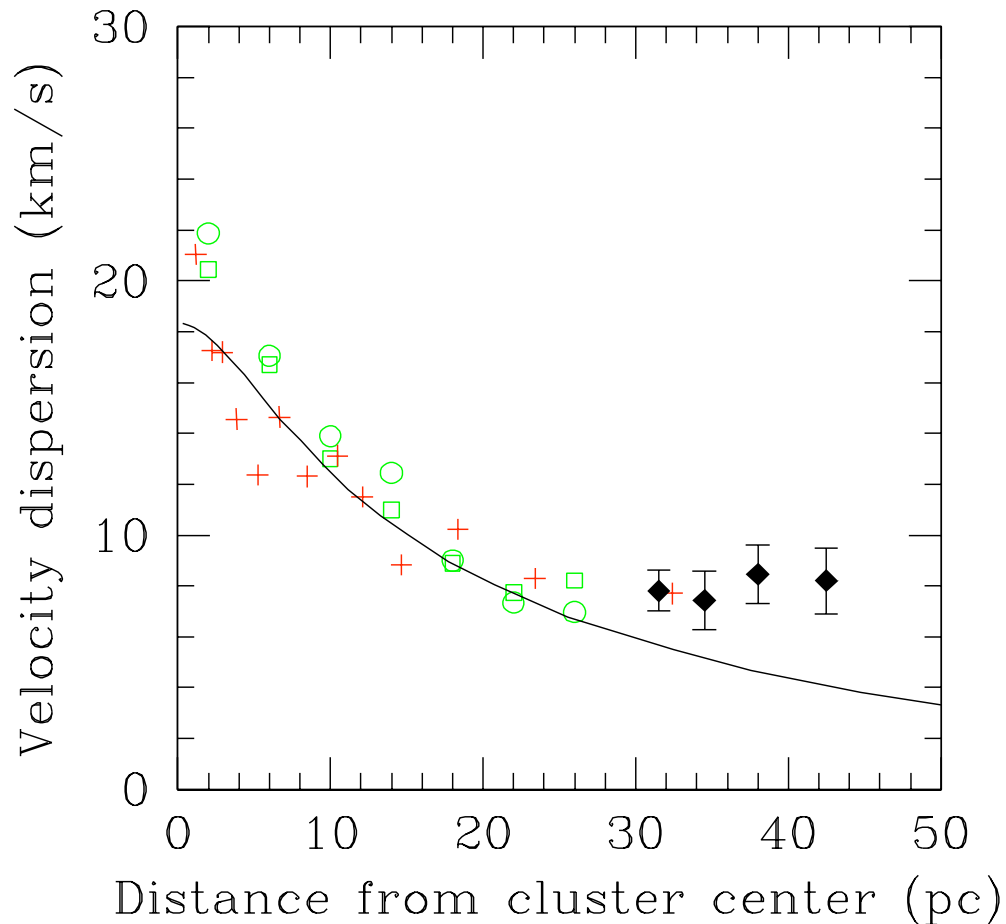
Rotation curves a general phenomenon



Andromeda M31



**Figure 4** Rotation curves of spiral galaxies obtained by combining CO data for the central regions, optical for disks, and for outer disk and halo (Sofue et al. 1999a)



Velocity dispersion profile of  $\omega$  Cen. Proper motion data (**Circles** and **Squares**; van Leeuwen et al. 2000) and radial velocities (**Crosses**; Meylan et al. 1995 up to 20 pc, last two points from Meylan & Mayor 1986) agree well showing the cluster is isotropic. The solid line is the best fit model to the radial velocity data as in Meylan et al. (1995) (their fig 1). Our velocity dispersion data (**Diamonds**) show the dispersion starts to be constant for  $R > 25$  pc, where the acceleration falls below  $10^{-7} \text{ cm s}^{-2}$ .

**Scarpa, Gilmozzi, et al. 2006,  
globular clusters**

# Newton's law

$$F(\text{grav}) = GMm/r^2$$

rules all (bound) systems in the universe

1: change  $M$ , by introducing CDM

2: change accel. by introducing  $a(0)$  at very small  
acceleration  
MOND

3: adding an Yukawa like term at very small distance

4: consider  $G$  to be time dependent

# MOND

- **MO**dified **N**ewtonian **D**ynamics 1983 M. Milgrom
- Change Newtons Law at accelerations below about  $10^{-10}$  m/sec<sup>2</sup>
- Provides an excellent description of the non-Newtonian behavior with just one additional universal parameter  $a(0)$



“MOND”

Milgrom 1983

is non-relativistic

but

“TeVeS”

Bekenstein 2004

is one

relativistic extension



virial law

$$E(\text{kin}) = -1/2 E(\text{pot})$$

Fritz Zwicky, geboren 1898 in Wäna und als Schweizerbürger im Kanton Glarus aufgewachsen, studierte an der Eidgenössischen Technischen Hochschule in Zürich und wurde dort 1922 zum Dr. sc. nat. promoviert. Seit 1925 lebt er in den USA; er ist heute Professor für Astrophysik am California Institute of Technology in Pasadena und Astronom der Mount Wilson- und der Mount-Palomar-Sternwarte. Auch auf humanitärem Gebiet ist Professor Zwicky sehr aktiv. Besonders wichtig aber sind seine Bemühungen als Morphologe, wie er sie in diesem Buch darstellt: Es geht ihm darum, als »Spezialist des Unmöglichen« mit einem Minimum an Arbeit und Zeit zu einem Optimum von Lösungen gegebener Probleme zu gelangen und dabei zugleich neue Probleme zu entdecken.

Helvetica Physica Acta  
1933

Fritz  
Zwicky

Entdecken  
Erfinden  
Forschen

Droemer  
Knaur

Fritz Zwicky

Entdecken, Erfinden, Forschen  
im morphologischen Weltbild

Droemer Knaur



# Fritz Zwicky

1933 and 1937 he proposed

- SN explosions to produce N-Stars
- Introduced the term Dark Matter
- observed Newton to fail for the Coma-CL
- Gravitational lensing of distant galaxies by foreground galaxies

# Clover Leaf



And in the solar system??

at least two dramatic effects

1: “Pioneer” anomaly

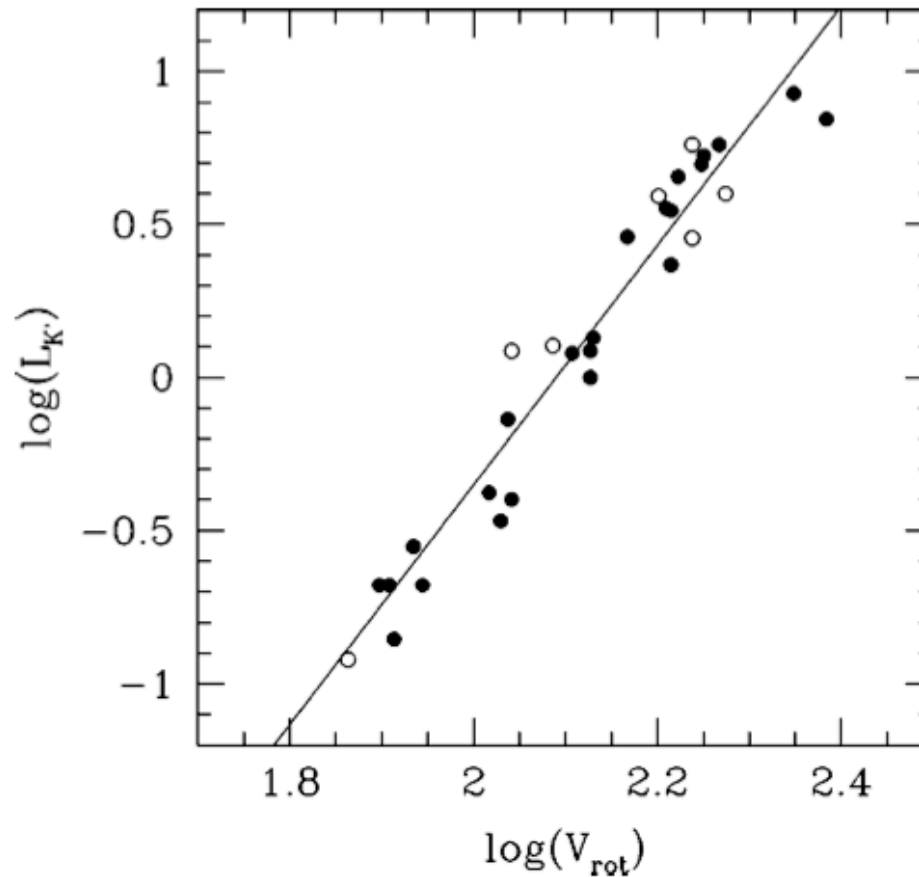
2: “fly by” effect

again seem both to be ruled by

$a(0)$

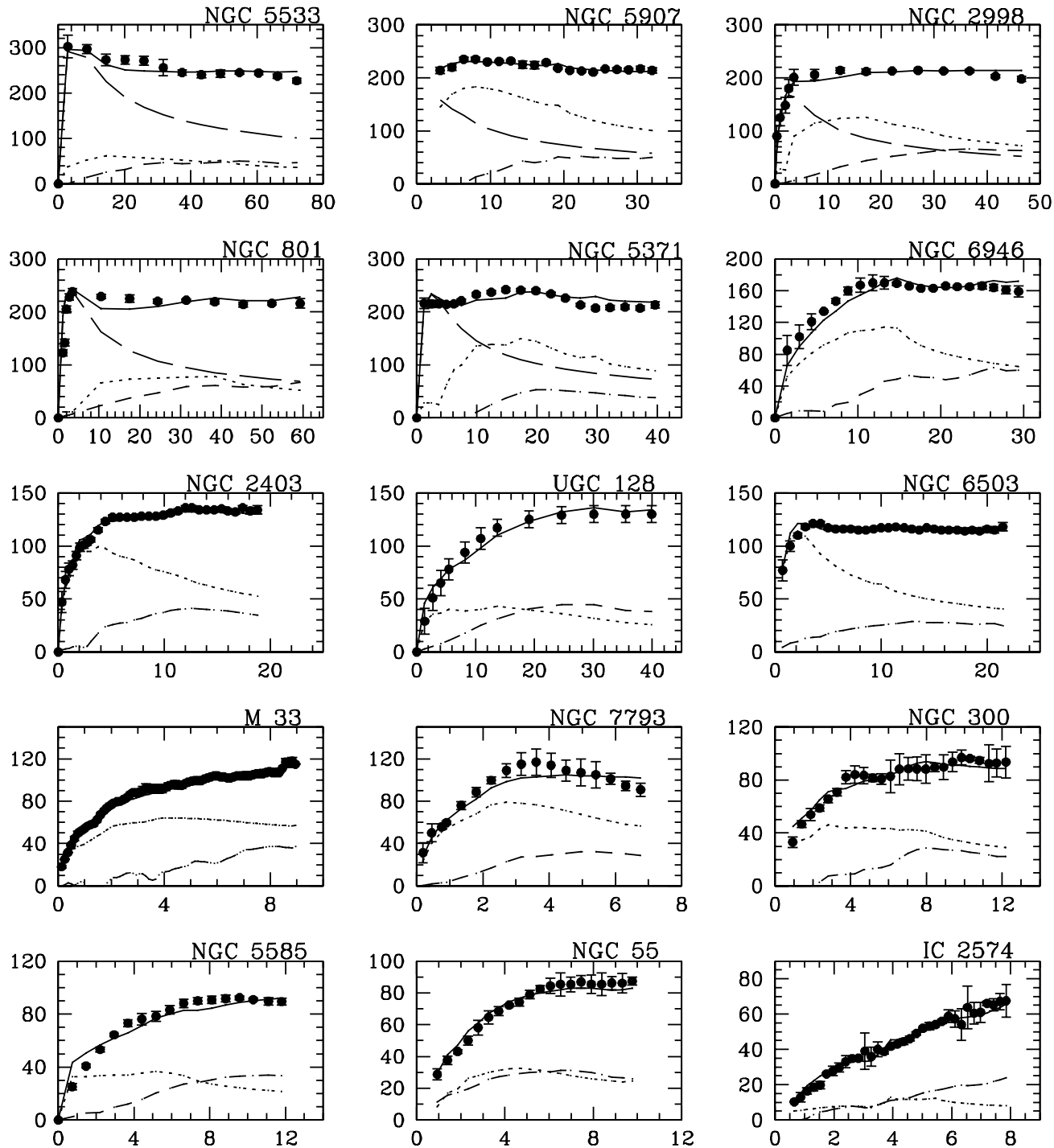
which is about  $cH$  !!

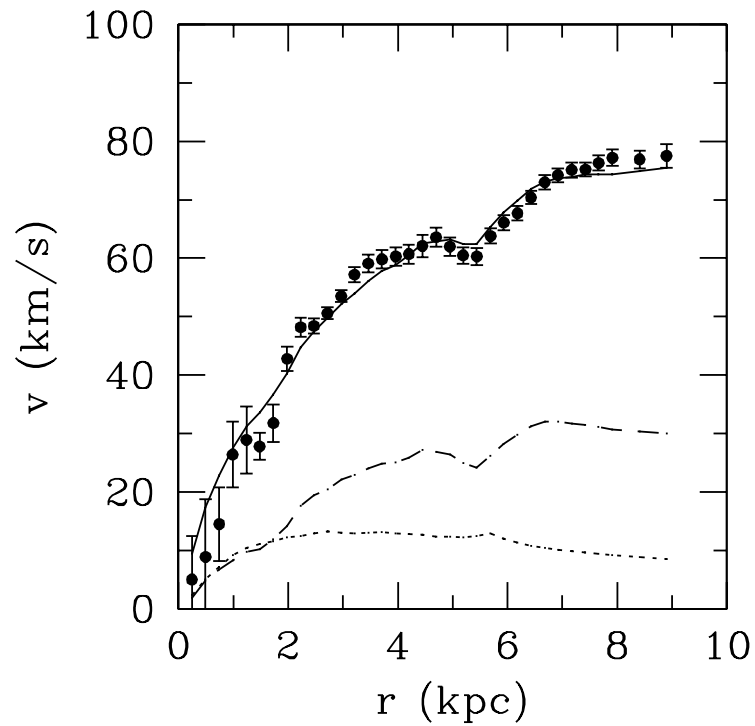
# Sanders and McGaugh



**Figure 2** The near-infrared Tully-Fisher relation of Ursa Major spirals (Sanders & Verheijen 1998). The rotation velocity is the asymptotically constant value. The velocity is in units of kilometer/second and luminosity in  $10^{10} L_{\odot}$ . The unshaded points are galaxies with disturbed kinematics. The line is a least-square fit to the data and has a slope of  $3.9 \pm 0.2$ .

# Rotation in MOND

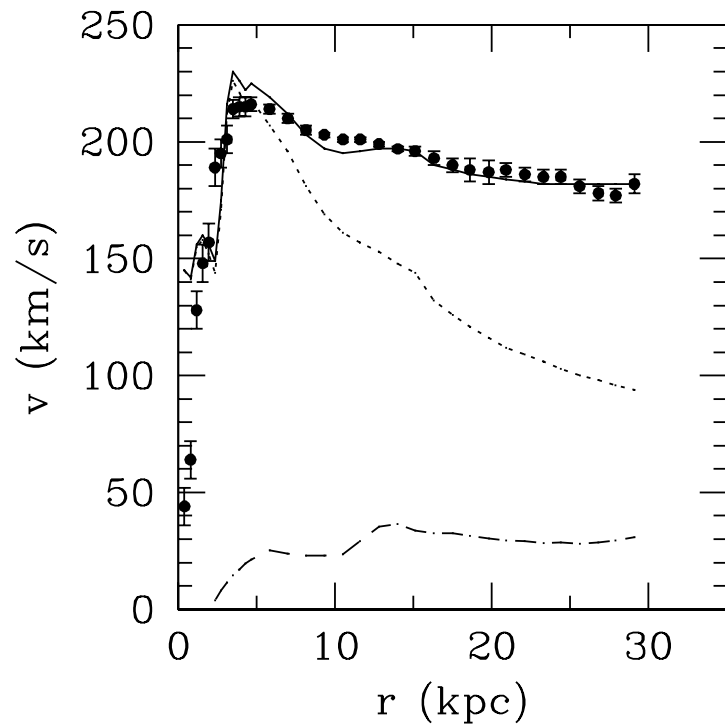




NGC 1560

$$\langle \mu_B \rangle = 23.2 \text{ mag/arcsec}^2$$

$$(M/L_B)_{\text{disk}} = 0.4$$



NGC 2903

$$\langle \mu_B \rangle = 20.5 \text{ mag/arcsec}^2$$

$$(M/L_B)_{\text{disk}} = 1.9$$



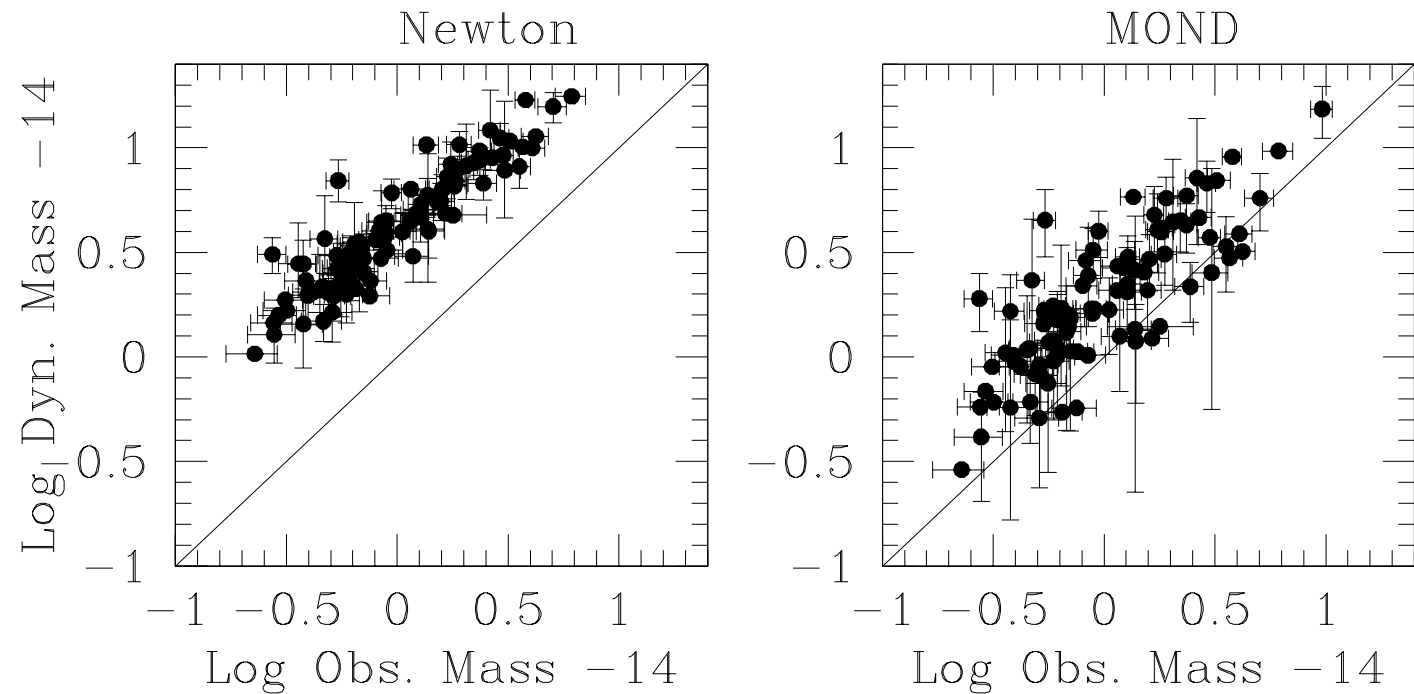


Figure 10: (*Left*) the Newtonian dynamical mass of clusters of galaxies within an observed cutoff radius ( $r_{out}$ ) vs. the total observable mass in 93 X-ray emitting clusters of galaxies (White et al. 1997). The solid line corresponds to  $M_{dyn} = M_{obs}$  (no discrepancy). (*Right*) the MOND dynamical mass within  $r_{out}$  vs. the total observable mass for the same X-ray emitting clusters. From Sanders (1999).

Sanders, 1999

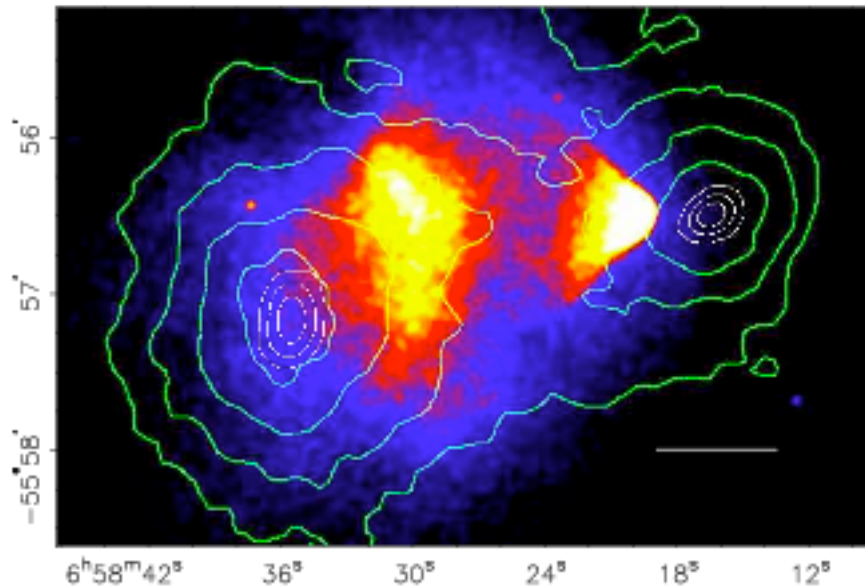


FIG. 6: The colliding clusters 1E0657-56. The bullet cluster (right) rammed through the cluster on the left. Hot gas stripped off both clusters is colored red-yellow. Green and white curves are level surfaces of gravitational lensing convergence; the two peaks of this do not coincide with those of the gas which makes up most of the visible mass, but are skewed in the direction of the galaxy concentrations. The white bar corresponds to 200 kpc. Figure reproduced from Ref. 85 by permission of the American Astronomical Society.

MOND works:  
however  
needs  
Neutrinos with mass  
see  
Zhao 2007

# Vera Rubin

## (final remarks)

10. Dark halos may finally be understood. We will know their extent and their relation to the intracluster dark mass. We may even know the rotation velocity of the halo. Will the concept of a “rotation curve” apply at such large distances from the disk? Will we learn if our halo brushes the halo of M31?
11. We will ultimately know what is the dark matter—the major mass constituent of the Universe. Elementary particle physics will teach us its origin and physical properties.
12. Perhaps we will be able to put to rest the last doubt about the applicability of Newtonian gravitational theory on a cosmic scale, or enthusiastically embrace its successor.

Thus, it is my personal opinion – and I am the only one responsible for it if proved wrong-- that if Newtonian dynamics fails below  $a_0$ , this should be true irrespectively of the total field and one should be able to observe MOND effect also here on earth. For instance, I think a refined version of the Cavendish experiment studying gravitational forces in the horizontal plane should detect MOND effects.

**R. Scarpa 2006**

# Setup in Wuppertal

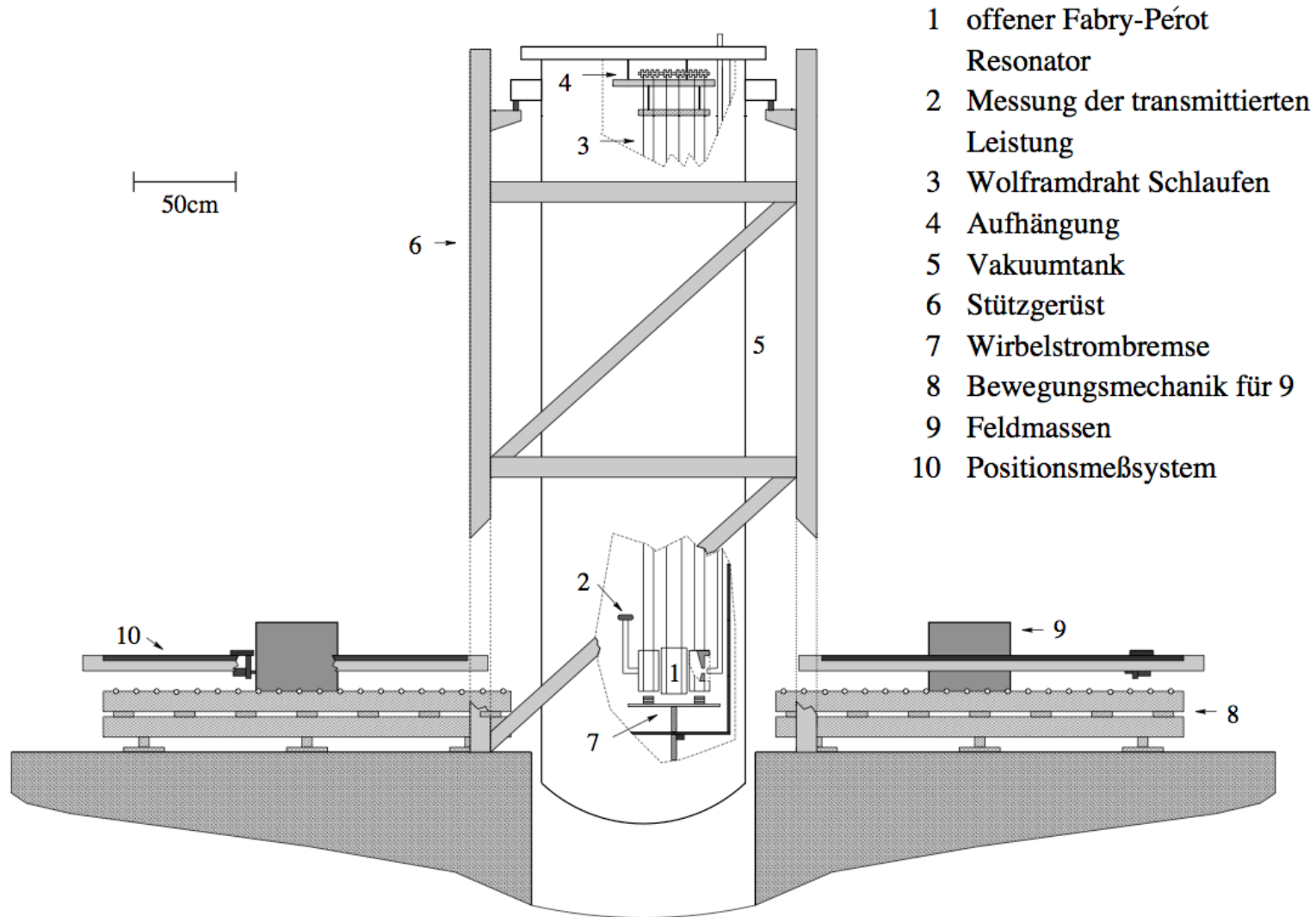
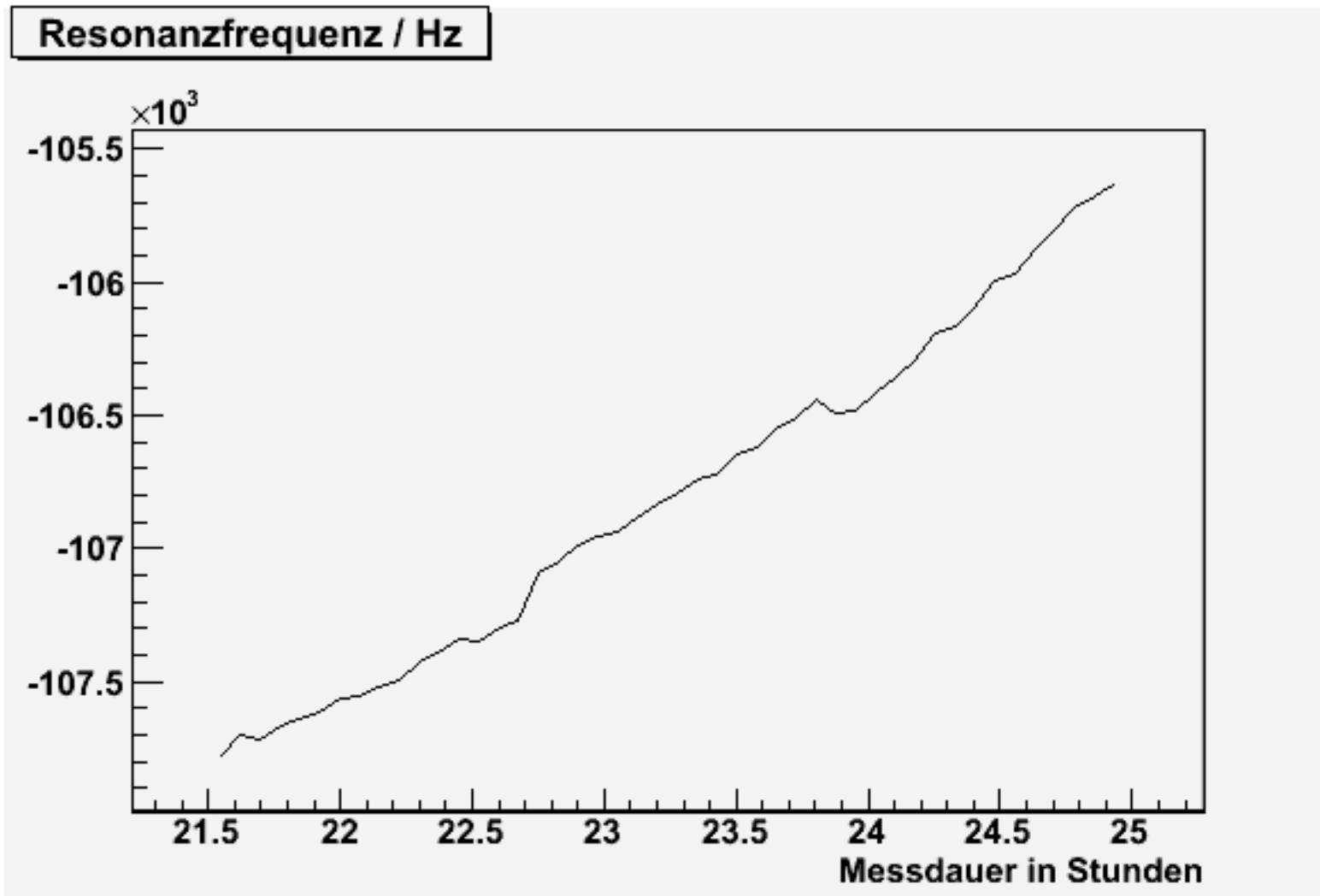


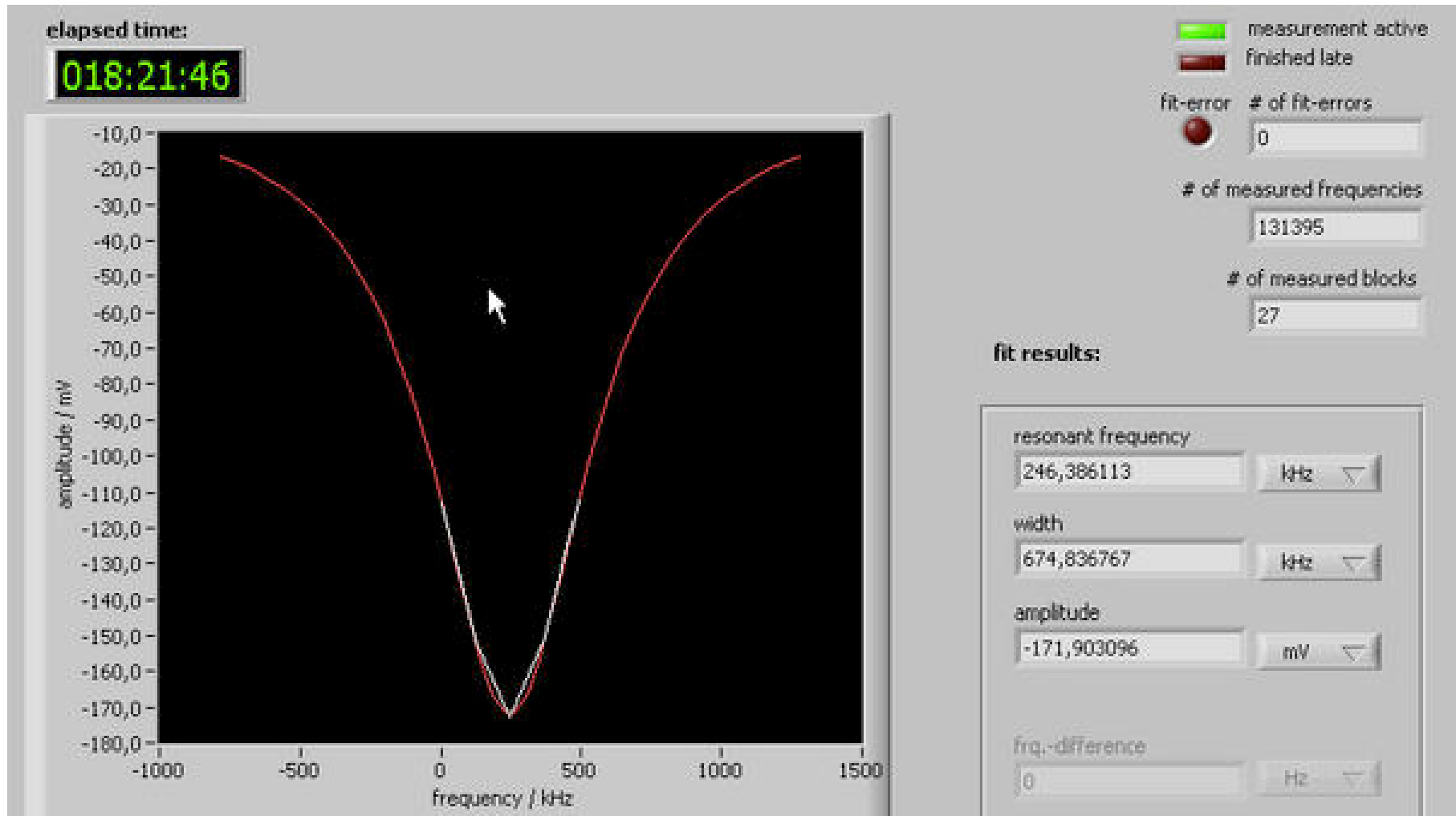
Abbildung 4.1: Schematische Darstellung des Wuppertaler Gravitations-Experimentes.

# “Gravi” at DESY works!!! (18.6.2008)

## 127Hz in 22GHz (1/10 nm)



# 21 GHz resonance in our cavity

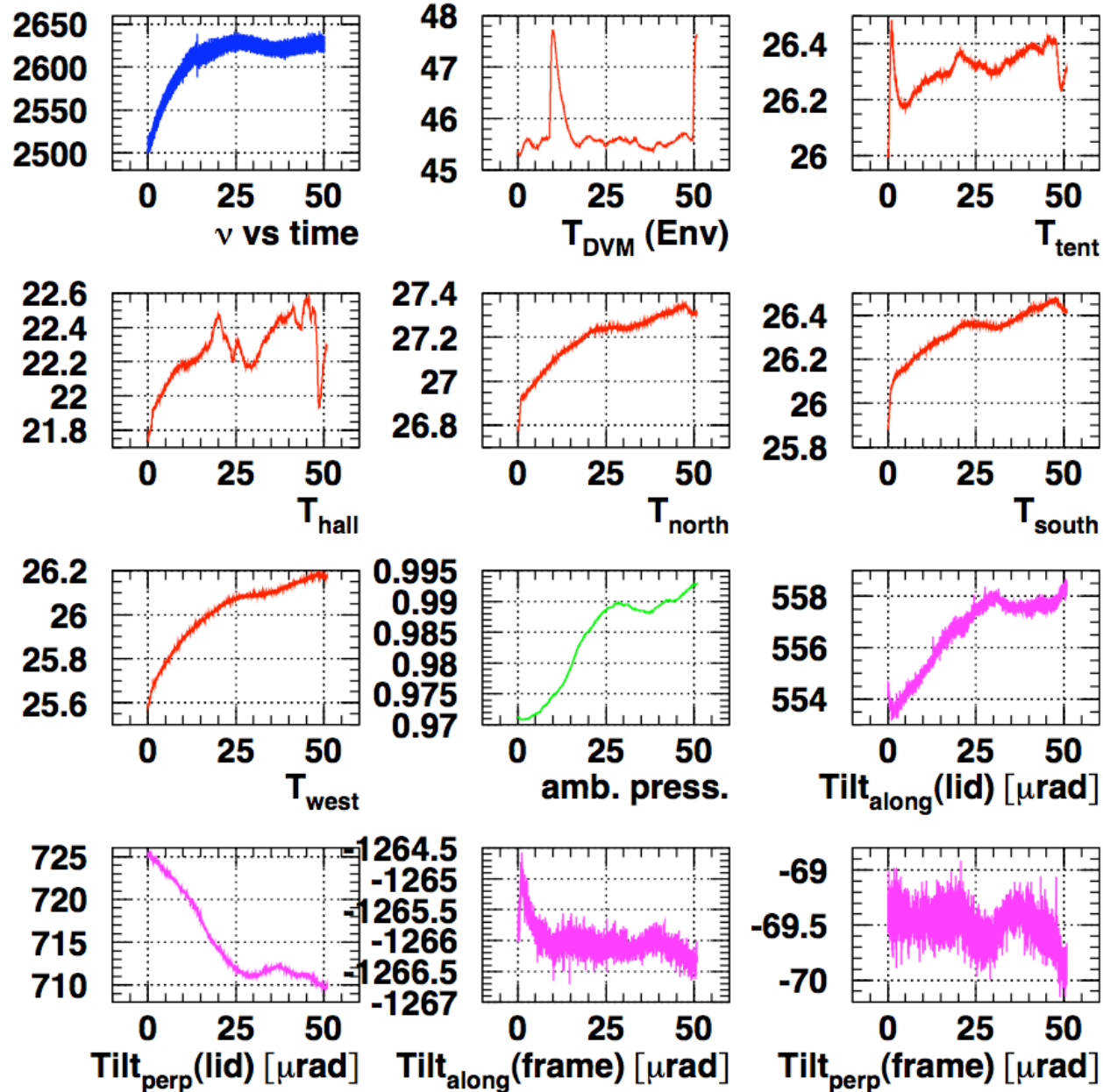


About 1/sec, 5 point fit, our basic signal

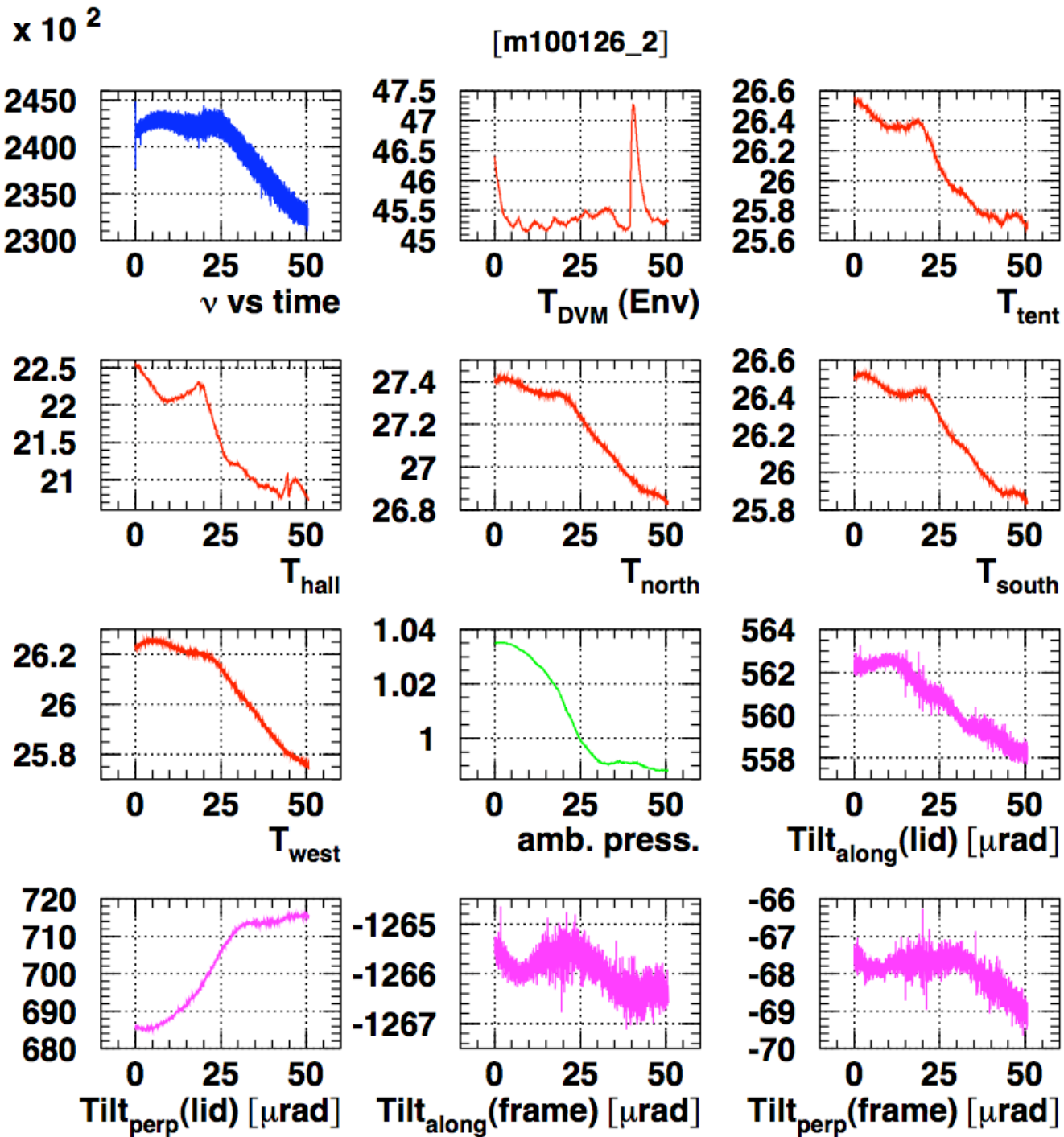
# Data vs run time

$\times 10^2$

[m100129]

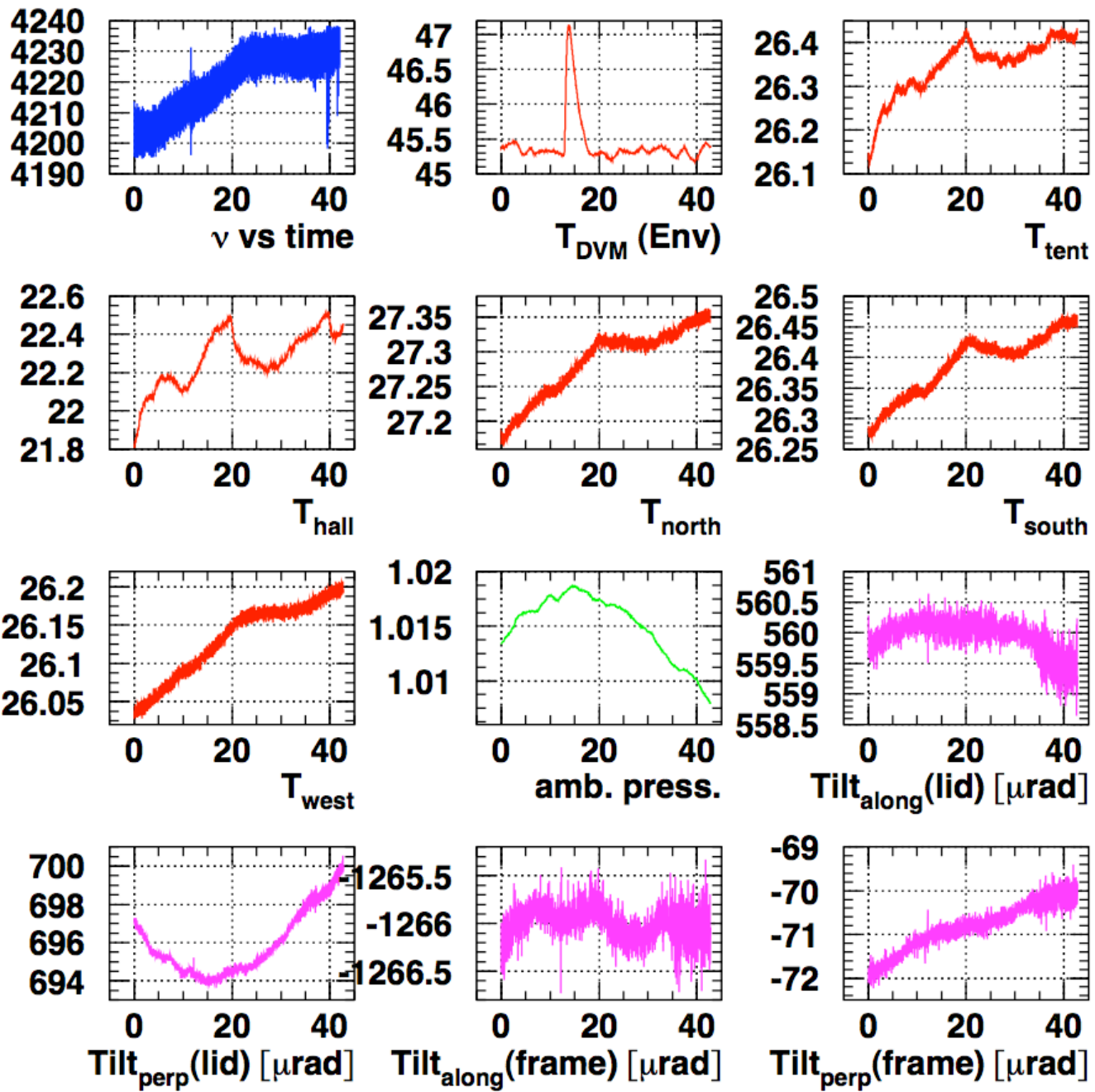




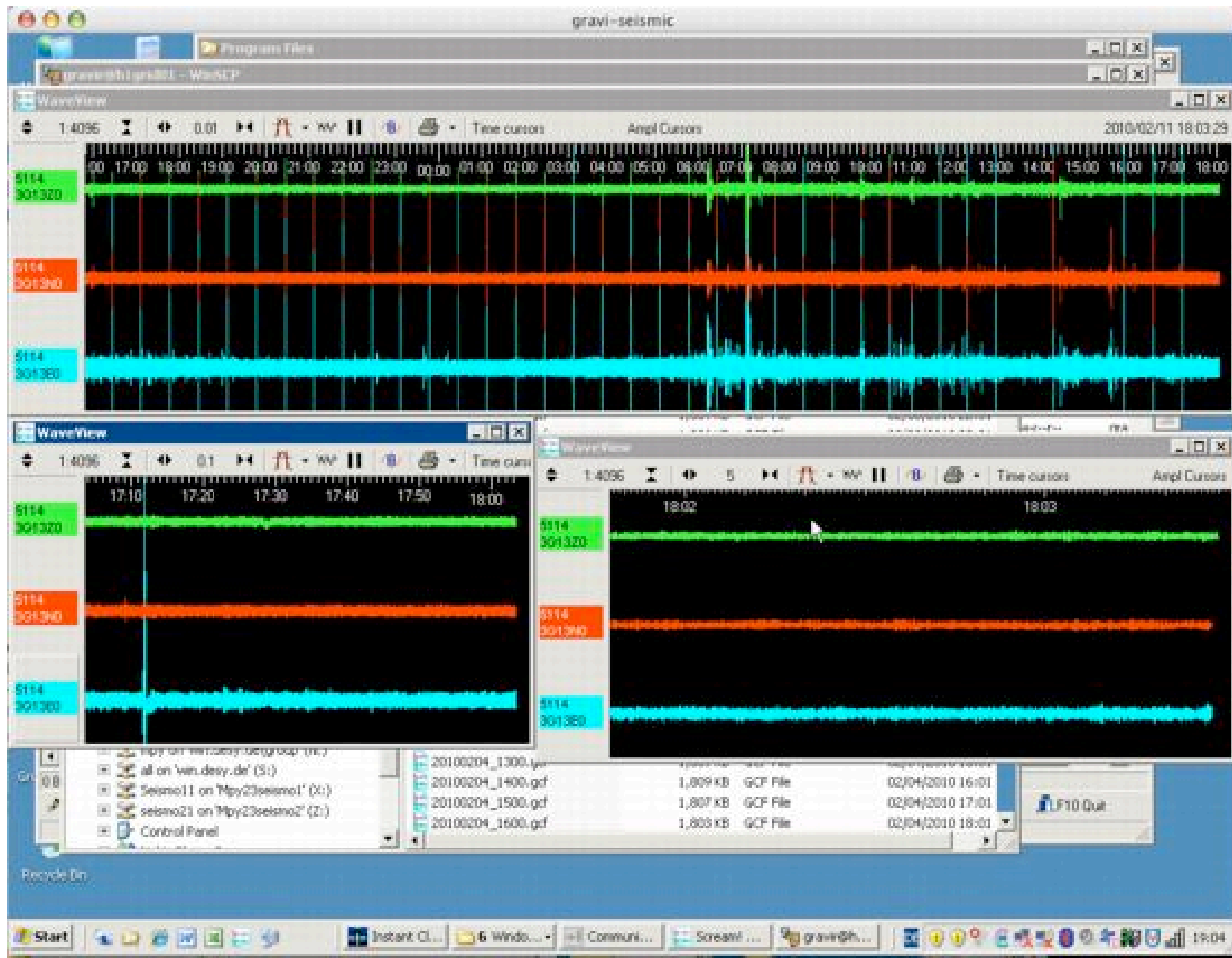


$\times 10^2$

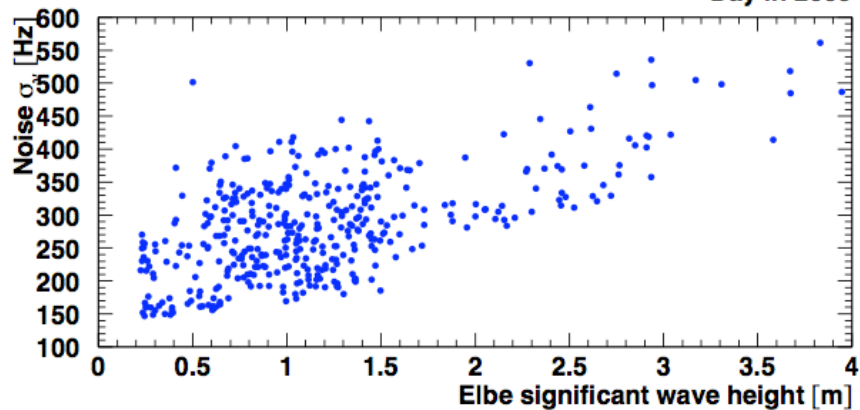
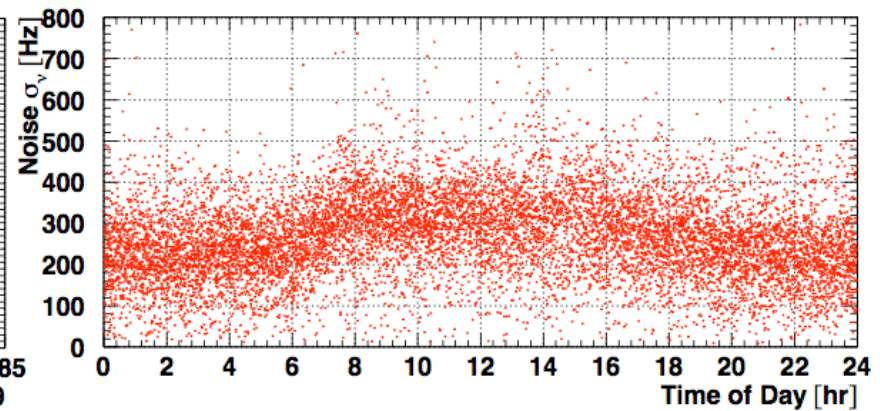
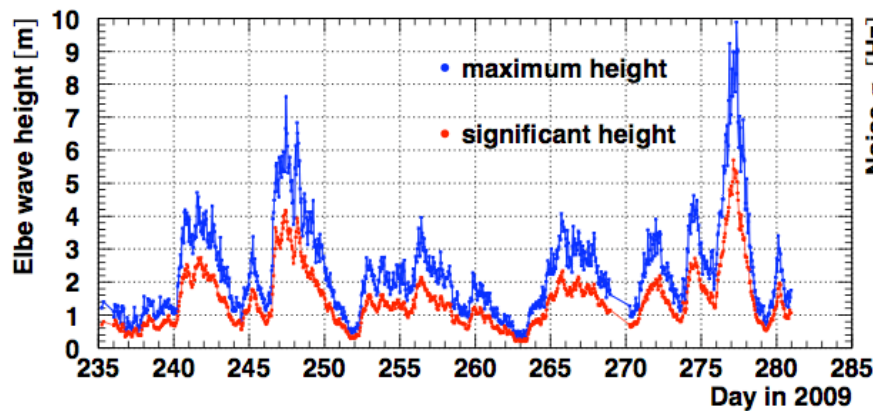
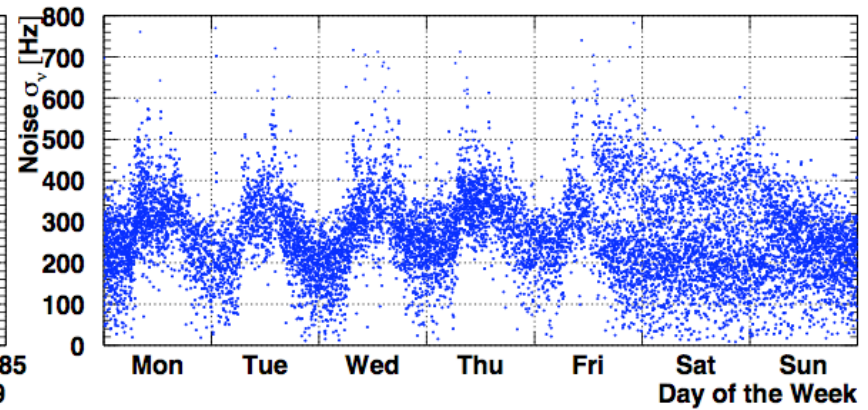
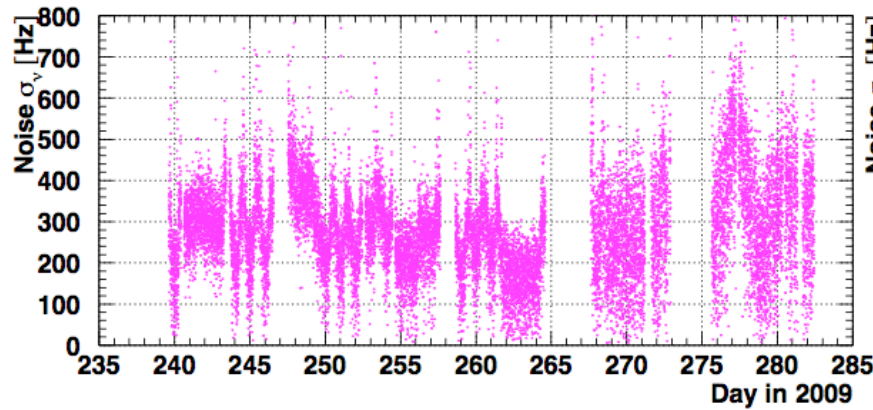
[m100206\_2]



# Seismic signal

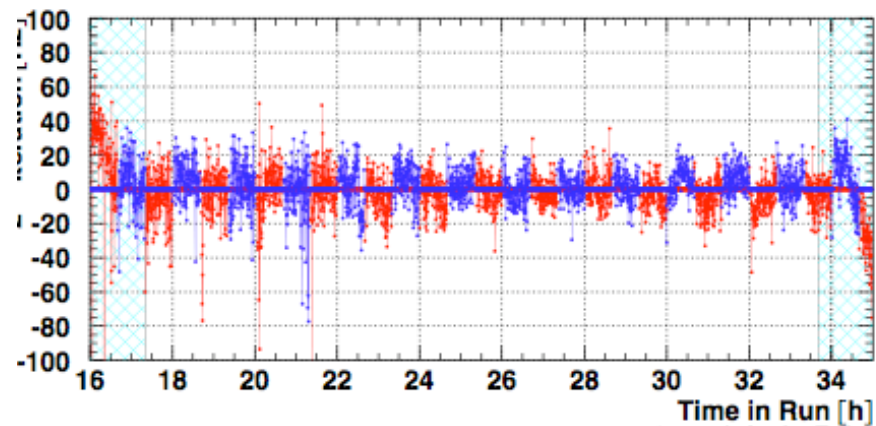
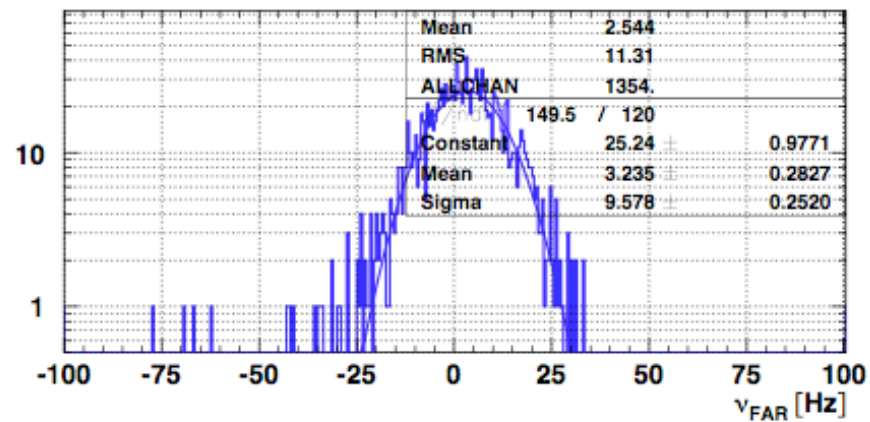
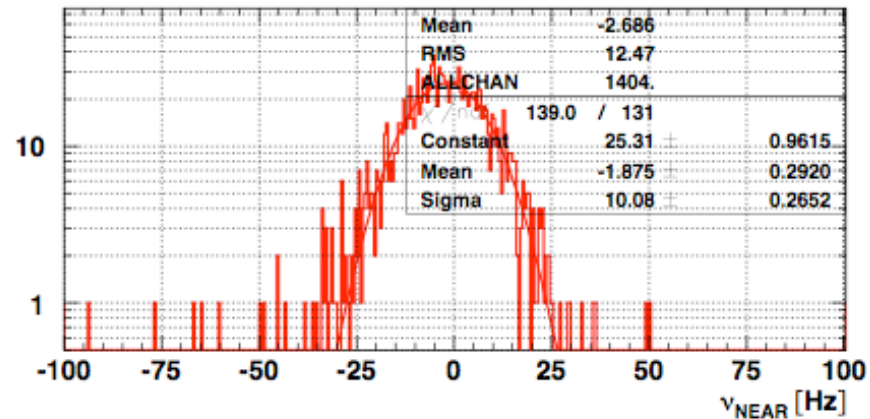


# NOISE

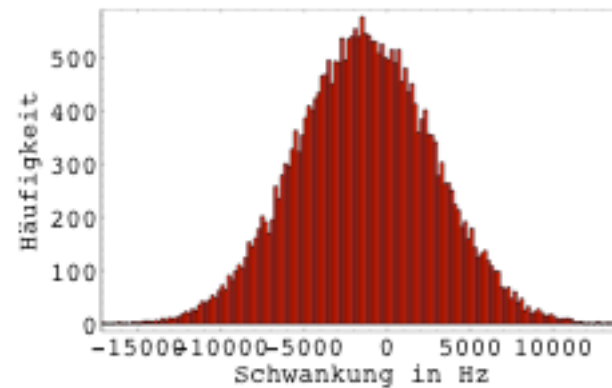
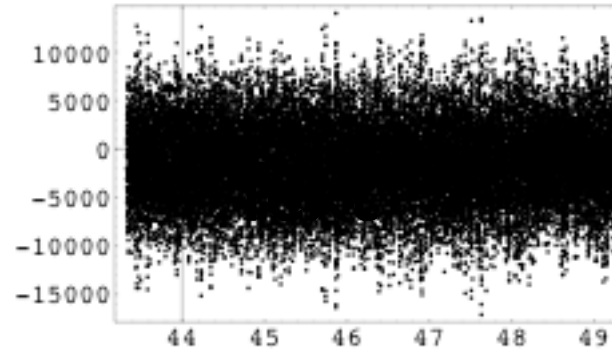
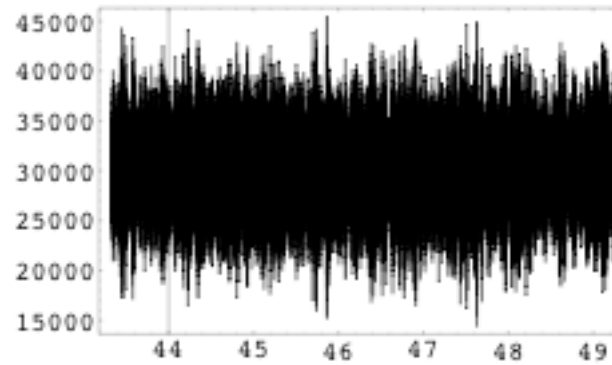




# Gaussian distribution ??

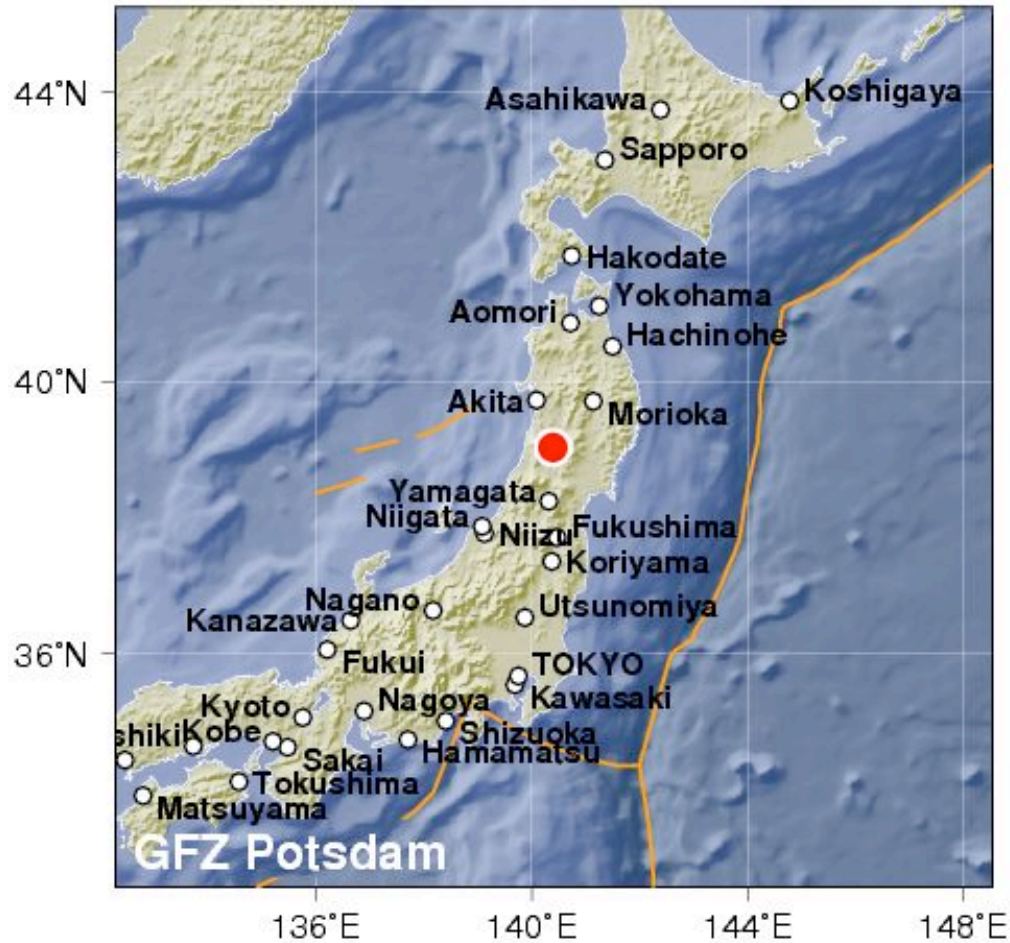


Noise  
nicely  
Gaussian,  
very  
crucial



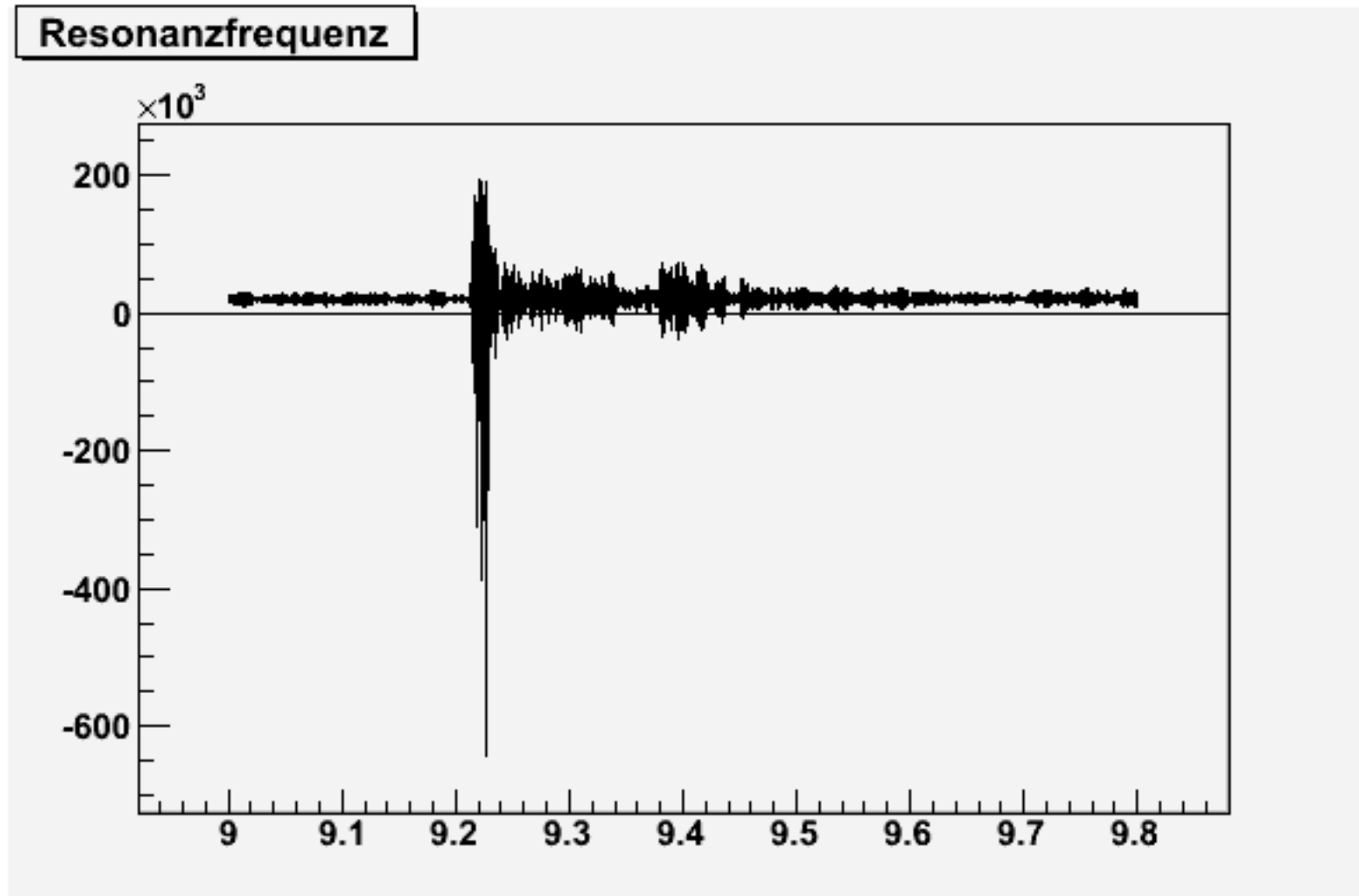
frequency  
vs.  
time

# Earthquake in Japan

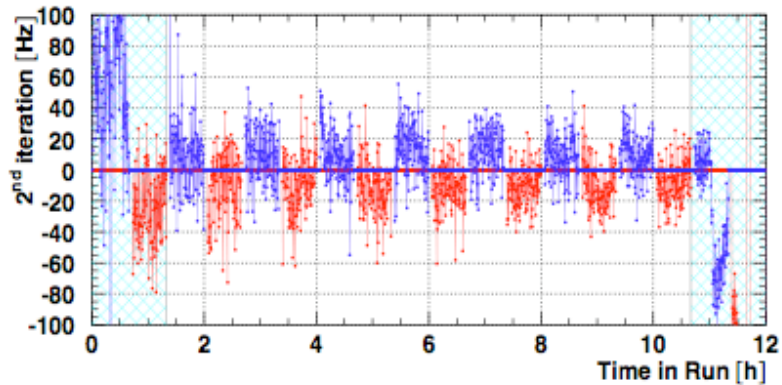




# Earthquake signal in GRAVI

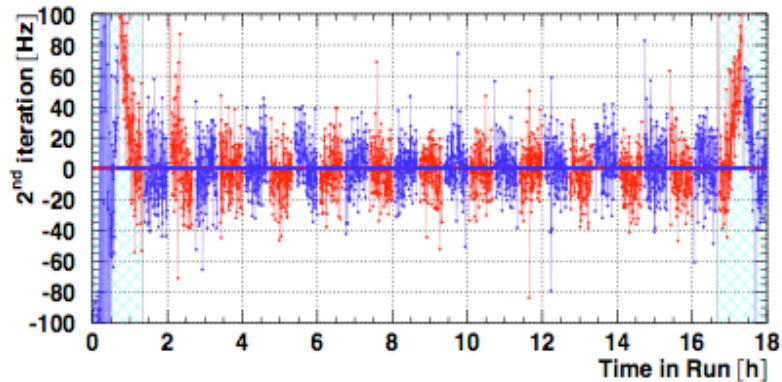


# F-Signal (in/out)

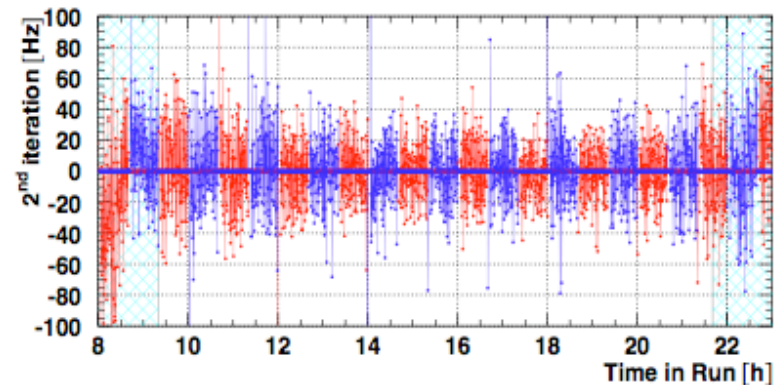


9 kg  
m090930

Newtons Law  
strictly proportional  
to the fieldmass



2.9 kg  
m090918

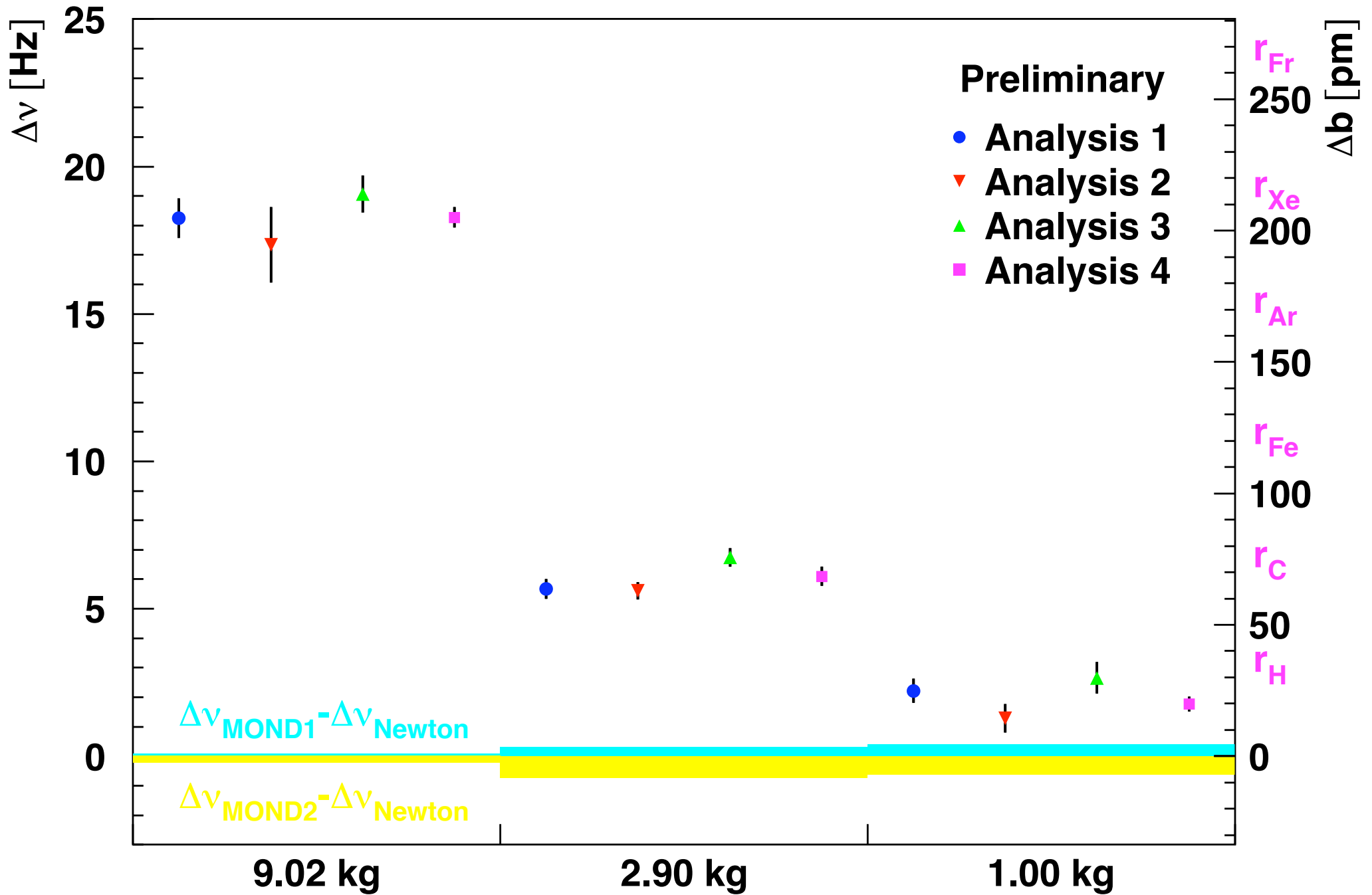


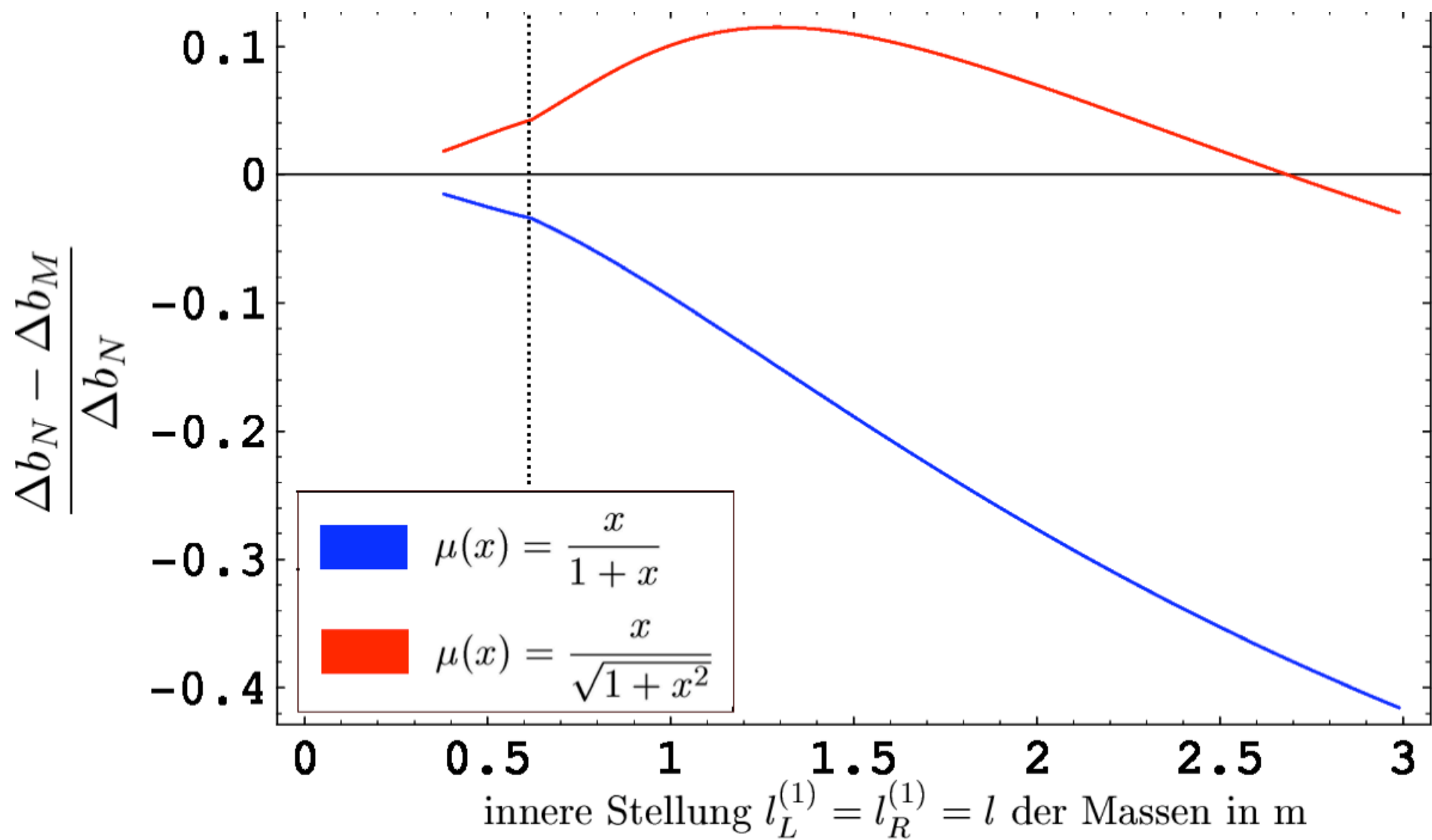
1.0 kg  
m091216

# Analysis

Have 4 independent lines of analysis  
Consider this to be absolutely crucial  
for such a fundamental experiment

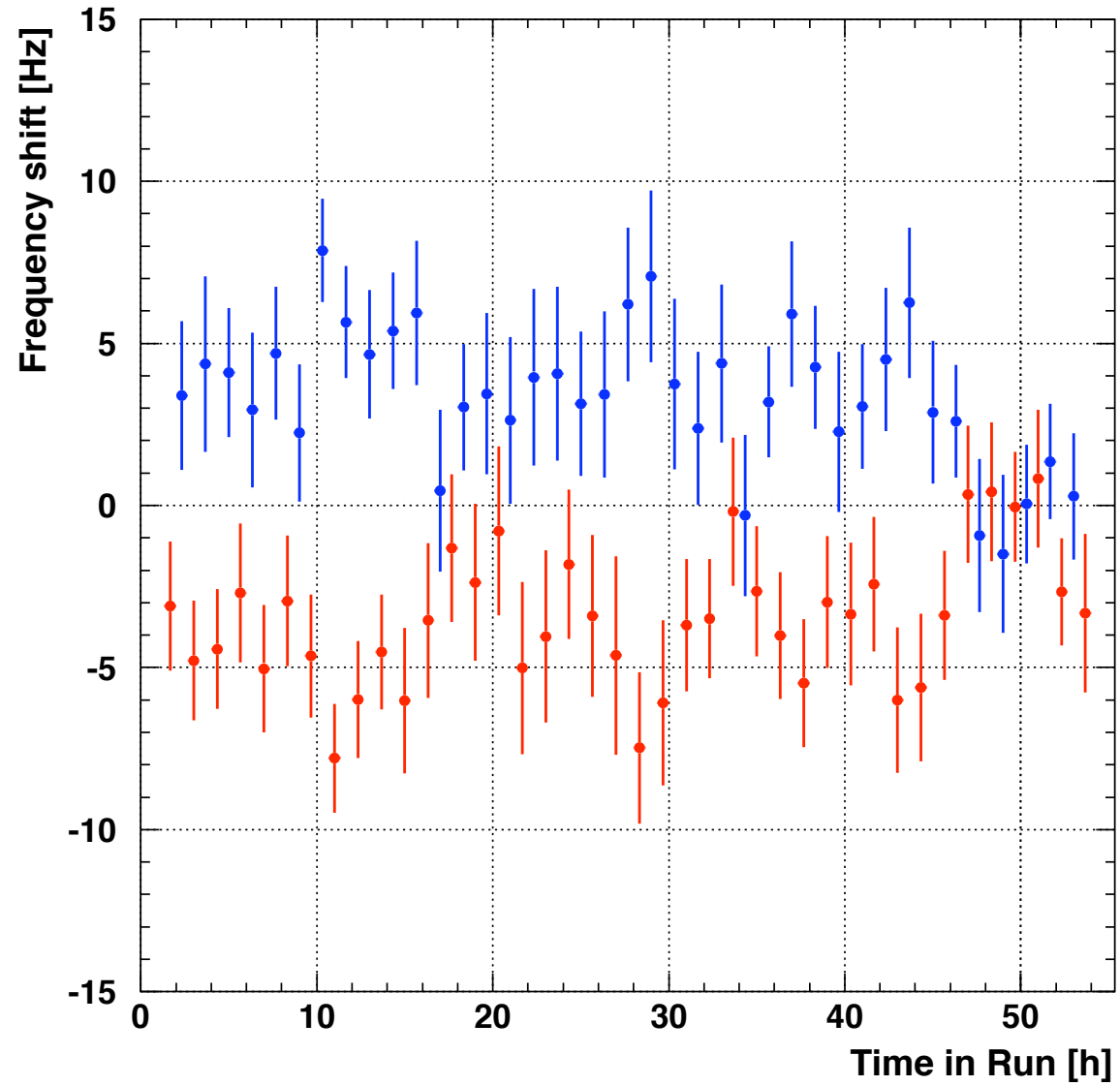
# “Preliminary”





# Problem?

$$\Delta v_{\text{fit/mean}} = 7.04 \pm 0.16 \text{ Hz}$$



# Summary

- Take data since about 9/2009
- Use 3 Masses (9, 2.9 and 1kg) at 77cm
- Get results roughly in agreement with Newton
- Have still some unexplained systematic problems
- Have to do some further adjustments to the setup, but we are optimistic to come out soon with a word on Newton vs. MOND

fascinating times  
testing  
**GRAVITY**  
in the lab

I like to thank DESY  
for unbureaucratic  
support

Not to forget!!!  
H.J. Seidel  
U. Cornett, T. Külper  
C. Muhl, D. Habercorn  
J. Schaffran