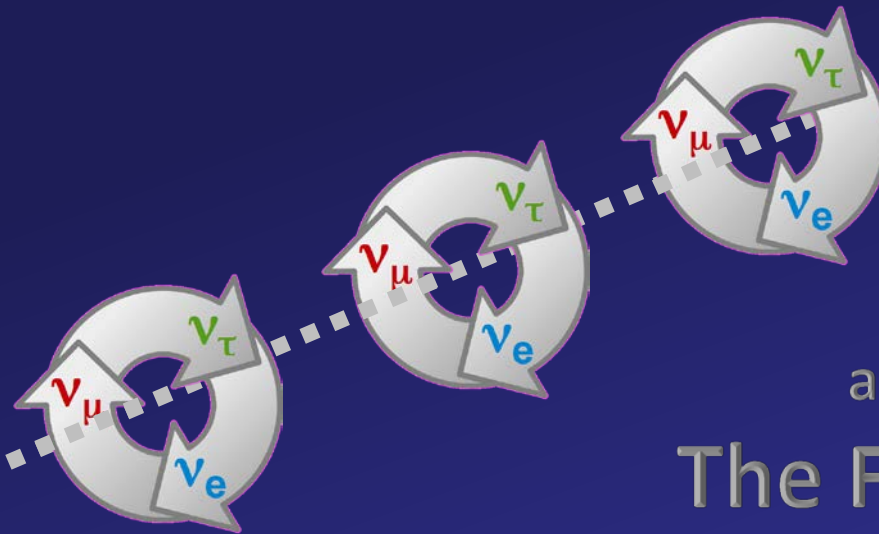


Colloquium – DESY Zeuthen

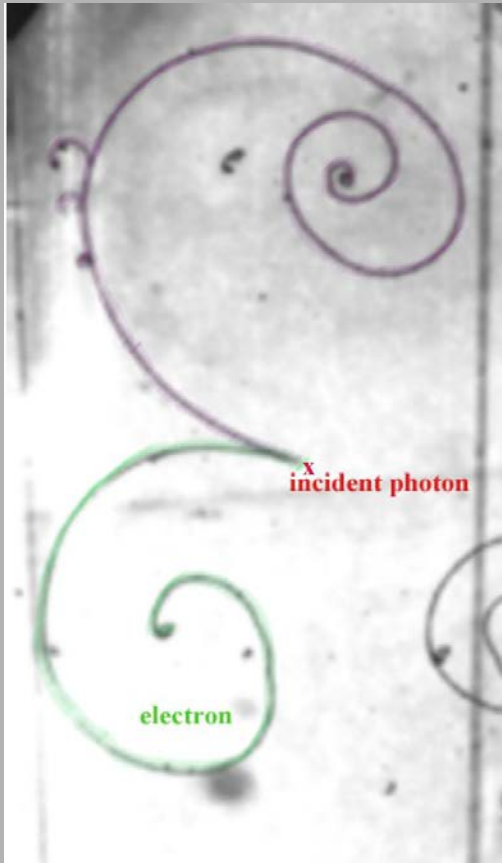
JUNO



and
The Fate of
Antimatter

In the beginning ...

tested in the lab



„a million times“

matter

anti-matter

created matter and antimatter
in equal amounts

... the Big Bang

Today ...

... we only find matter



Evolution of Matter

Galaxy A1689-zD1:
~700 million years
after the Big Bang

Big Bang

Radiation era

~300,000 years:
"Dark Ages" begin

~400 million years: Stars
and nascent galaxies form

on years: Dark ages end

~4.5 billion years: Sun, Earth, and solar system have formed

matter and antimatter annihilated ...

How ?

... some matter survived

Andrei Sakharov



1. Baryon-Number Violation **theoretical ideas**
2. CP-Violation **not enough !**
3. Thermal Non-Equilibrium **understood**

• 13.71

Today ...

Baryon to Photon Ratio:

$$\eta = \frac{n_B - n_{\bar{B}}}{n_\gamma} \approx 5 \cdot 10^{-10}$$

$$n_\gamma \approx 0.4/\text{mm}^3$$

$$n_B \approx 0.2/\text{m}^3$$

$$n_{\bar{B}} \approx 0$$

**Standard Model fails by
many orders of magnitude**

... we only find matter

Concept of JARA-FAME

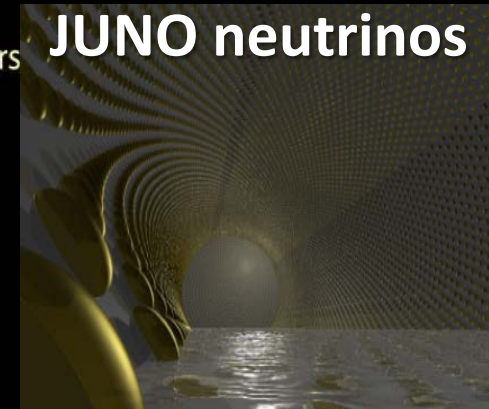
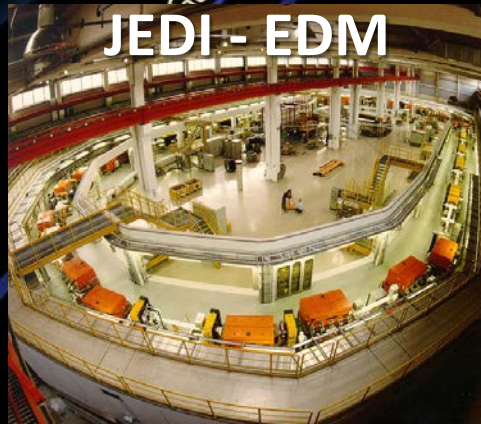
antimatter
separated

antimatter
disappeared

More CP-violation needed !

baryo-genesis

lepto-genesis



Big Bang

Radiation era

~300,000 years:
"Dark Ages" begin

Stars
form

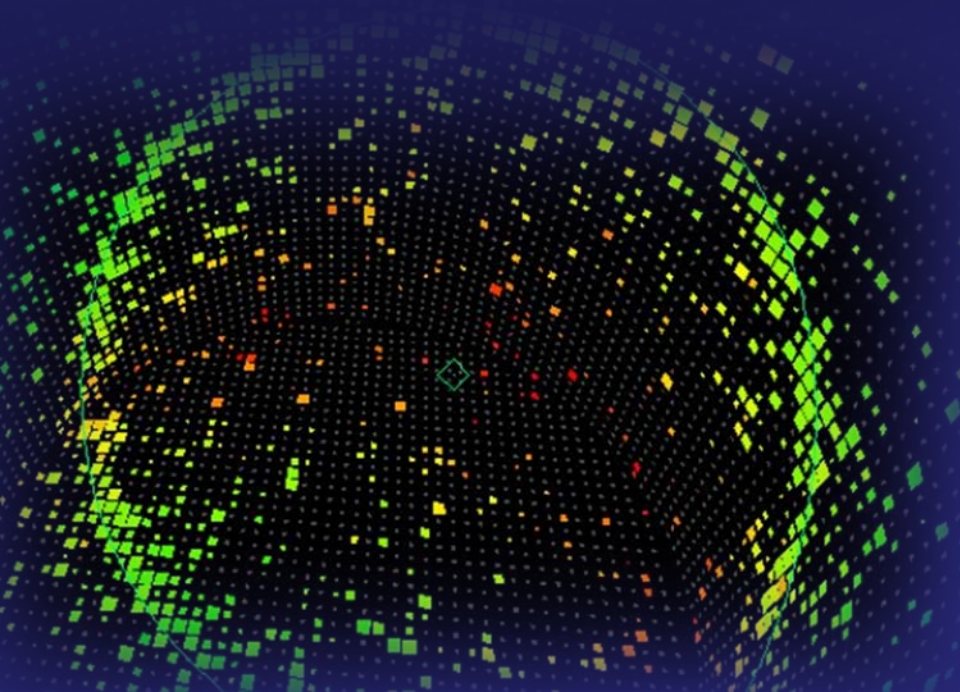
~1 billion years: Dark ages end

• 13.7 billion years: Present



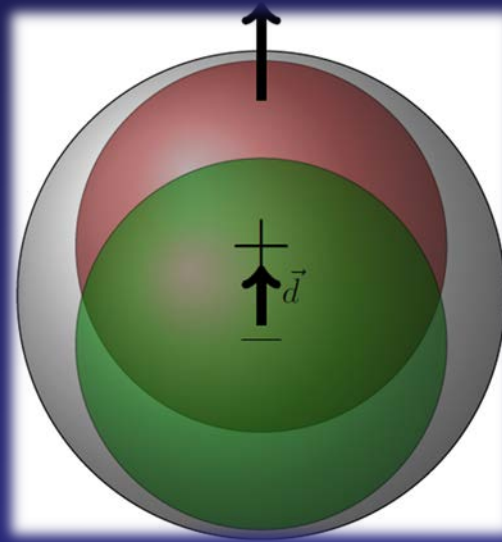
Content

- The Fate of Antimatter
- The p/d EDM
- Neutrino Oscillations
- Reactor Neutrino Experiments
- CP-Violation
- The Mass Hierarchy
- The JUNO Project
- Summary

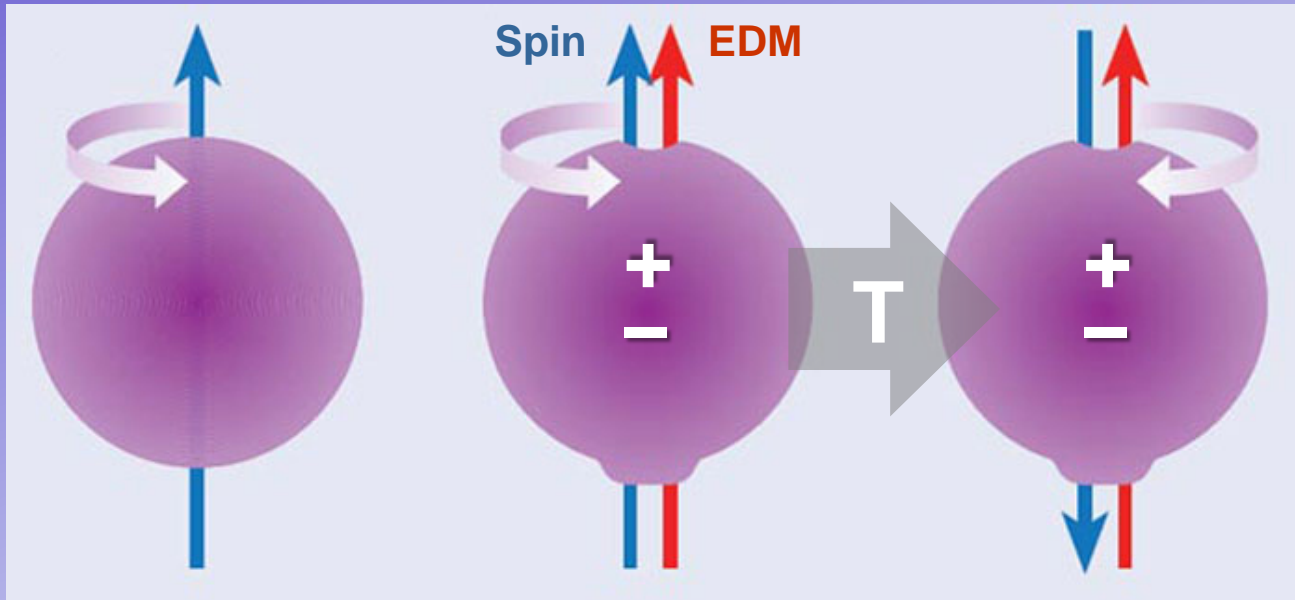




Electric Dipole Moments



ELECTRIC DIPOLE MOMENTS

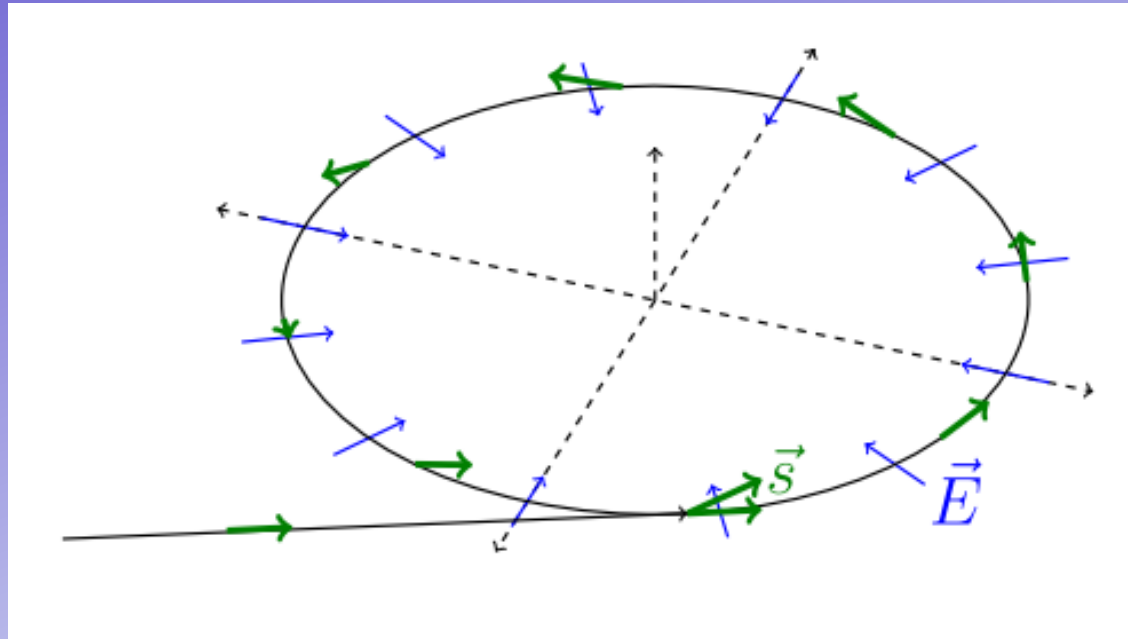


Electric Dipole Moments

→ Violate P- and T-Symmetry

→ CPT-Theorem: violate CP-Symmetry

STORAGE RING EDM



frozen spin @ $p_p = 700.740 \text{ MeV}/c$ (magic momentum)

EDM turns spin out of accelerator plane

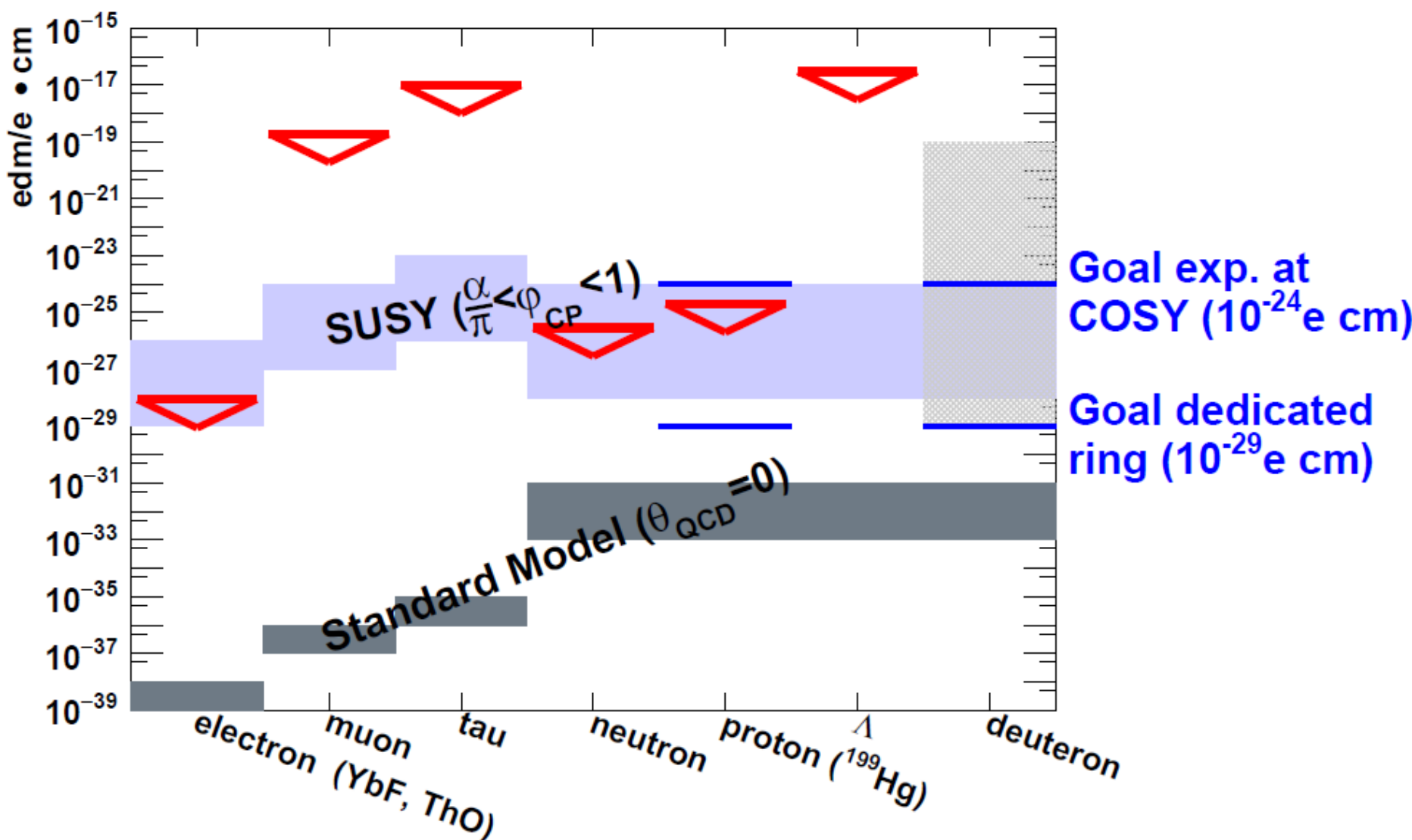
PRECURSER EXPERIMENT



Jülich
Electric
Dipole moment
Investigation

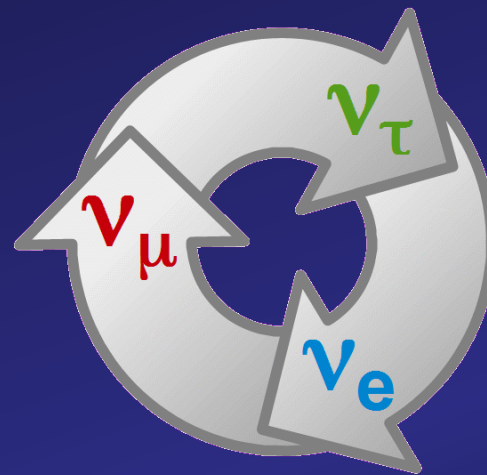


magnetic ring
RF-Wien filters
introduce f-Field





Neutrinos-Oscillations



NEUTRINO OSCILLATIONS

Flavour
Eigenstates

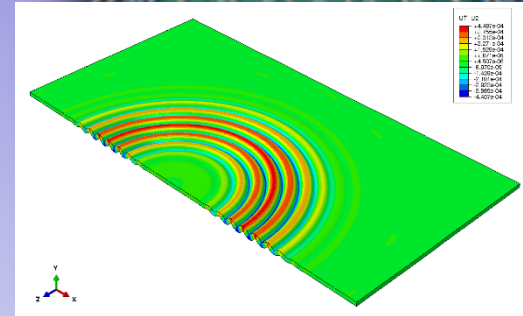
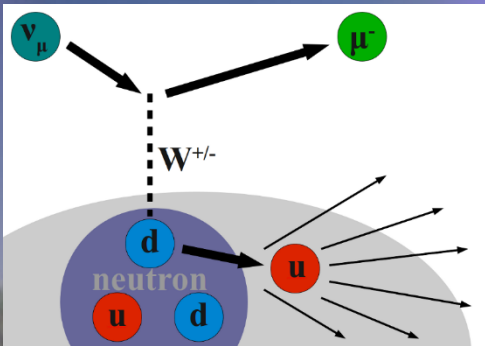
Mass
Eigenstates

$$| \nu_{\alpha} \rangle = \sum_{i=1}^3 U_{\alpha,i} | \nu_i \rangle$$

$\begin{matrix} e & \mu & \tau \\ 1 & 2 & 3 \end{matrix}$

How neutrinos
interact

How neutrinos
propagate



NEUTRINO OSCILLATIONS

Flavour
Eigenstates

Mass
Eigenstates

$$| \nu_{\alpha} \rangle = \sum_{i=1}^3 U_{\alpha,i} | \nu_i \rangle$$

$e \ \mu \ \tau$
 $1 \ 2 \ 3$

How neutrinos
interact

How neutrinos
propagate



$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix}
 \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix}
 \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Pontecorvo **M**aki **N**akagawa **S**akata - matrix

NEUTRINO OSCILLATIONS

Flavour

Mass states

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix}$$

$$\Delta m^2 \approx 2.4 \cdot 10^{-3} \text{ eV}^2$$

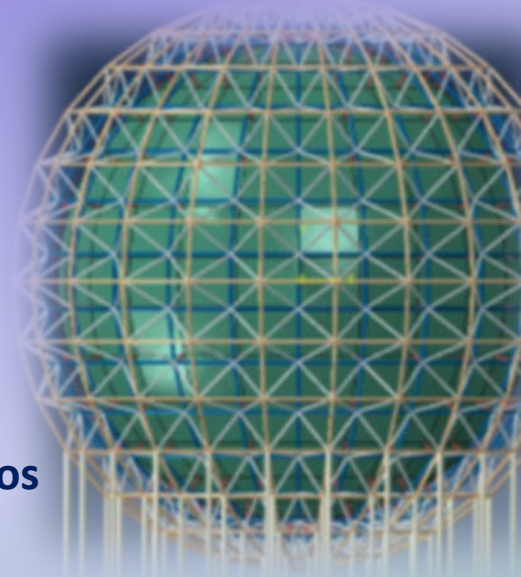
$$\theta_{23} \approx 45^\circ$$

Atmospheric Oscillations

2 3

neutrinos propagate

interact



$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Pontecorvo **M**aki **N**akagawa **S**akata - matrix

NEUTRINO OSCILLATIONS

Flavour
Eigenstates

Mass

$$\begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$\Delta m^2 \approx 7.6 \cdot 10^{-5} \text{ eV}^2$$

$$\theta_{23} \approx 35^\circ$$

$e \mu$

Solar Oscillations

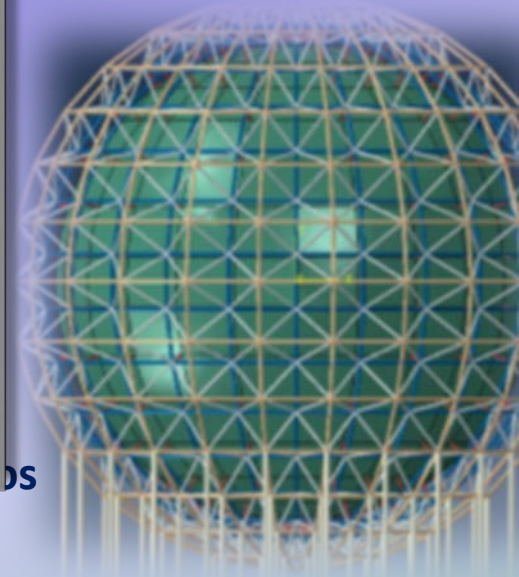
How neutrinos

interact

propagate

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Pontecorvo **M**aki **N**akagawa **S**akata - matrix



NEUTRINO OSCILLATIONS

Flavour		Mass
	$\begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix}$	
	$\Delta m^2 \approx 2.4 \cdot 10^{-3} \text{ eV}^2$	
	$\theta_{13} \approx 9^\circ$	

Reactor Oscillations

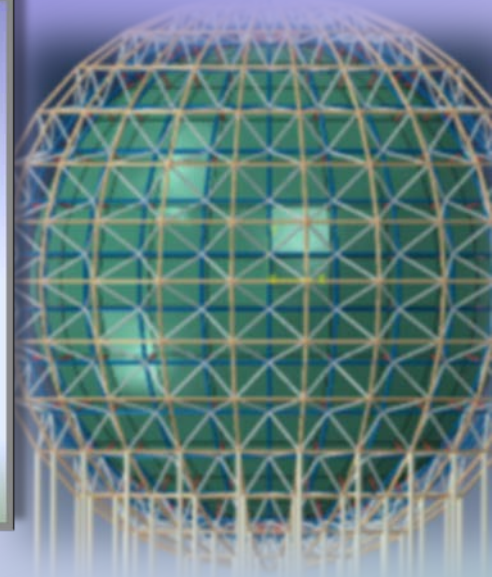
interact

propagate



$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Pontecorvo **M**aki **N**akagawa **S**akata - matrix



NEUTRINO OSCILLATIONS

Special case: 2 flavours only

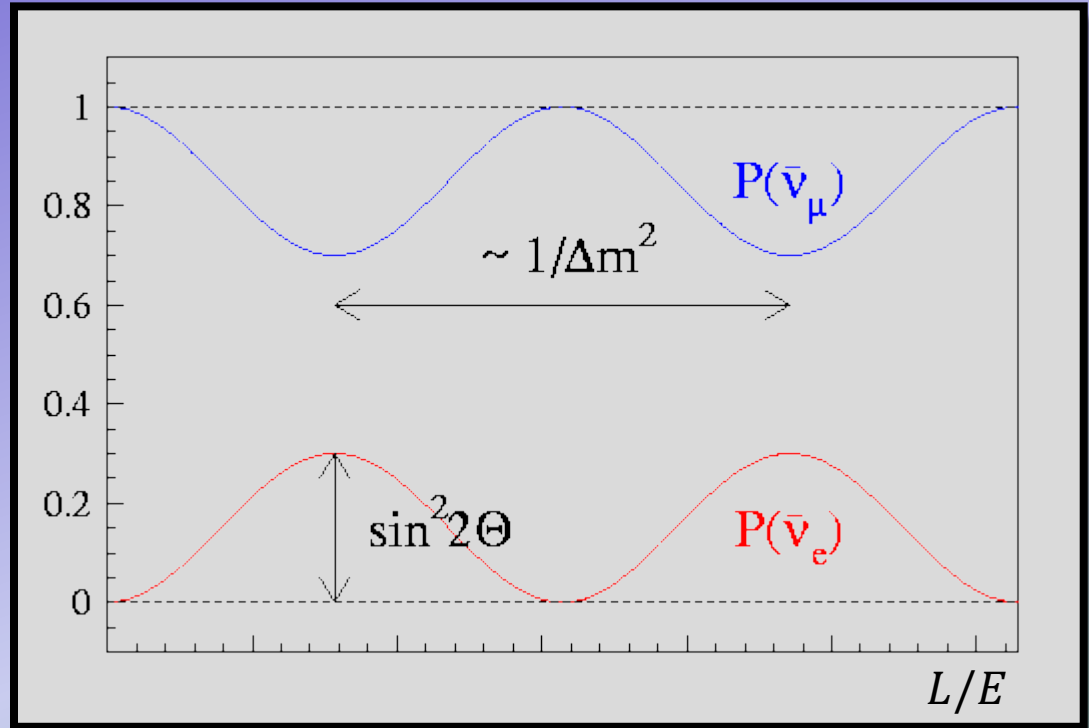
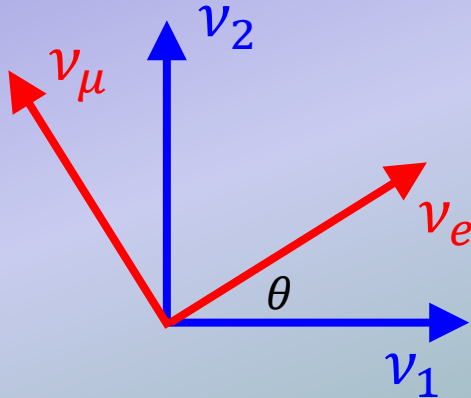
$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

Mass difference $\Delta m^2 = m_2^2 - m_1^2$

Mixing angle θ

Oscillation length:

$$\frac{4\pi E_\nu}{\Delta m^2} \approx 2.48 \text{ m} \frac{E_\nu [\text{MeV}]}{\Delta m^2 [\text{eV}^2]}$$



Disappearance: $P(\nu_\alpha \rightarrow \nu_\alpha) = 1 - \sin^2 2\theta \sin^2 \frac{\Delta m^2 L}{4 E_\nu}$

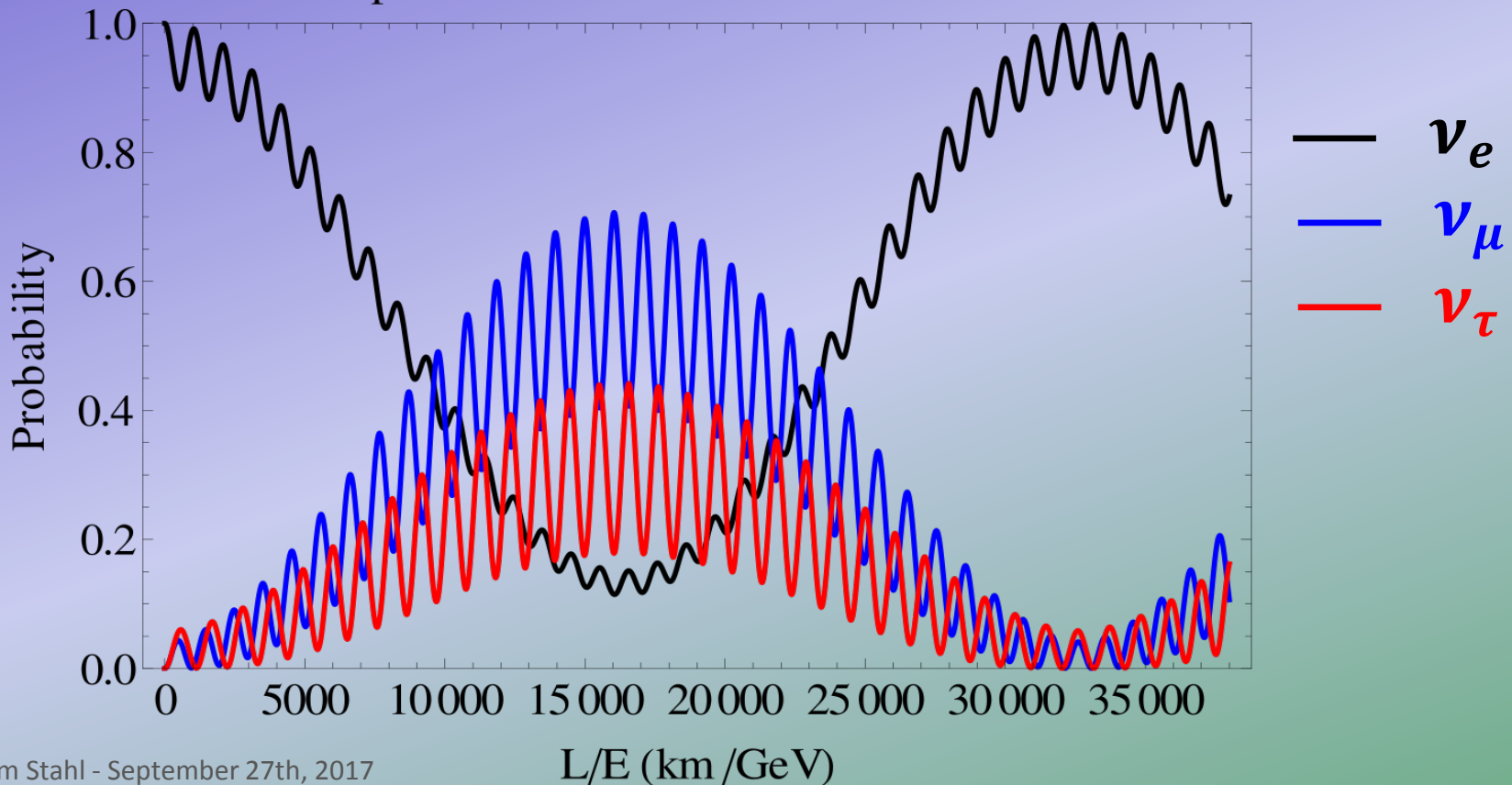
Appearance: $P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2 2\theta \sin^2 \frac{\Delta m^2 L}{4 E_\nu}$

OSCILLATIONEN

$$P_{\alpha \rightarrow \beta} = \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2 \frac{(m_i^2 - m_j^2)L}{4E} + 2 \sum_{i>j} \text{Im}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin \frac{(m_i^2 - m_j^2)L}{2E}$$

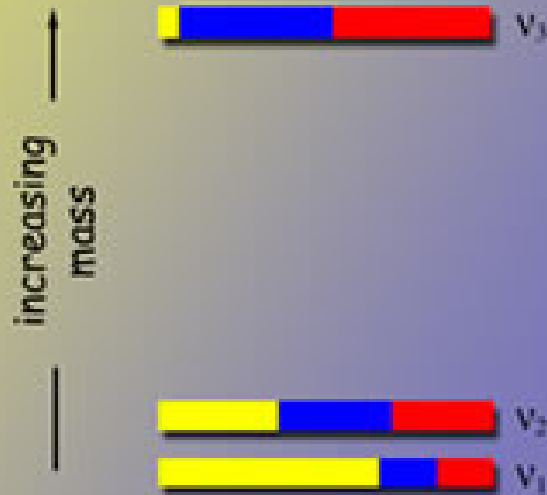
Oscillations require some $m_i \neq 0$

Oscillation probabilities for an initial electron neutrino

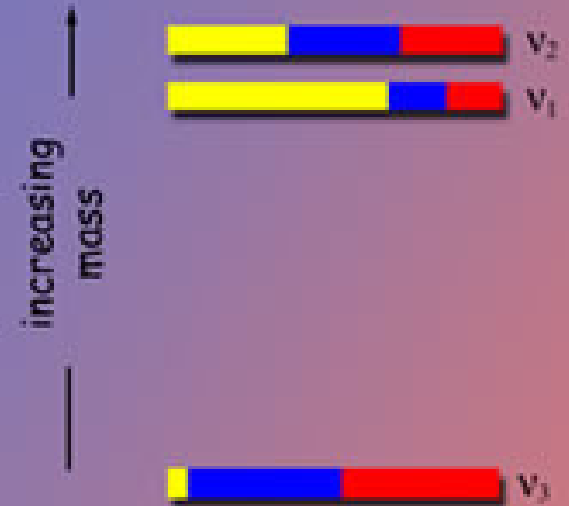







"normal" hierarchy



"inverted" hierarchy



- ν_e  electron neutrino flavor
- ν_μ  muon neutrino flavor
- ν_τ  tau neutrino flavor

Normal or inverted hierarchy ?
Only 3 generations ?
Majorana or Dirac neutrinos ?
CP-violation ?
Absolute mass scale ?

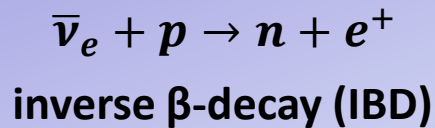
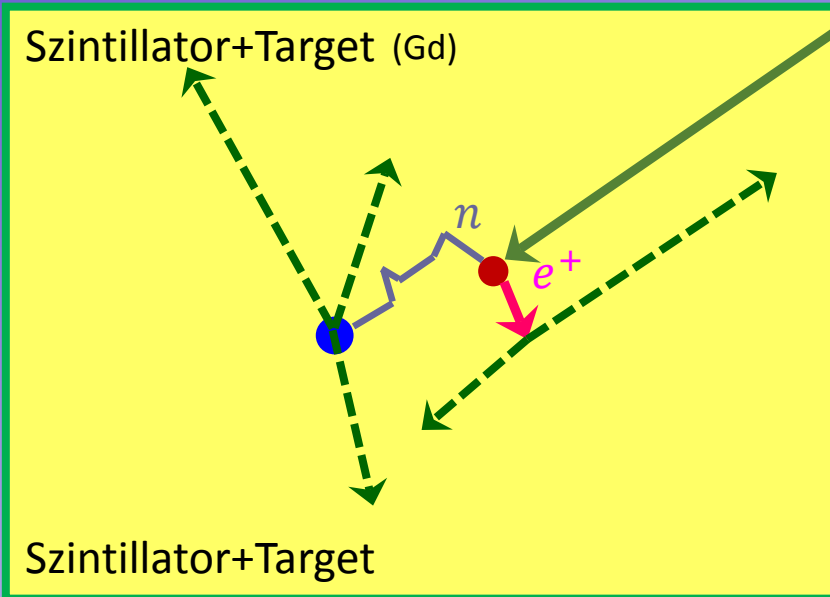
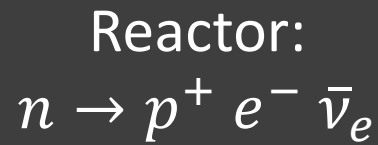
- Osc.: reactor, atmos., beam
- Osc.: precision measurements
- $2\beta 0\nu$ -experiments
- Osc.: beam
- KATRIN, but beyond ?



Reactor Neutrino Experiments



DETECTION



1. prompt event: $e^+ \rightarrow \gamma\gamma$
2. delayed event: n thermalization
+ capture on Gd

Energy Measurement:

1. prompt event

scintillation from positron
gammas from annihilation

$$E(\bar{\nu}_e) = E_{\text{prompt}} + Q - 2m_e$$

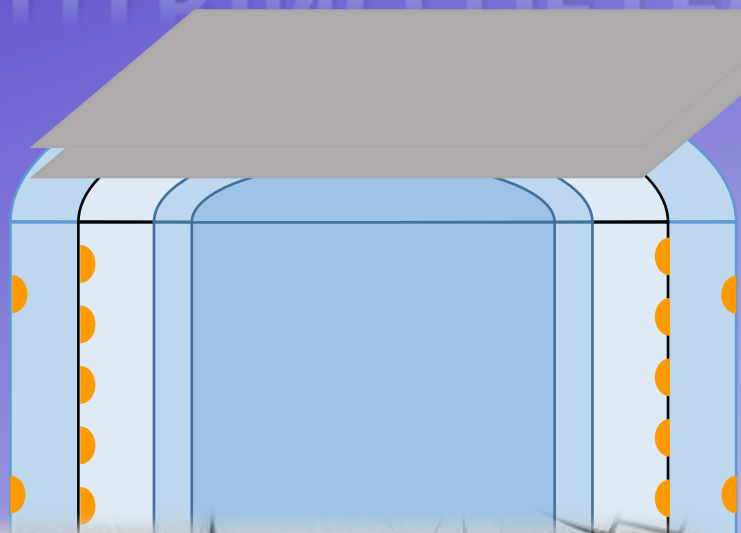
2. delayed event

Gd: 30 μsec delay, 8 MeV

H: 200 μsec delay, 2.2 MeV



NEUTRINO DETECTOR



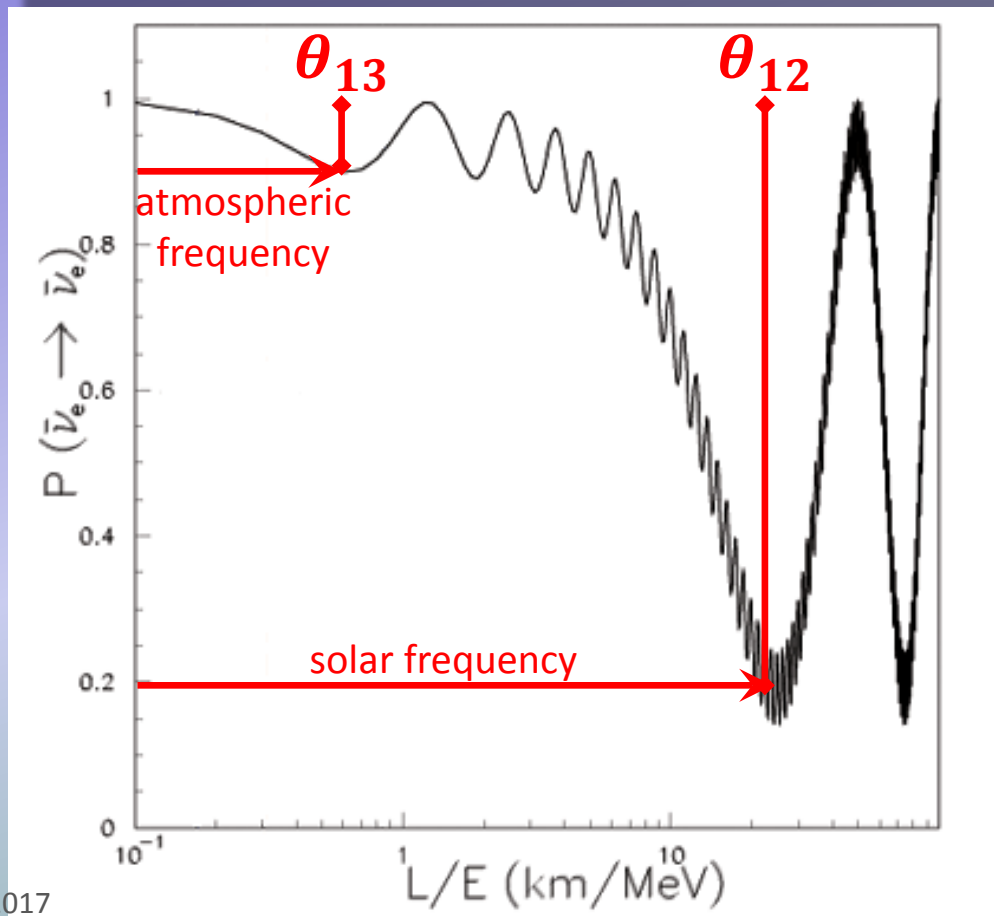
1. **Target:** Scintillator + 0.1 % Gd
2. **γ -catcher:** Scintillator
3. **Buffer:** Oil
4. **Veto:** Water or Scintillator
5. **Muon Tracker**

RENO far detector

OSCILLATION PATTERN

$$P_{\alpha \rightarrow \beta} = \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2 \frac{(m_i^2 - m_j^2)L}{4E} + 2 \sum_{i>j} \text{Im}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin \frac{(m_i^2 - m_j^2)L}{2E}$$

Reactor Experiment
 $\bar{\nu}_e$ -disappearance

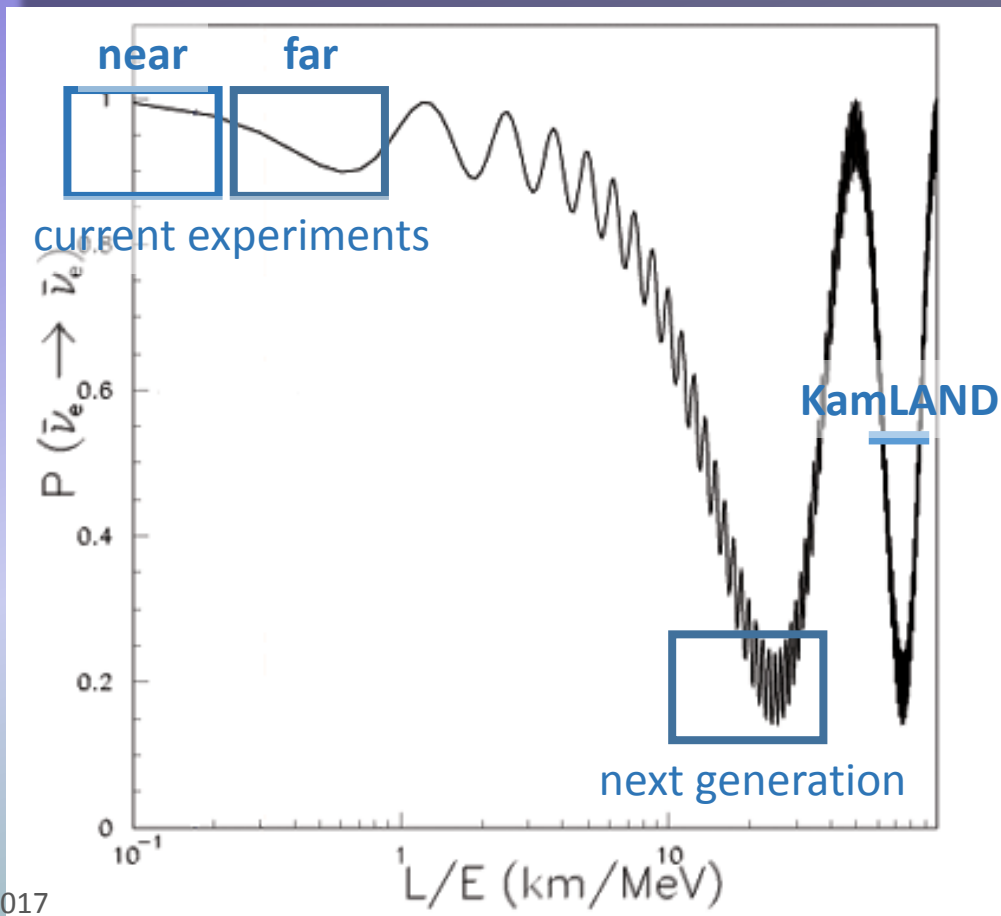


— $\bar{\nu}_e$

OSCILLATION PATTERN

$$P_{\alpha \rightarrow \beta} = \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2 \frac{(m_i^2 - m_j^2)L}{4E} + 2 \sum_{i>j} \text{Im}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin \frac{(m_i^2 - m_j^2)L}{2E}$$

Reactor Experiment
 $\bar{\nu}_e$ -disappearance



— $\bar{\nu}_e$

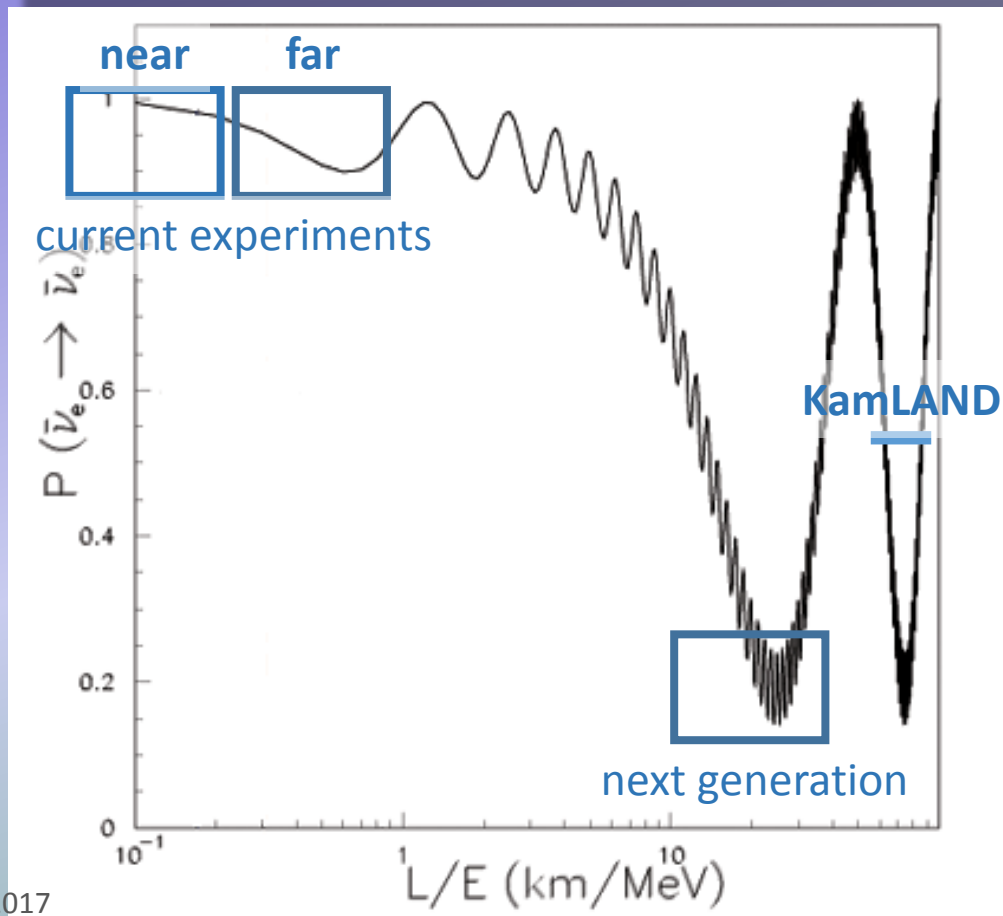
OSCILLATION PATTERN

leading terms:

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} = 1 - \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{ee}^2 L}{4E} - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \frac{\Delta m_{21}^2 L}{4E}$$

$$\sin^2 \frac{\Delta m_{ee}^2 L}{4E} = \cos^2 \theta_{12} \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \theta_{12} \sin^2 \frac{\Delta m_{32}^2 L}{4E}, \quad \Delta m_{ij}^2 = m_i^2 - m_j^2$$

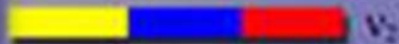
Reactor Experiment
 $\bar{\nu}_e$ -disappearance



— $\bar{\nu}_e$

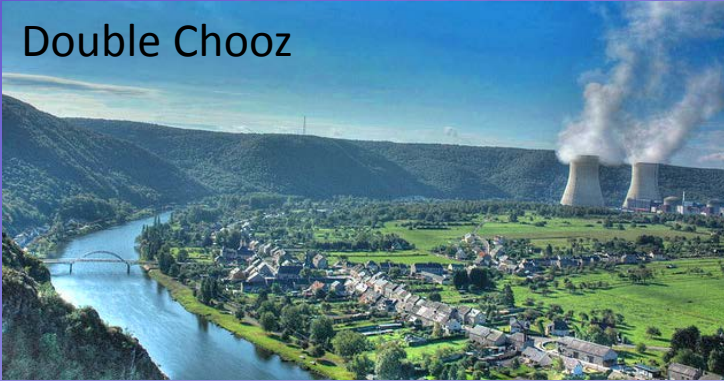


The Mixing Angle θ_{13}



THE 3 EXPERIMENTS

Double Chooz



Reno

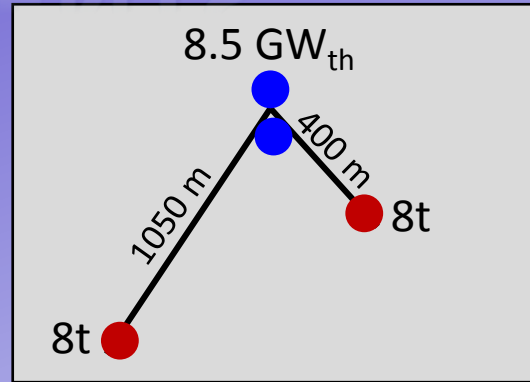
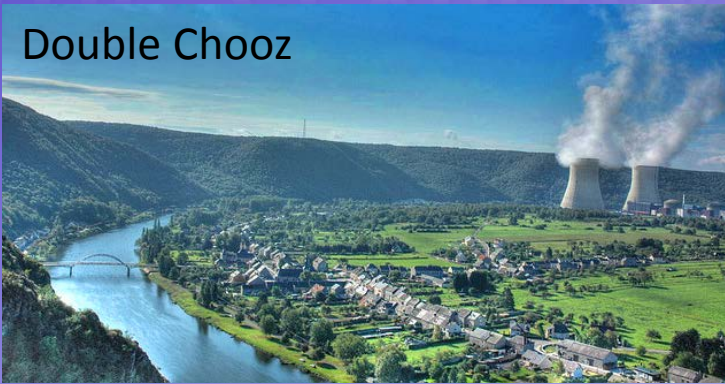


Daya Bay



THE 3 EXPERIMENTS

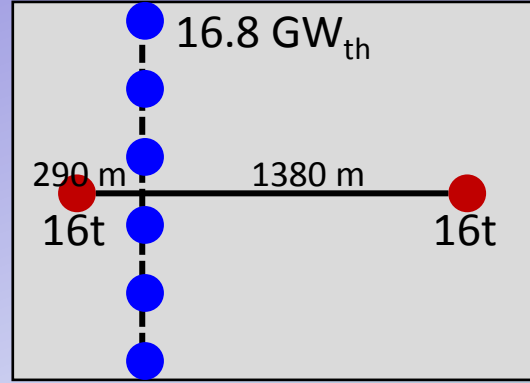
Double Chooz



energy res. 6% (1 MeV)
 overburden: 300m
 statistics: 40 000 IBD

far detector

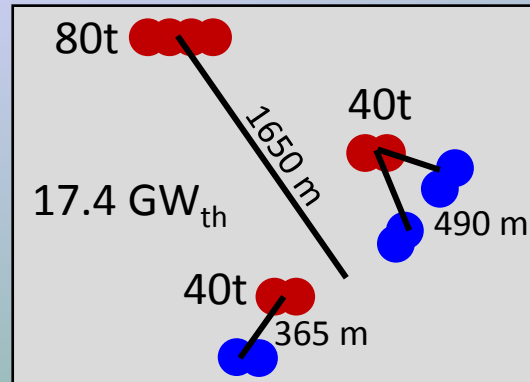
Reno



energy res. 7% (1 MeV)
 overburden: 450m
 statistics: 60 000 IBD

far detector

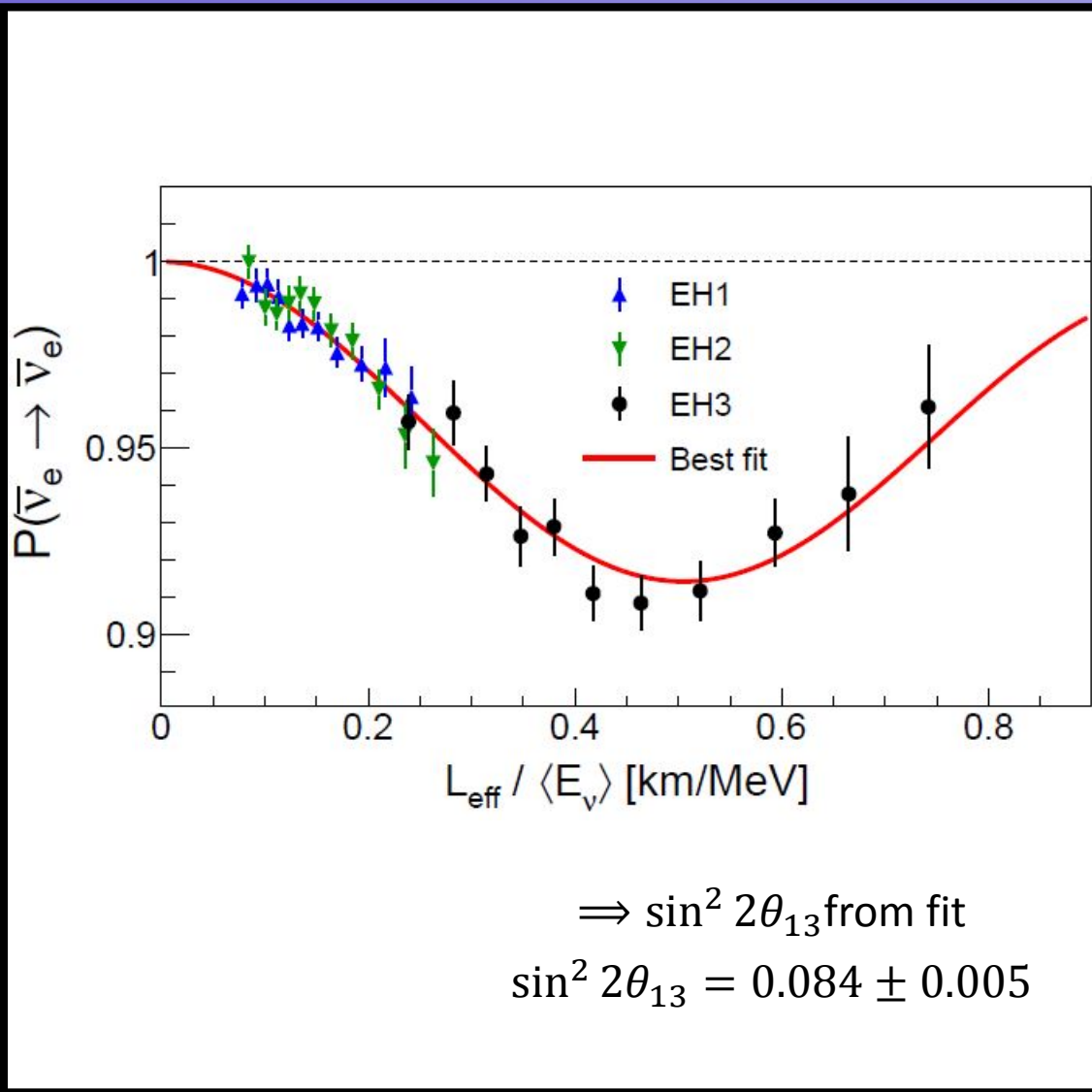
Daya Bay



energy res. 8% (1 MeV)
 overburden: 860m
 statistics: 150 000 IBD

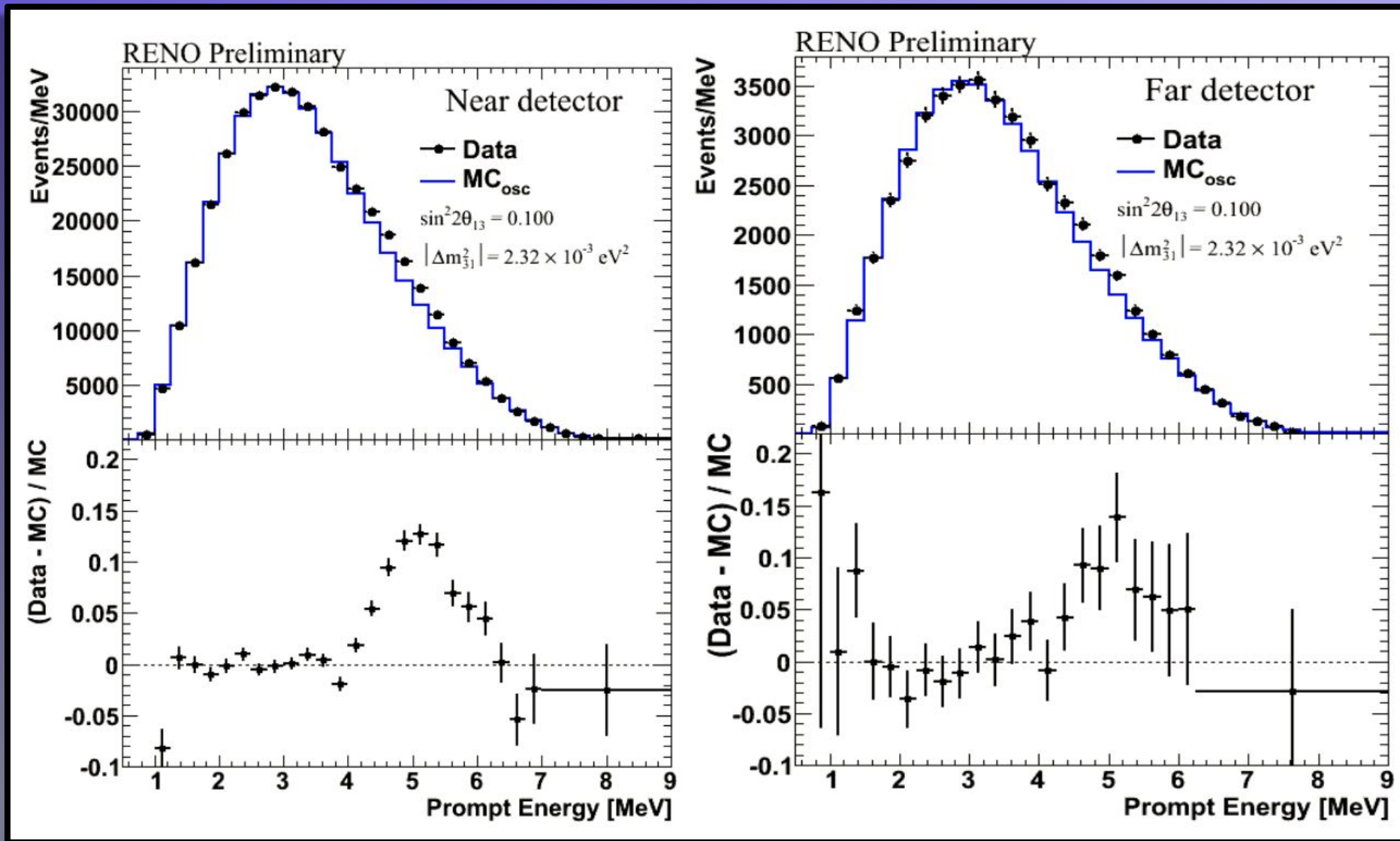
far detector

MEASURED SPECTRUM



F.P. An et al.
arXiv 1505.03456
May, 2015

THE 5 MeV EXCESS



S.-H. Seo (for the RENO collab.); arXiv 1410.7987

WORLD COMPARISON

DC-IV-PRELIMINARY @ CERN

Double Chooz
JHEP 1410, 086 (2014)

Preliminary
(CERN seminar 2016)
 $\sin^2(2\theta_{13}) = (0.119 \pm 0.016)$

Daya Bay
PRL 115, 111802 (2015)

RENO
PRL 116 211801(2016)

T2K
PRD 91, 072010 (2015)

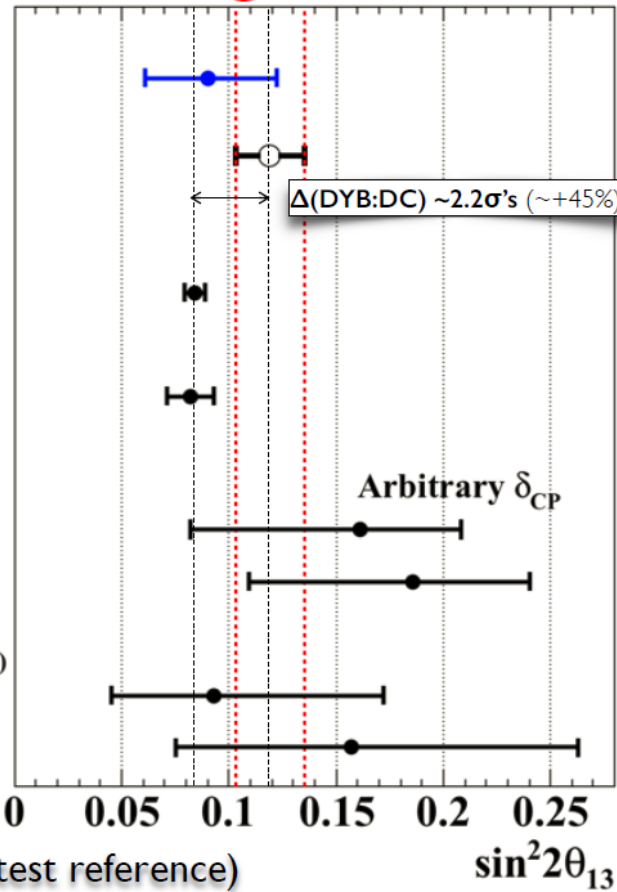
$\Delta m_{32}^2 > 0$

$\Delta m_{32}^2 < 0$

NOvA
Preliminary (private communication)

$\Delta m_{32}^2 > 0$

$\Delta m_{32}^2 < 0$



(Many thanks to NOvA: latest reference)

$$\sin^2 2\theta_{13} = 0.119 \pm 0.016$$

$$\sin^2 2\theta_{13} = 0.084 \pm 0.005$$

$$\sin^2 2\theta_{13} = 0.086 \pm 0.008$$

$$\theta_{13} = 8.4^\circ \pm 0.3^\circ$$



CP-Violation



Matter

=



Anti-Matter?

CP-VIOLATION

$$P_{\alpha \rightarrow \beta} = \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2 \frac{(m_i^2 - m_j^2)L}{4E} + 2 \sum_{i>j} \text{Im}(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin \frac{(m_i^2 - m_j^2)L}{2E}$$

**Example:
Neutrino Beams**

Numerically relevant terms only

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) =$$

$$4 \cdot s_{13}^2 \cdot c_{13}^2 \cdot s_{23}^2 \cdot \sin^2 \frac{\Delta m_{13}^2 L}{4E} \quad \theta_{13}$$

$$+ 8 \cdot c_{13}^2 \cdot s_{12} s_{13} s_{23} \cdot (c_{12} c_{23} \cdot \cos \delta - s_{12} s_{13} s_{23}) \cdot \cos \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \sin \frac{\Delta m_{12}^2 L}{4E} \quad \text{CP-even}$$

$$+ 8 \cdot c_{13}^2 \cdot c_{12} c_{23} s_{12} s_{13} s_{23} \cdot \sin \delta \cdot \sin \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \sin \frac{\Delta m_{12}^2 L}{4E} \quad \text{CP-odd}$$

$$+ 4 \cdot s_{12}^2 \cdot c_{13}^2 \cdot (c_{12}^2 c_{23}^2 + s_{12}^2 s_{23}^2 s_{13}^2 - 2c_{12} c_{23} s_{12} s_{13} s_{23} \cos \delta) \cdot \sin \frac{\Delta m_{12}^2 L}{4E} \quad \text{solare Skala}$$

$$+ 8 \cdot c_{13}^2 \cdot s_{13}^2 \cdot s_{23}^2 \cdot \cos \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \frac{a \cdot L}{4E} \cdot (1 - 2s_{13}^2) \quad \text{Materie-Effekt (CP-odd)}$$

Jarlskog's Determinant for neutrinos: $0.28 \sin \delta$ (quarks: $4 \cdot 10^{-5}$)

CP-VIOLATION

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) =$$

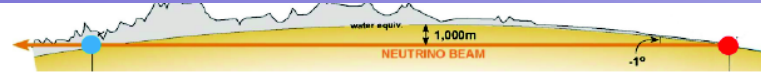
$$4 \cdot s_{13}^2 \cdot c_{13}^2 \cdot s_{23}^2 \cdot \sin^2 \frac{\Delta m_{13}^2 L}{4E}$$

$$+ 8 \cdot c_{13}^2 \cdot s_{12} s_{13} s_{23} \cdot (c_{12} c_{23} \cos \delta - s_{12} s_{13} s_{23}) \cdot \cos \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \sin \frac{\Delta m_{12}^2 L}{4E} \quad \text{CP-even}$$

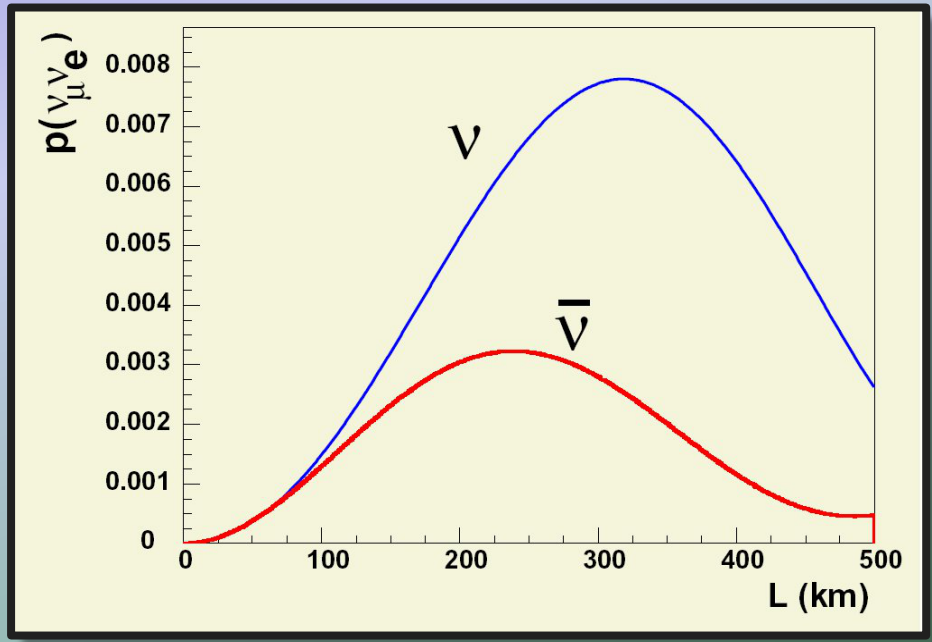
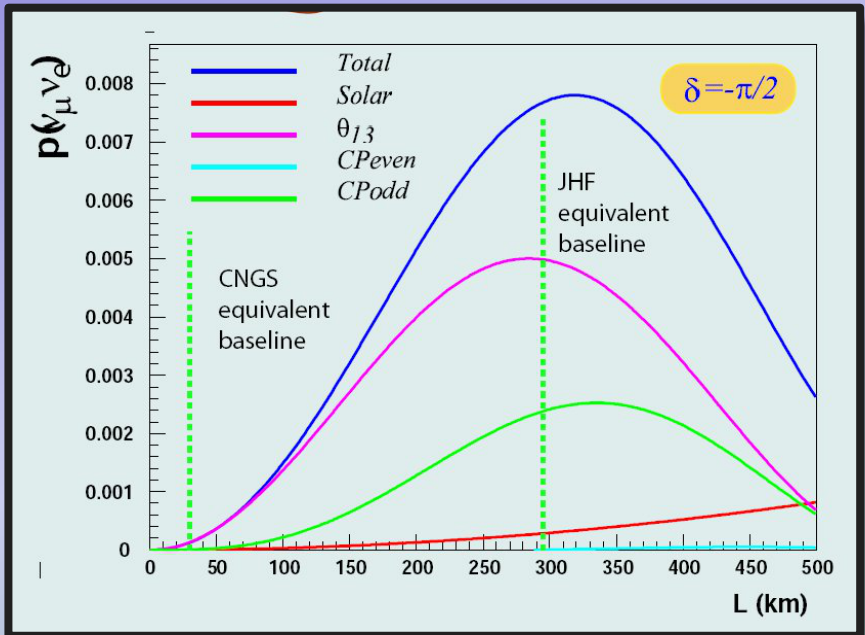
$$\pm 8 \cdot c_{13}^2 \cdot c_{12} c_{23} s_{12} s_{13} s_{23} \sin \delta \cdot \sin \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \sin \frac{\Delta m_{12}^2 L}{4E} \quad \text{CP-odd}$$

$$+ 4 \cdot s_{12}^2 \cdot c_{13}^2 \cdot (c_{12}^2 c_{23}^2 + s_{12}^2 s_{23}^2 s_{13}^2 - 2c_{12} c_{23} s_{12} s_{13} s_{23} \cos \delta) \cdot \sin \frac{\Delta m_{12}^2 L}{4E} \quad \text{solare Skala}$$

$$+ 8 \cdot c_{13}^2 \cdot s_{13}^2 \cdot s_{23}^2 \cdot \cos \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \frac{a \cdot L}{4E} (1 - 2s_{13}^2) \quad \text{Materie-Effekt (CP-odd)}$$



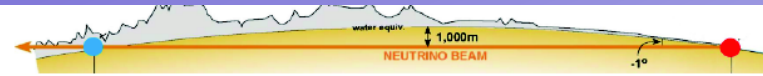
θ_{13}



CP-VIOLATION

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) =$$

$$\begin{aligned}
 & 4 \cdot \boxed{s_{13}^2} \cdot c_{13}^2 \cdot s_{23}^2 \cdot \sin^2 \frac{\Delta m_{13}^2 L}{4E} && \theta_{13} \\
 & + 8 \cdot c_{13}^2 \cdot s_{12} s_{13} s_{23} \cdot (c_{12} c_{23} \cdot \boxed{\cos \delta} - s_{12} s_{13} s_{23}) \cdot \cos \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \sin \frac{\Delta m_{12}^2 L}{4E} && \text{CP-even} \\
 & \pm 8 \cdot c_{13}^2 \cdot c_{12} c_{23} s_{12} s_{13} s_{23} \cdot \boxed{\sin \delta} \cdot \sin \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \sin \frac{\Delta m_{12}^2 L}{4E} && \text{CP-odd} \\
 & + 4 \cdot \boxed{s_{12}^2} \cdot c_{13}^2 \cdot (c_{12}^2 c_{23}^2 + s_{12}^2 s_{23}^2 s_{13}^2 - 2c_{12} c_{23} s_{12} s_{13} s_{23} \cos \delta) \cdot \sin \frac{\Delta m_{12}^2 L}{4E} && \text{solare Skala} \\
 & \pm 8 \cdot c_{13}^2 \cdot s_{13}^2 \cdot s_{23}^2 \cdot \cos \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \boxed{\frac{a \cdot L}{4E}} (1 - 2s_{13}^2) && \text{Materie-Effekt (CP-odd)}
 \end{aligned}$$



CP-VIOLATION: INDIRECT

$$P(\overline{\nu}_\mu \rightarrow \overline{\nu}_e) =$$

$$4 \cdot \boxed{s_{13}^2} \cdot c_{13}^2 \cdot s_{23}^2 \cdot \sin^2 \frac{\Delta m_{13}^2 L}{4E}$$

reactor

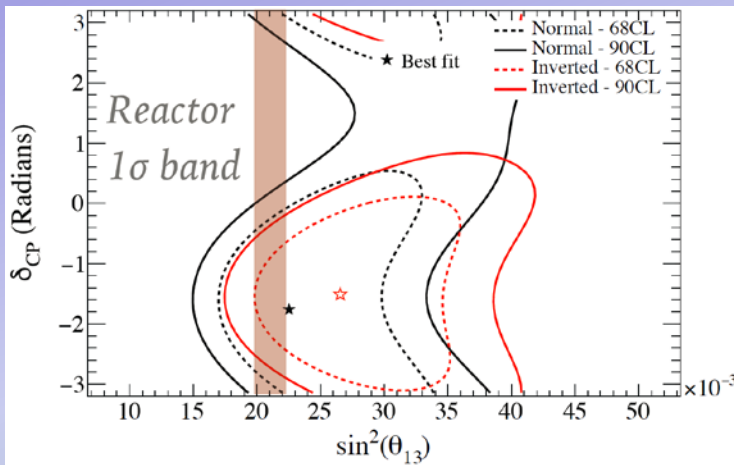


$$+ 8 \cdot c_{13}^2 \cdot s_{12} s_{13} s_{23} \cdot (c_{12} c_{23} \boxed{\cos \delta} - s_{12} s_{13} s_{23}) \cdot \cos \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \sin \frac{\Delta m_{12}^2 L}{4E} \quad \text{CP-even}$$

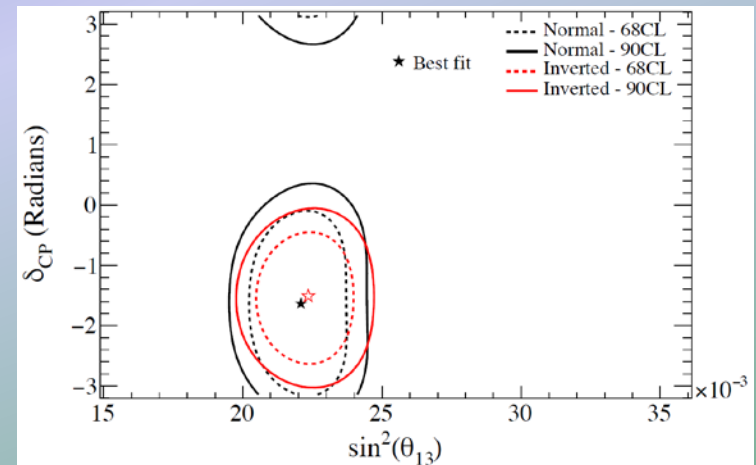
$$\pm 8 \cdot c_{13}^2 \cdot c_{12} c_{23} s_{12} s_{13} s_{23} \boxed{\sin \delta} \cdot \sin \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \sin \frac{\Delta m_{12}^2 L}{4E} \quad \text{CP-odd}$$

$$+ 4 \cdot \boxed{s_{12}^2} \cdot c_{13}^2 \cdot (c_{12}^2 c_{23}^2 + s_{12}^2 s_{23}^2 s_{13}^2 - 2c_{12} c_{23} s_{12} s_{13} s_{23} \cos \delta) \cdot \sin \frac{\Delta m_{12}^2 L}{4E} \quad \text{solare Skala}$$

$$\pm 8 \cdot c_{13}^2 \cdot s_{13}^2 \cdot s_{23}^2 \cdot \cos \frac{\Delta m_{23}^2 L}{4E} \cdot \sin \frac{\Delta m_{13}^2 L}{4E} \cdot \boxed{\frac{a \cdot L}{4E}} (1 - 2s_{13}^2) \quad \text{Materie-Effekt (CP-odd)}$$



T2K (run 1-8)



T2K + reactor exp.



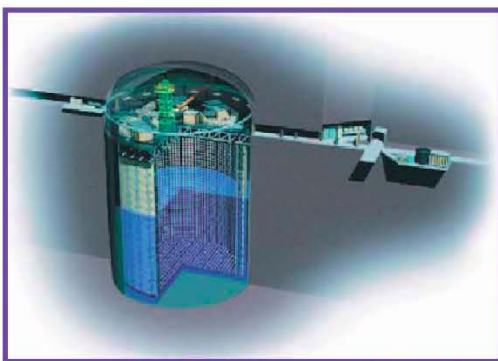
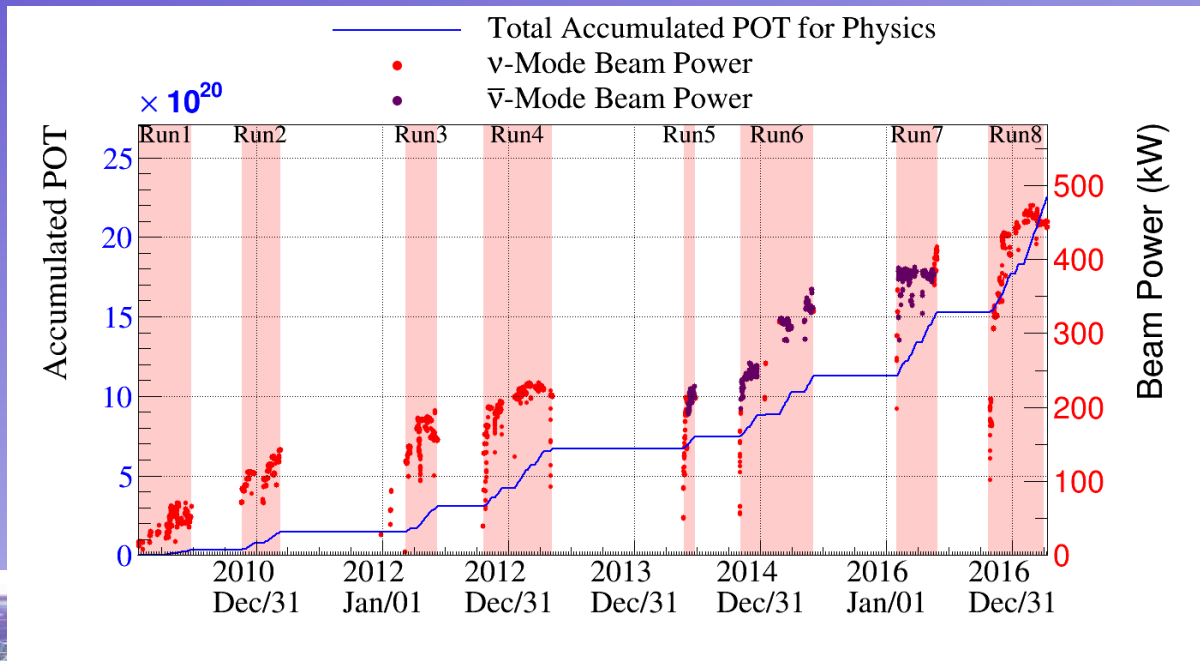
Run 1-7:

$$\nu: 7.48 \times 10^{20} \text{ pot}$$

$$\bar{\nu}: 7.53 \times 10^{20} \text{ pot}$$

Run 8:

$$\nu: 7.48 \times 10^{20} \text{ pot}$$



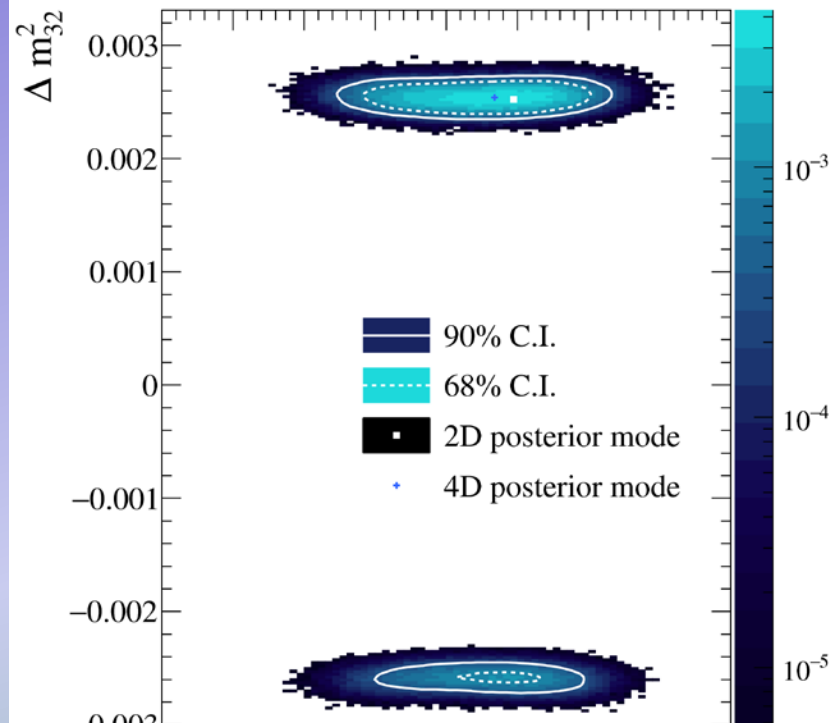
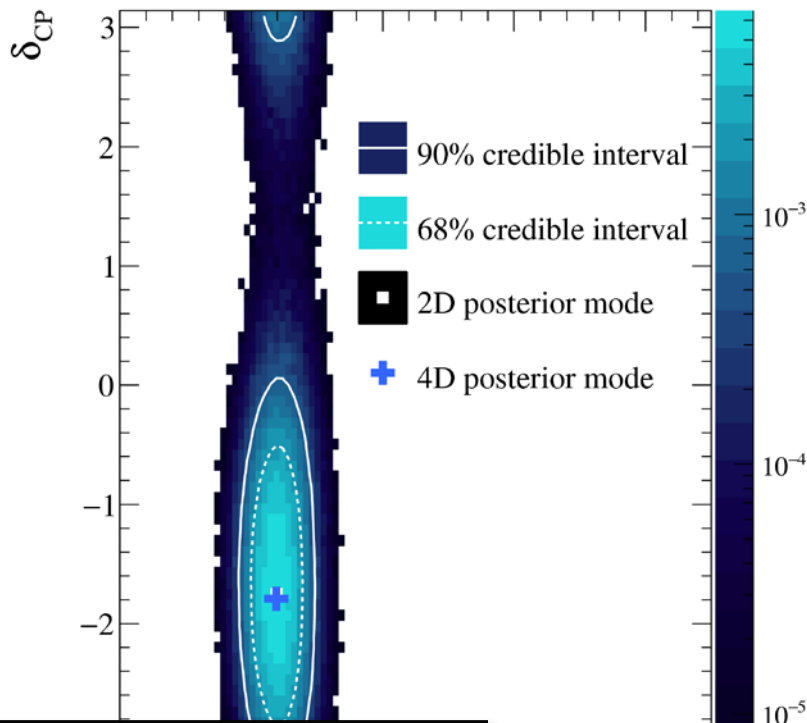
Super-Kamiokande
(ICRR, Univ. Tokyo)



J-PARC Main Ring
(KEK-JAEA, Tokai)



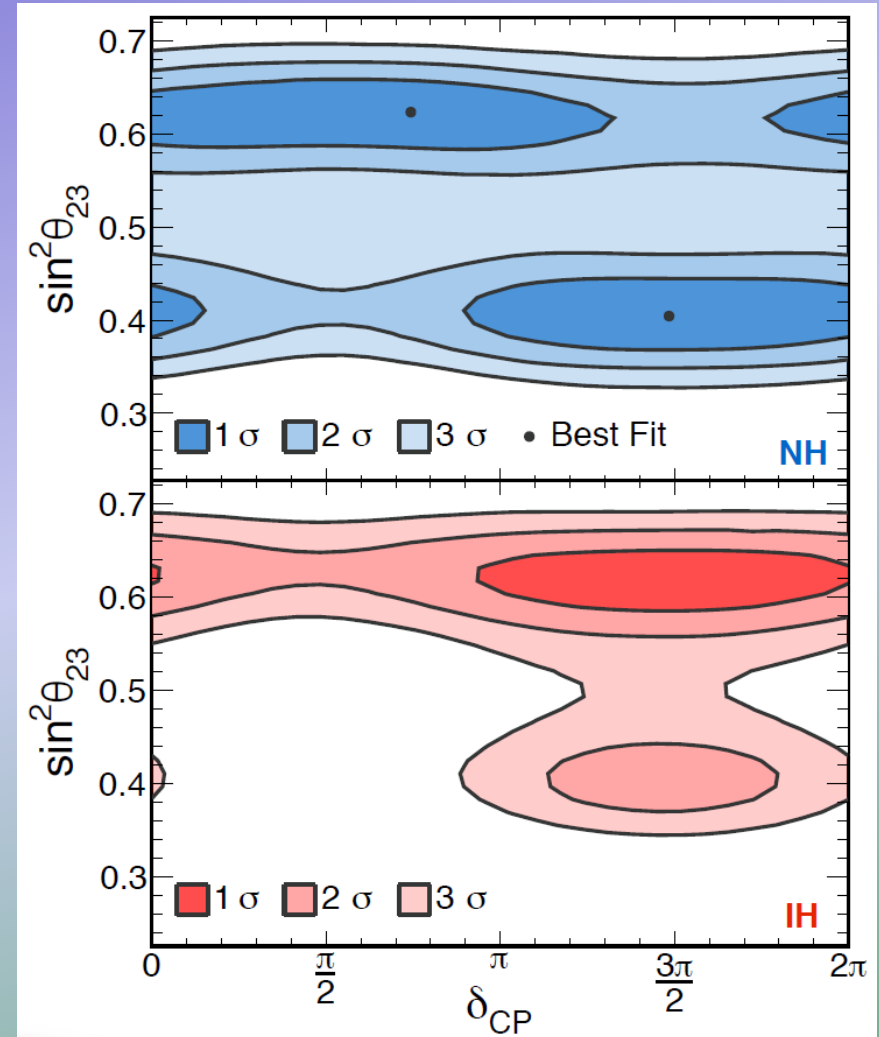
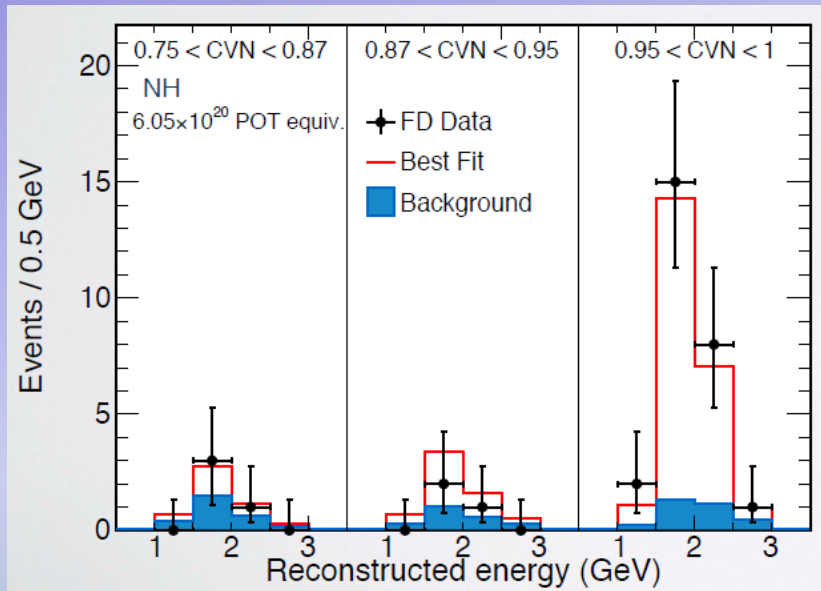
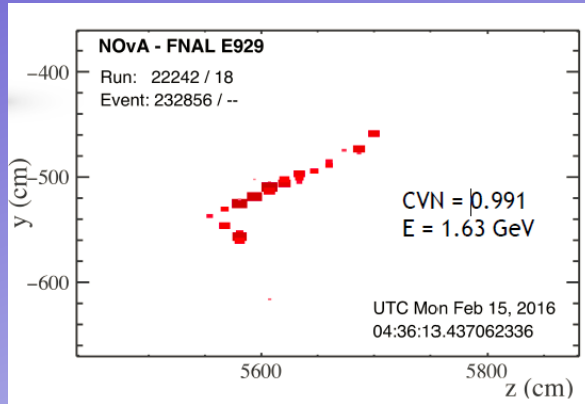
T2K-RESULTS



No CP-violation
 „excluded“
 at 90% c.l.

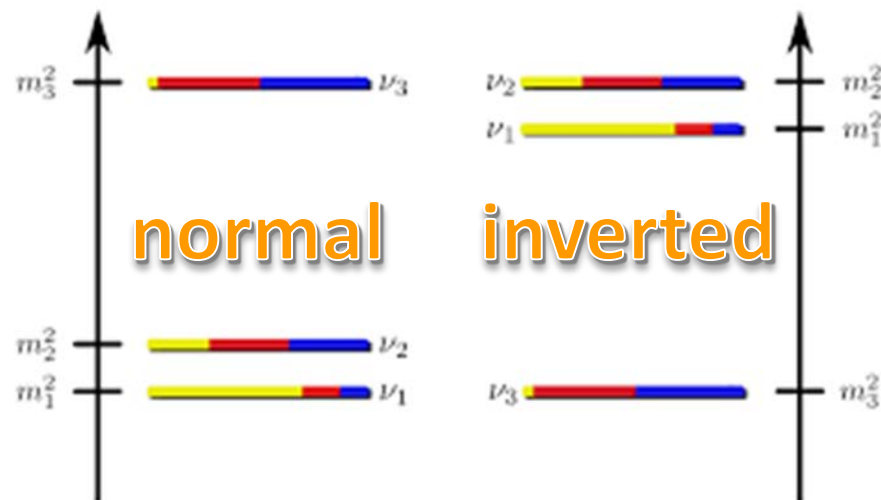
Parameter	Best-fit	$\pm 1\sigma$
δ_{CP}	-1.789	[-2.450; -0.880]
$\sin^2 \theta_{13}$	0.0219	[0.0208; 0.0233]
$\sin^2 \theta_{23}$	0.534	[0.490 ; 0.580]
Δm^2_{32}	$2.539 \times 10^{-3} \text{ eV}^2/c^4$	$[-3.000; -2.952] \times 10^{-3} \text{ eV}^2/c^4$ $[2.424; 2.664] \times 10^{-3} \text{ eV}^2/c^4$

NOVA-APPEARANCE

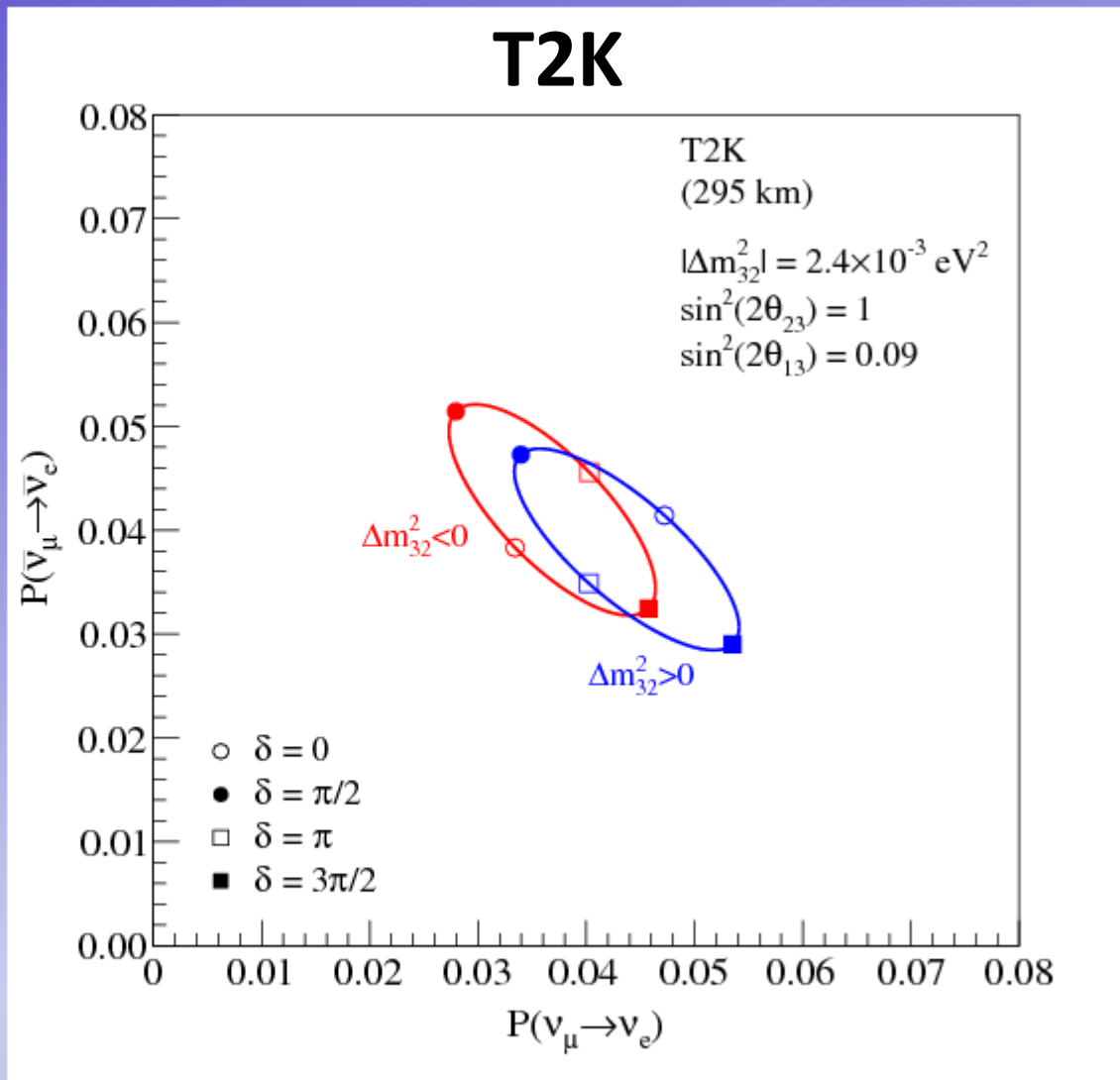




The Mass Hierarchy



MASS HIERARCHY AND CP



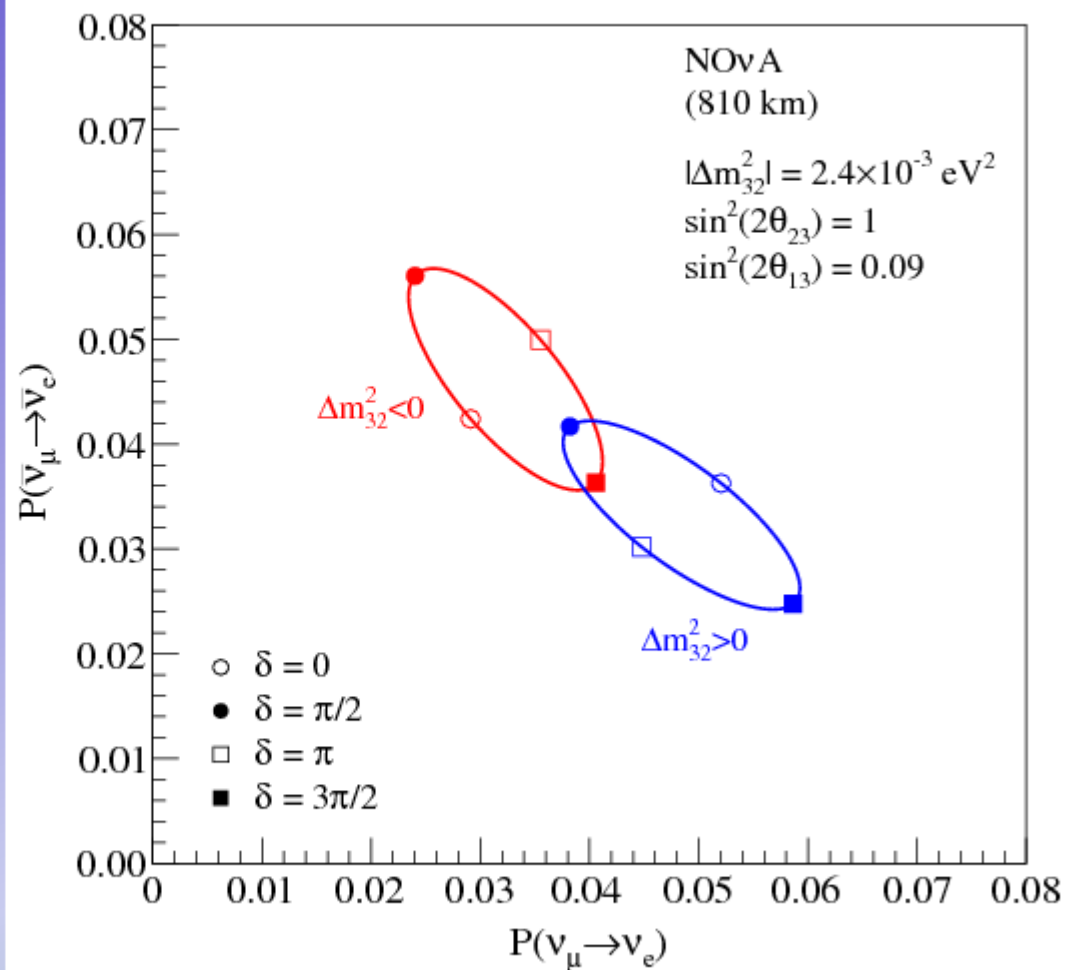
arXiv:1506.07917

„theorist’s plot“

- No experimental uncertainties
- No uncertainties of external parameters

MASS HIERARCHY AND CP

NOVA

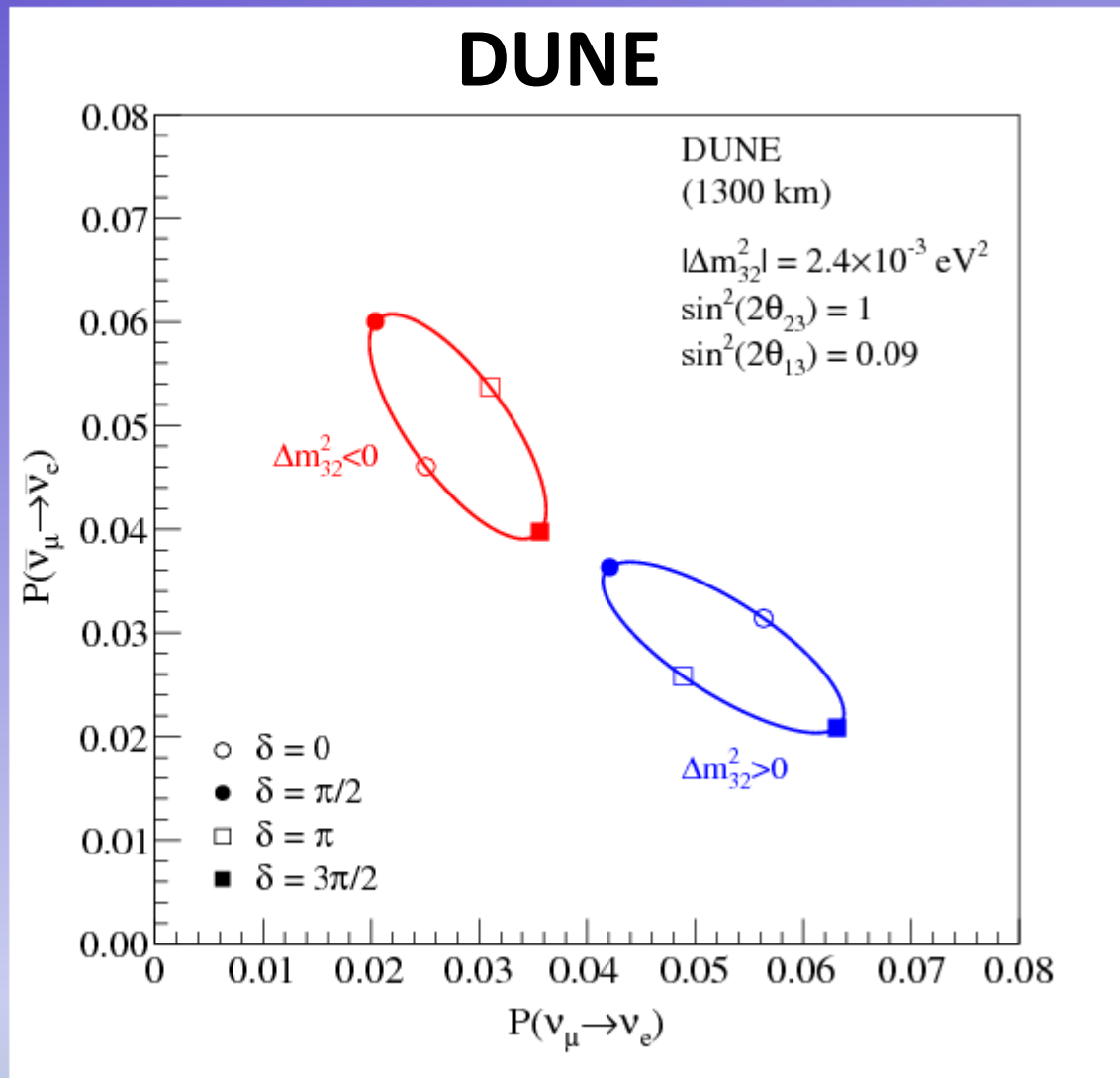


„theorist’s plot“

- No experimental uncertainties
- No uncertainties of external parameters

arXiv:1506.07917

MASS HIERARCHY AND CP



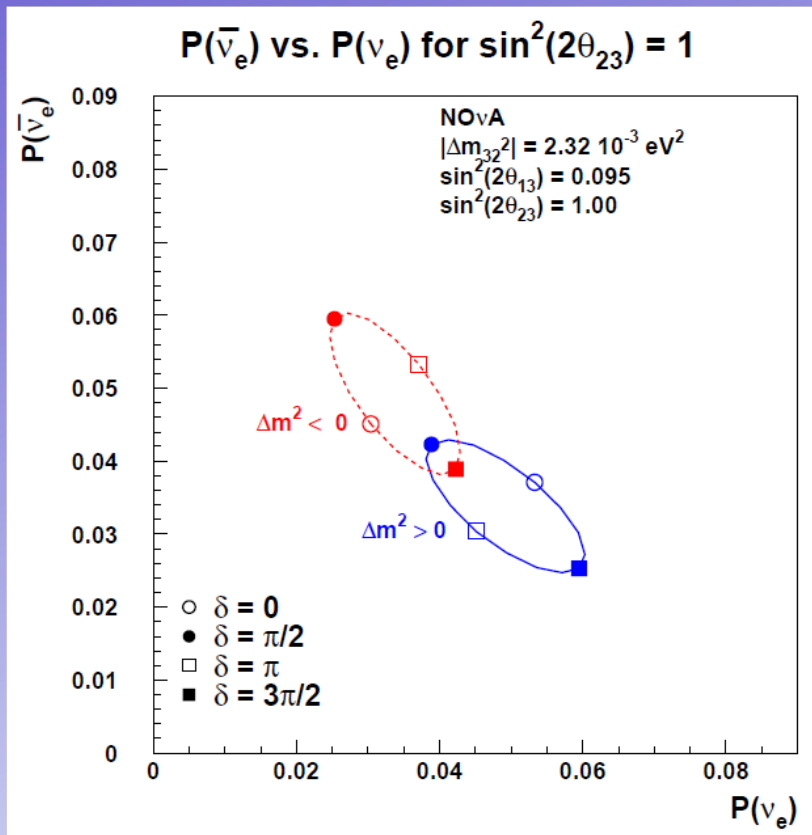
arXiv:1506.07917

„theorist’s plot“

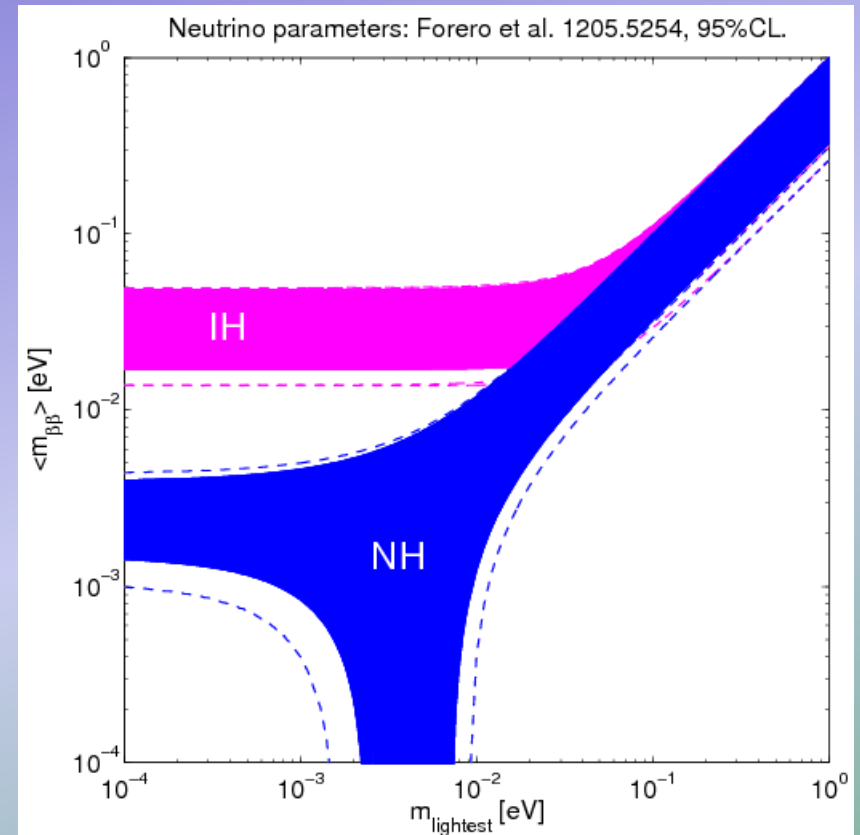
- No experimental uncertainties
- No uncertainties of external parameters

Impact of the Mass Hierarchy

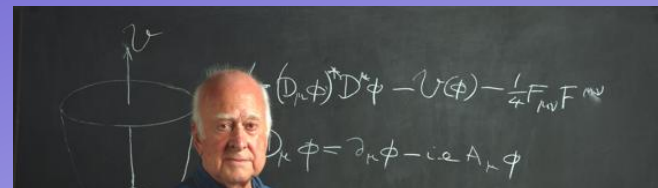
CP Violation



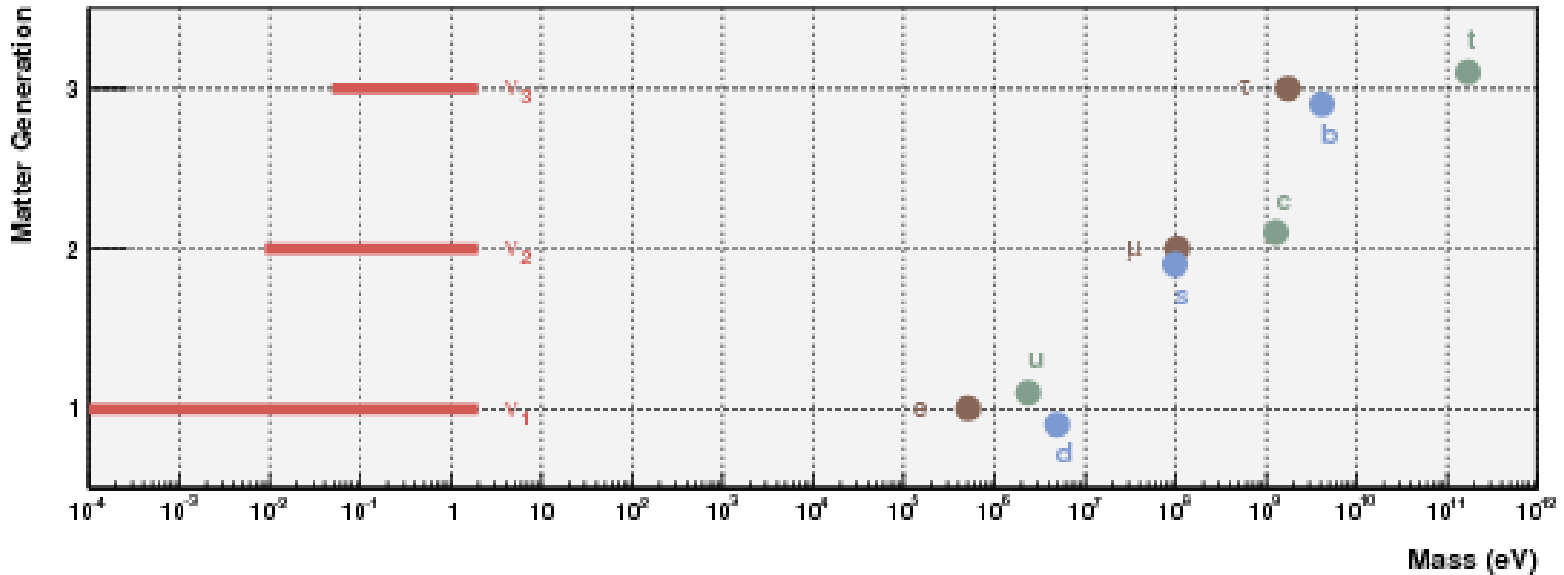
Majorana Neutrinos



Neutrino Masses



Ch



st

$$\begin{aligned}
 & \langle \bar{\Psi}_{L\nu} | 1 | \Psi_{R\nu} \rangle \\
 &= \langle (1 - \gamma_5) \Psi_\nu | 1 | (1 + \gamma_5) \Psi_\nu \rangle \\
 &= \langle \bar{\Psi}_\nu (1 + \gamma_5) | (1 - \gamma_5) \Psi_\nu \rangle \\
 &= \langle \bar{\Psi}_\nu | (1 + \gamma_5) (1 - \gamma_5) | \Psi_\nu \rangle \\
 &= \langle \bar{\Psi}_\nu | 1 | (1 + \gamma_5) \Psi_\nu \rangle
 \end{aligned}$$

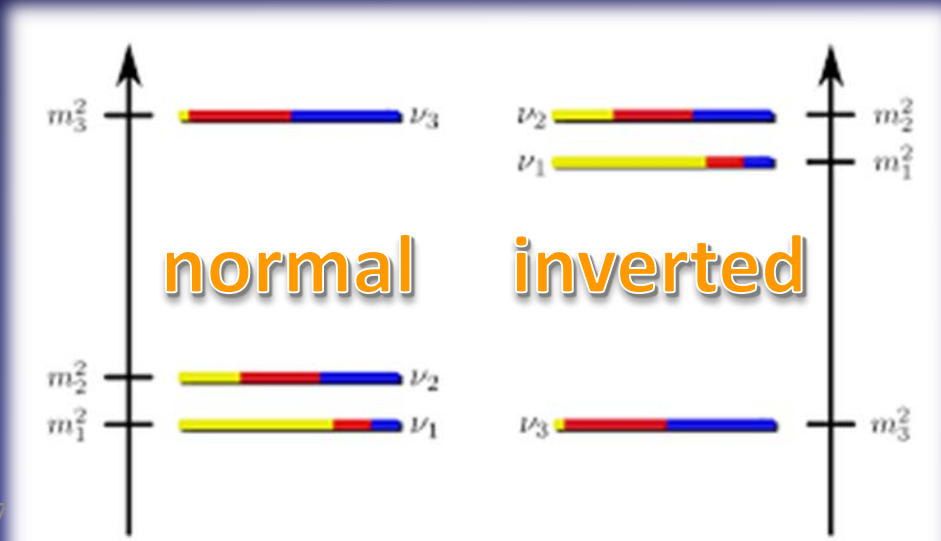
Is there something on top of the Higgs Mechanism?
Do they have the same hierarchy as the other fermions?

It's a sterile component

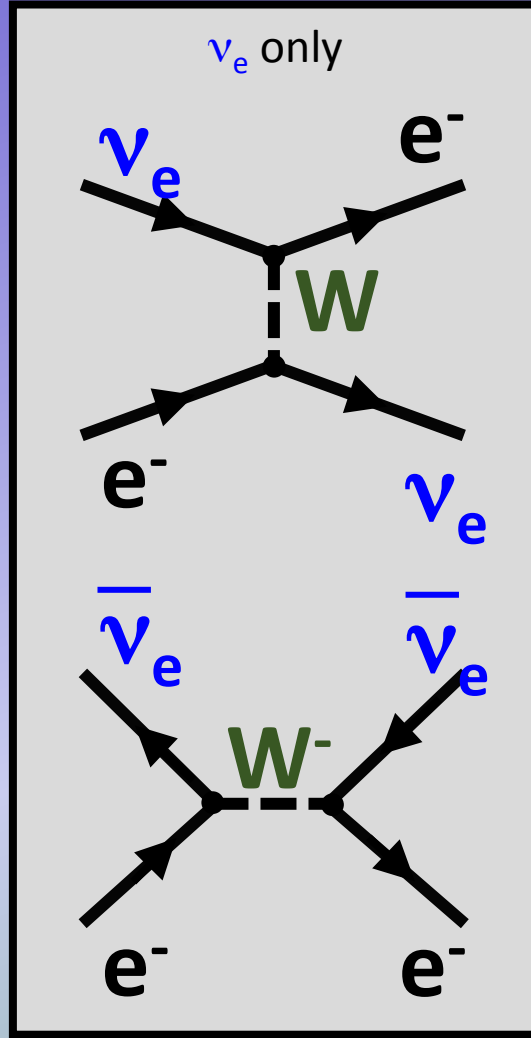
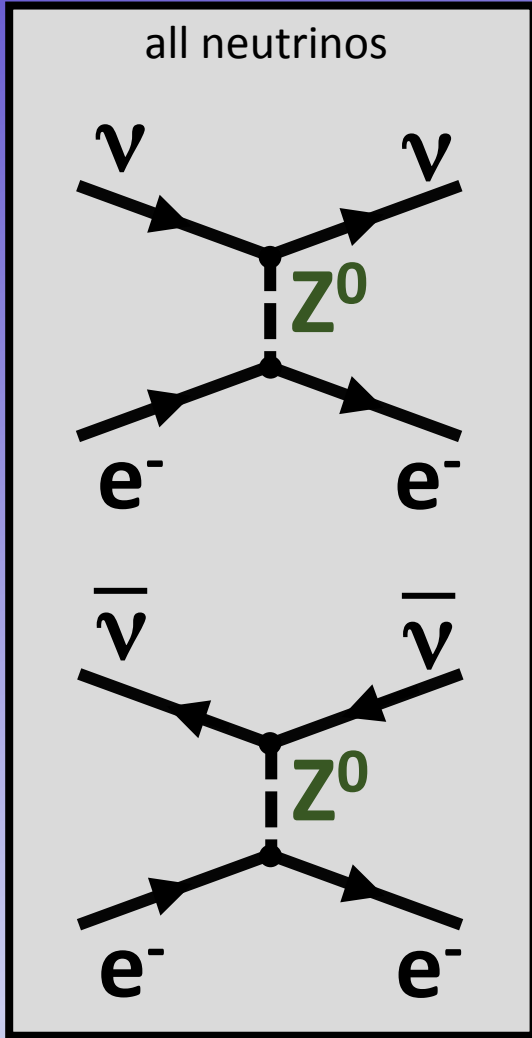
|
chirality flipping



The Mass Hierarchy

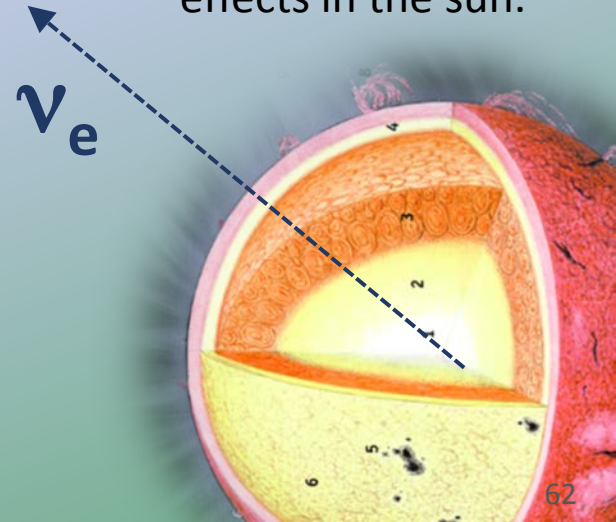


Method 1: MATTER EFFECTS



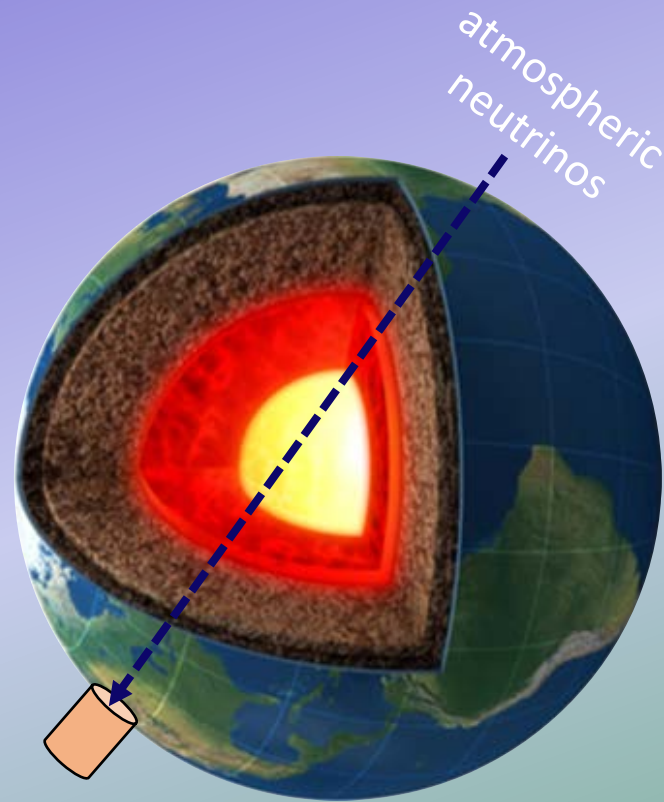
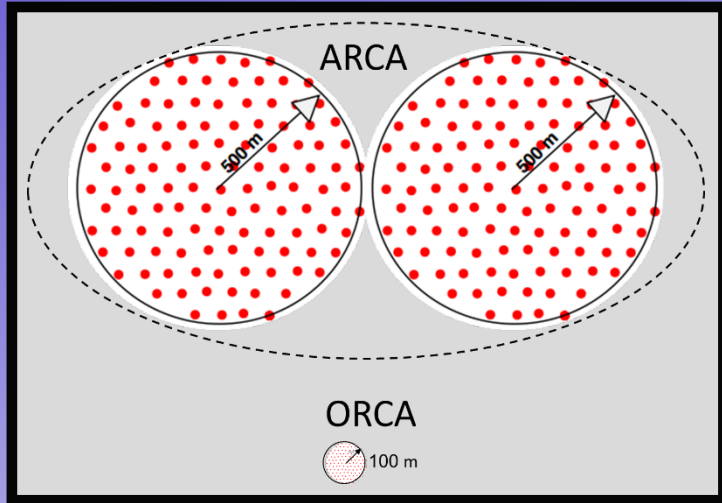
$$P_{\bar{\nu}_\mu \rightarrow \bar{\nu}_e} = \dots + \dots - \dots \frac{aL}{4E} \cos \frac{\Delta m_{23}^2 L}{4E} \sin \frac{\Delta m_{13}^2 L}{4E}$$

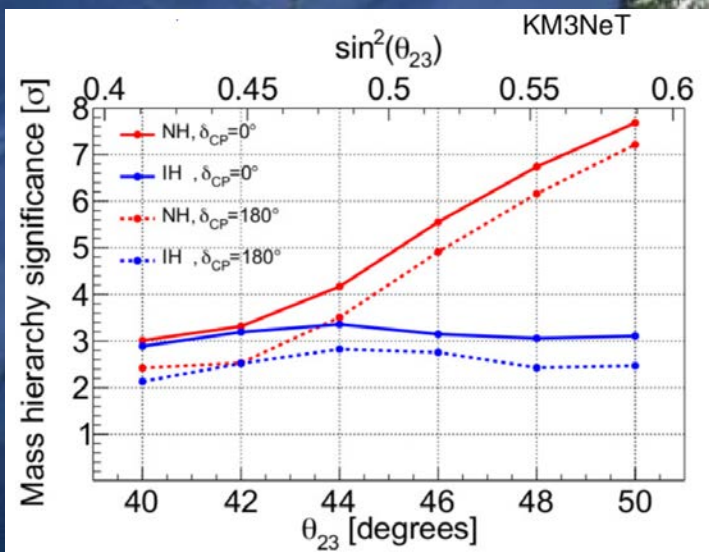
Hierarchy of ν_1 and ν_2
Determined from matter effects in the sun.



MATTER EFFECTS: ORCA

Resonant transition (MSW) near the core of the earth





KM3NeT Phase-1 Infrastructure (March 2016):

- 3 Installation sites
- 2 PMT preparation sites
- 4 DOM integration sites
- 3 DOM integration sites proposed/planned
- 3 base module integration sites
- 3 DU integration sites
- 3 DU test and preparation to deployment sites

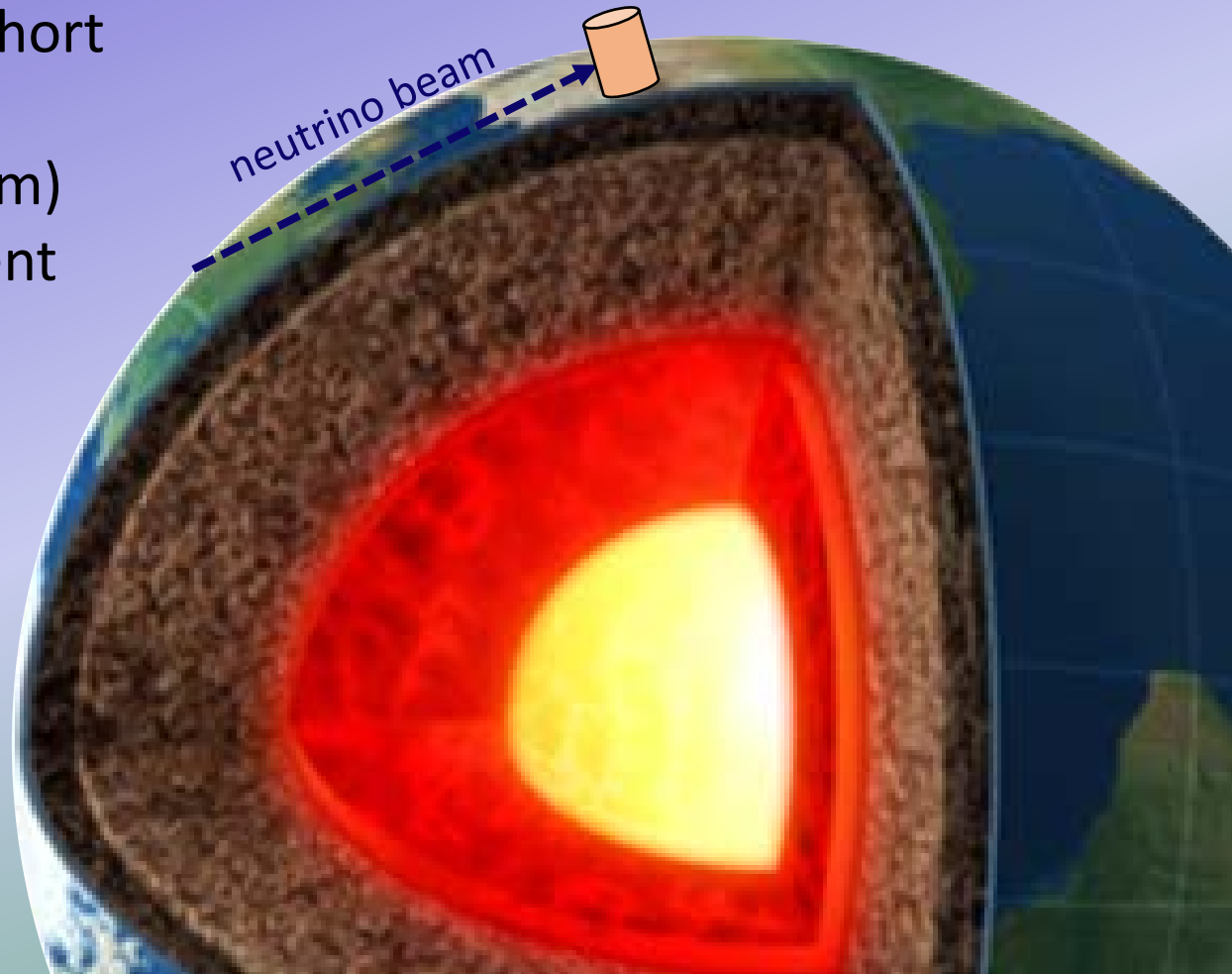


Method 1: MATTER EFFECTS

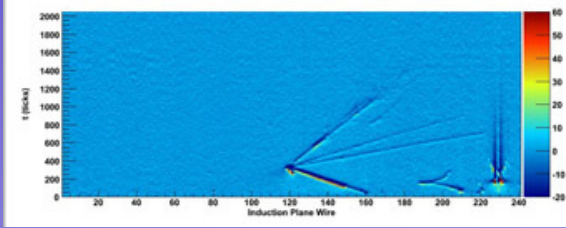
Matter Effect is proportional to L !

Long baseline Beams

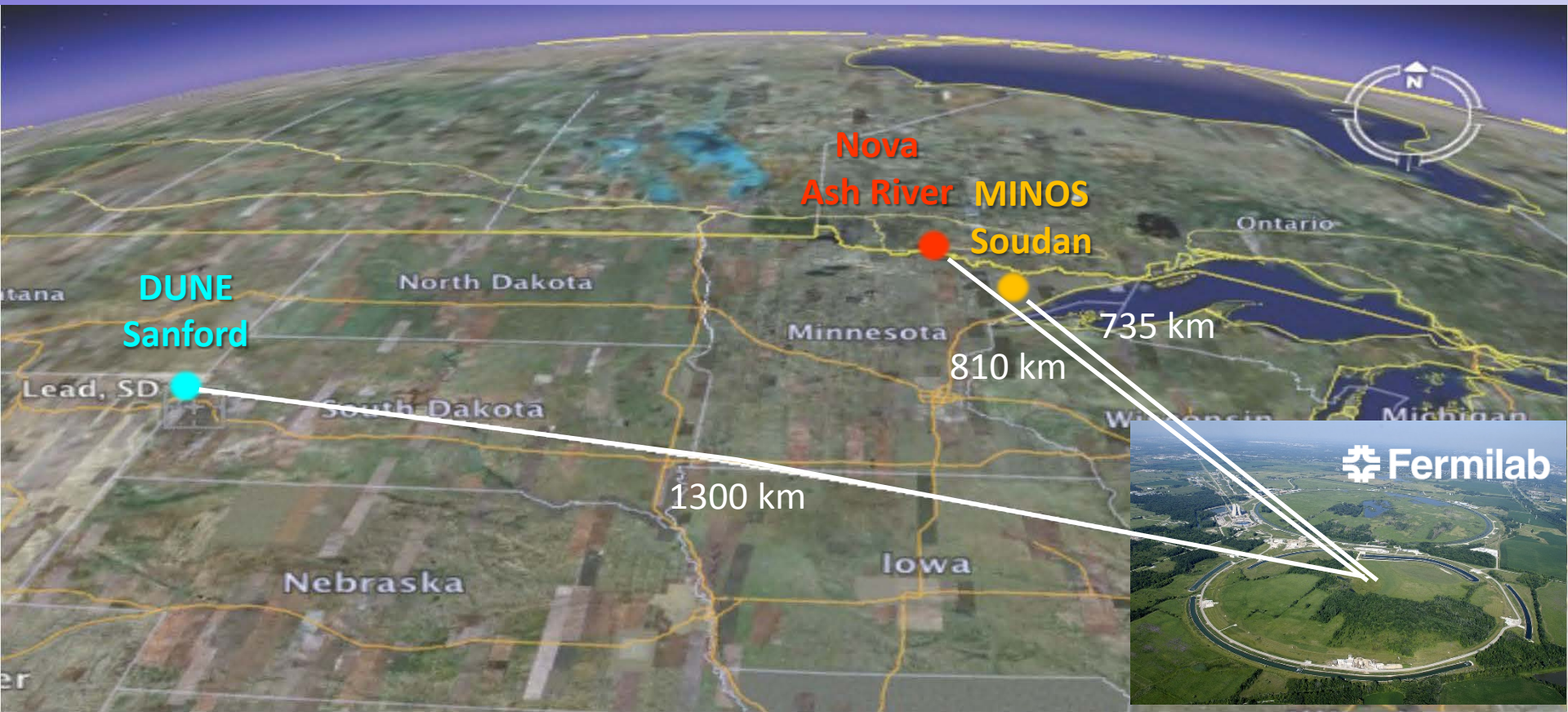
- T2K (295 km) too short
- Nova (810 km) 2σ
- LBNE/DUNE (1300 km) excellent



FUTURE: DUNE



1...4 Liquid Argon TPCs; 10 kt each



DUNE: R&D

ICARUS: 600 t

MicroBooNE: 170 t

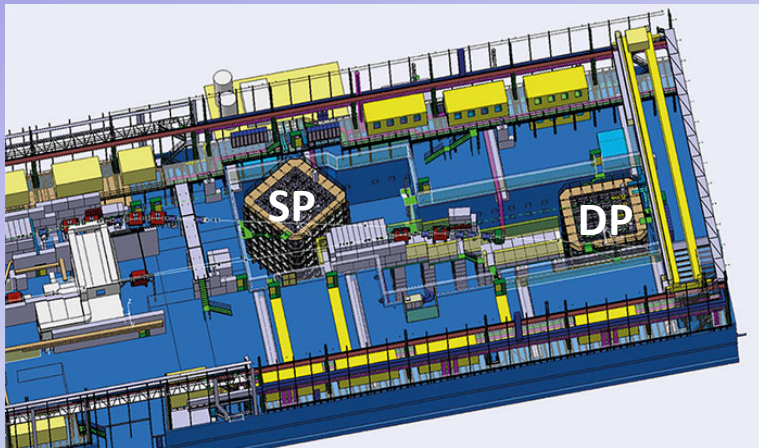
⋮

Prototypes

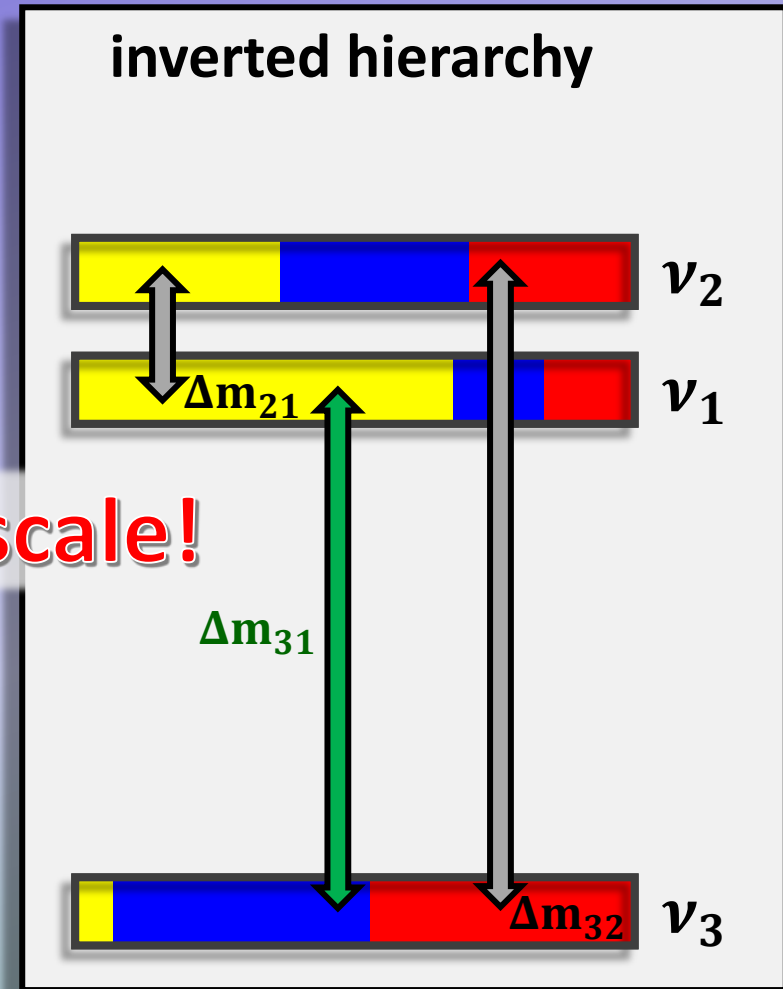
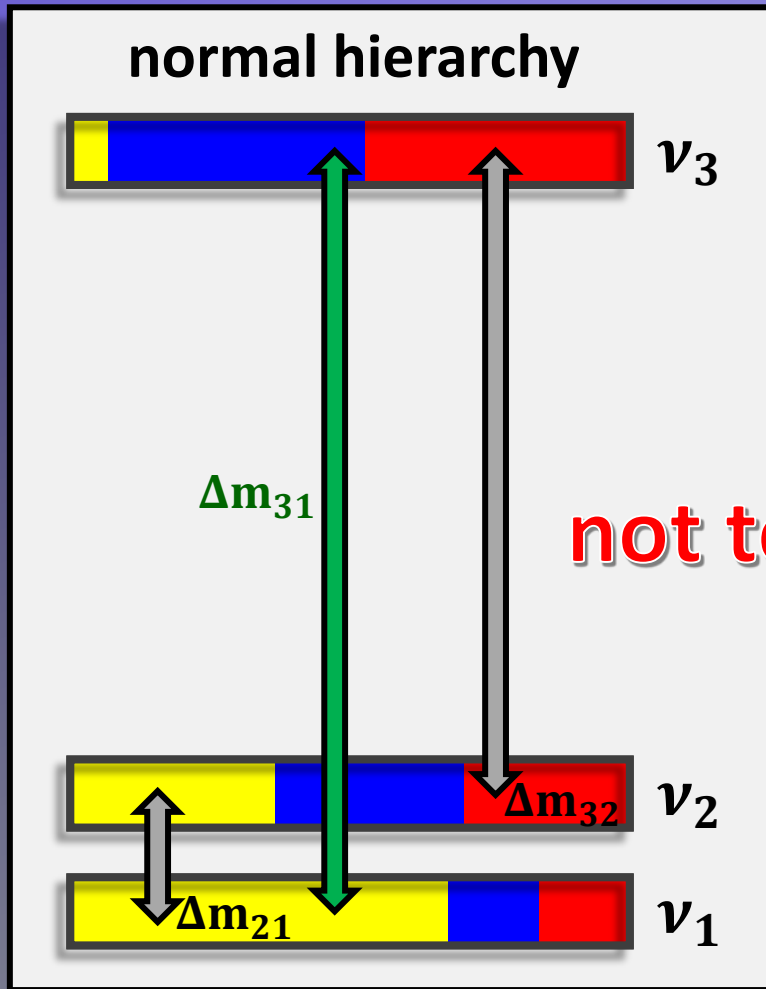
⋮

DUNE: 10 000 t

CERN: Neutrino Platform
600t prototypes DUNE SP/DP



Method 2: 3-Flavour-Interference



not to scale!

$$\Delta m_{31,\text{normal}} > \Delta m_{31,\text{inverted}}$$

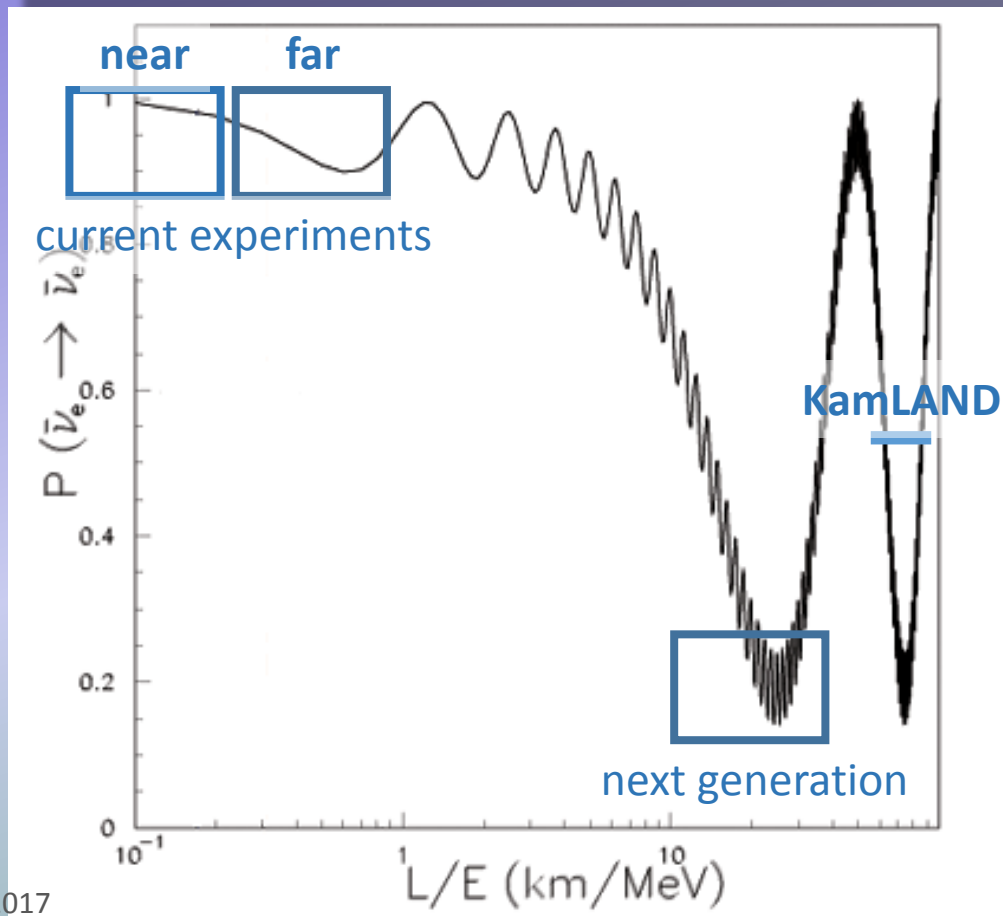
OSCILLATION PATTERN

leading terms:

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} = 1 - \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{ee}^2 L}{4E} - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \frac{\Delta m_{21}^2 L}{4E}$$

$$\sin^2 \frac{\Delta m_{ee}^2 L}{4E} = \cos^2 \theta_{12} \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \theta_{12} \sin^2 \frac{\Delta m_{32}^2 L}{4E}, \quad \Delta m_{ij}^2 = m_i^2 - m_j^2$$

Reactor Experiment
 $\bar{\nu}_e$ -disappearance



— $\bar{\nu}_e$

OSCILLATION PATTERN

Yu-Feng Li et al.,
Phys.Rev. D88 (2013) 013008

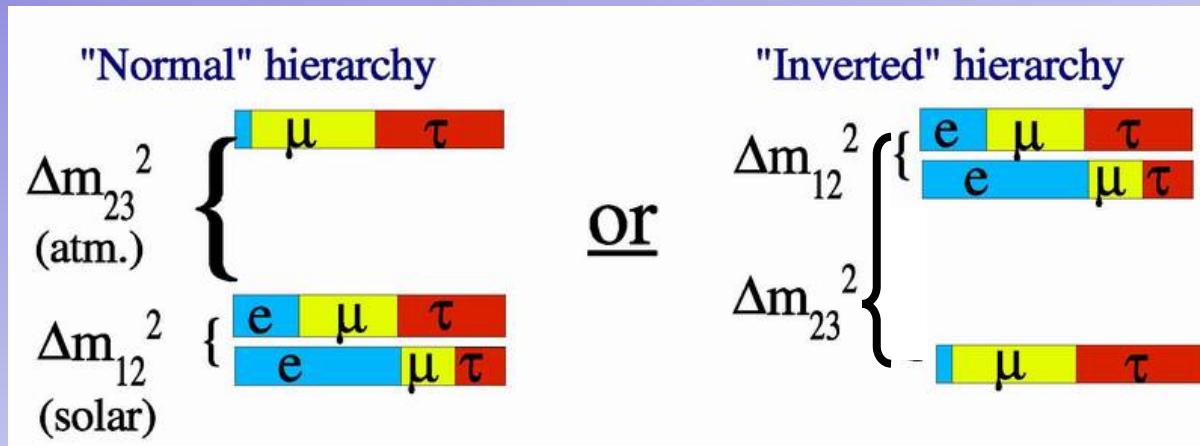
leading terms:

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} = 1 - \cos^2 \theta_{12} \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{31}^2 L}{4E} \quad \leftarrow \text{high frequency}$$

$$- \sin^2 \theta_{12} \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{32}^2 L}{4E} \quad \leftarrow \text{high frequency}$$

$$- \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \frac{\Delta m_{21}^2 L}{4E} \quad \leftarrow \text{low frequency}$$

$$\Delta m_{ij} = m_i^2 - m_j^2$$



$$|\Delta m_{31}^2| = |\Delta m_{23}^2| + |\Delta m_{12}^2|$$

$$|\Delta m_{31}^2| = |\Delta m_{23}^2| - |\Delta m_{12}^2|$$

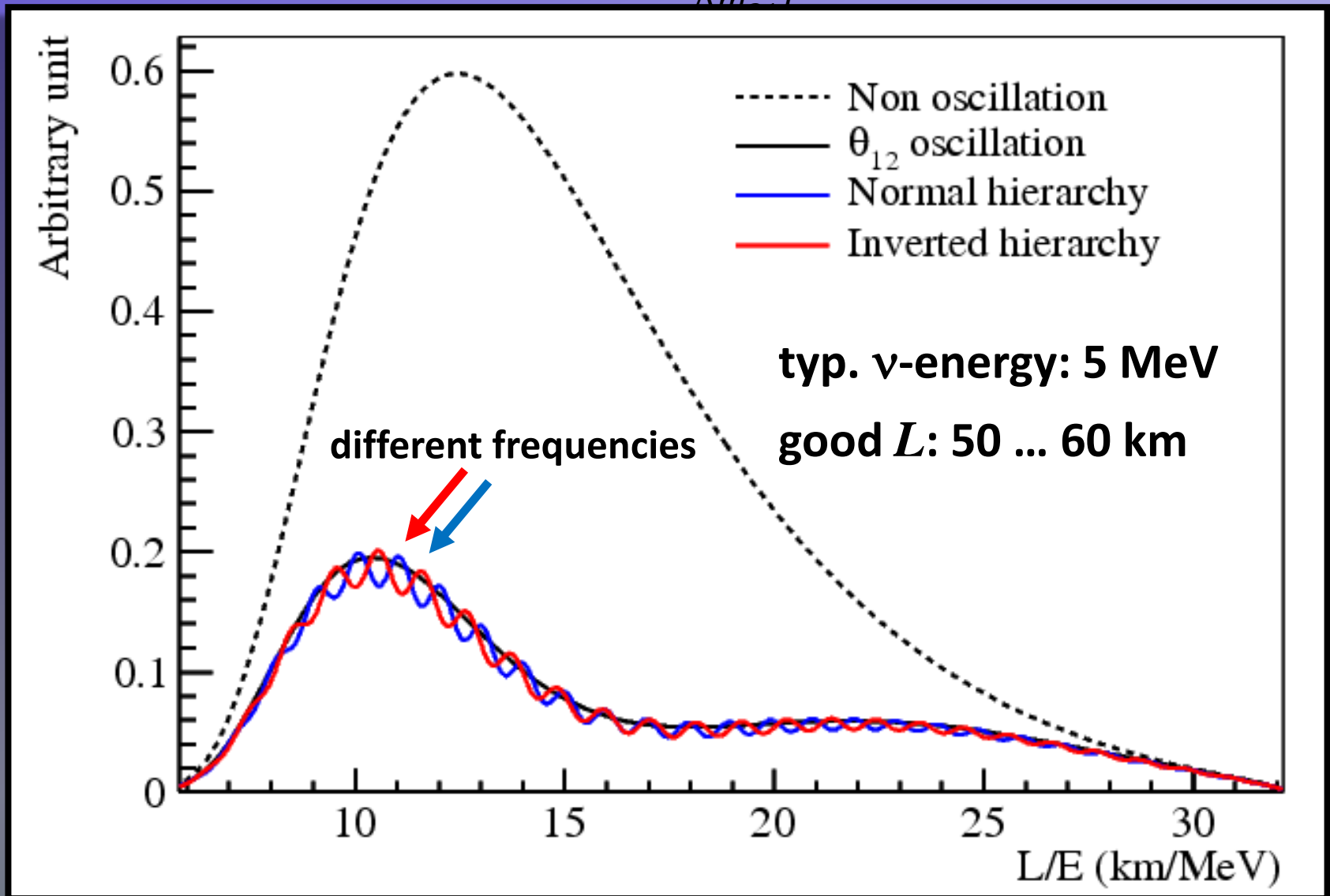
Yu-Feng Li et al.,
Phys.Rev. D88 (2013) 013008

OSCILLATION PATTERN

Yu-Feng Li et al.,
Phys.Rev. D88 (2013) 013008

leading terms:

$$\Delta m^2 \cdot L$$



EXPERIMENTAL LANDSCAPE

Matter-Effects:

approval
status

1. atmospheric neutrinos

PINGU



4 σ

ORCA



4 σ

INO



2 σ

2. beam neutrinos

Nova



2 σ

DUNE/LBNE



>5 σ

3-Flavour Interference:

JUNO



3-4 σ

RENO50



?



The JUNO Project



THE JUNO PROJECT



550 scientists, 70 institutions, 1/3 from Europe



Jiangmen Underground Neutrino Observatory

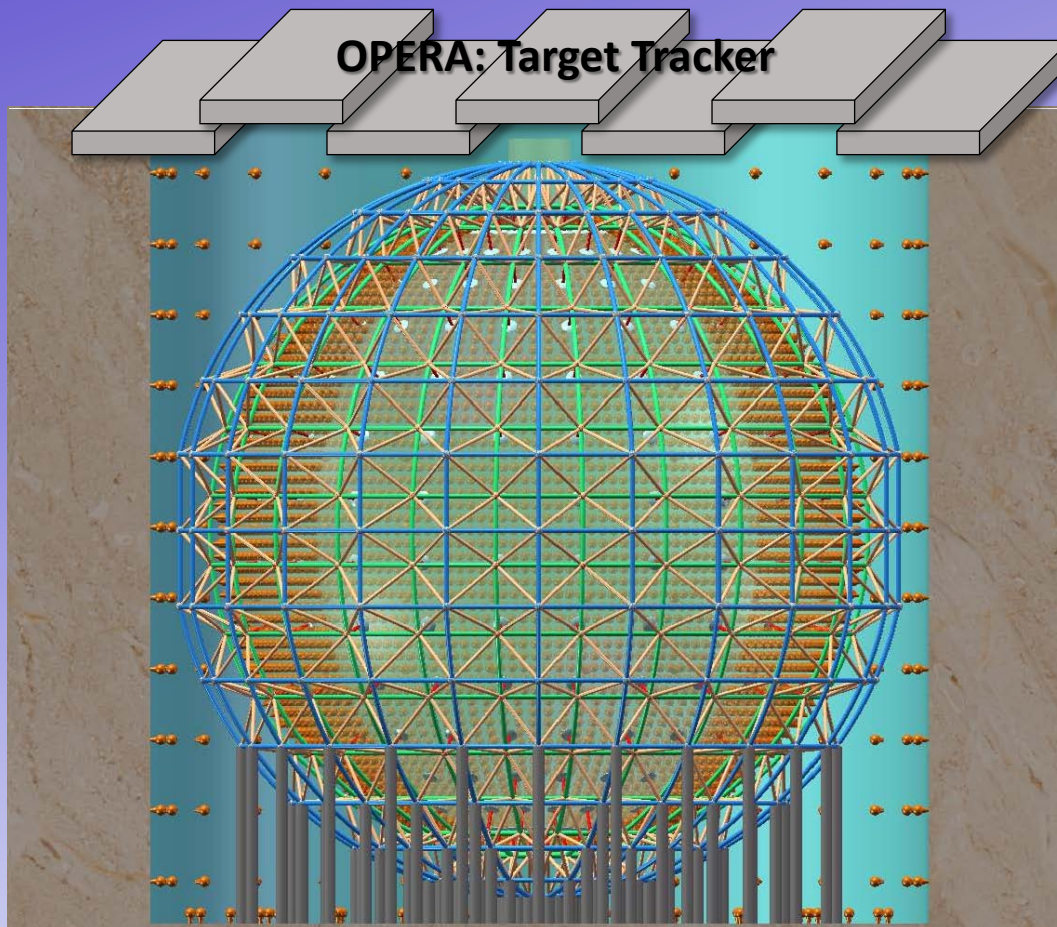
supported by



Armenia, Belgium, Brazil, Chili, Chinese Republic, Czech Republic, Germany, Finland, France, Italy, Latvia, Pakistan, Russia, Slovakia, Thailand, Taiwan, and the United States



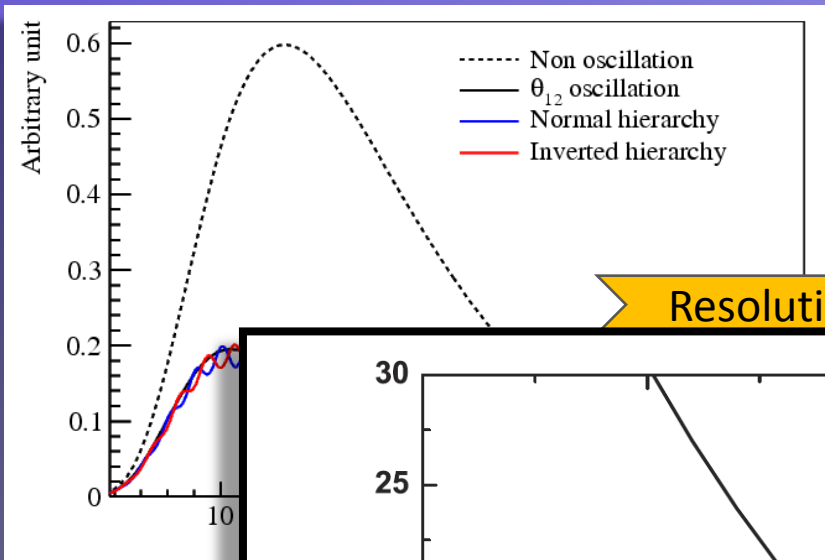
THE JUNO PROJECT



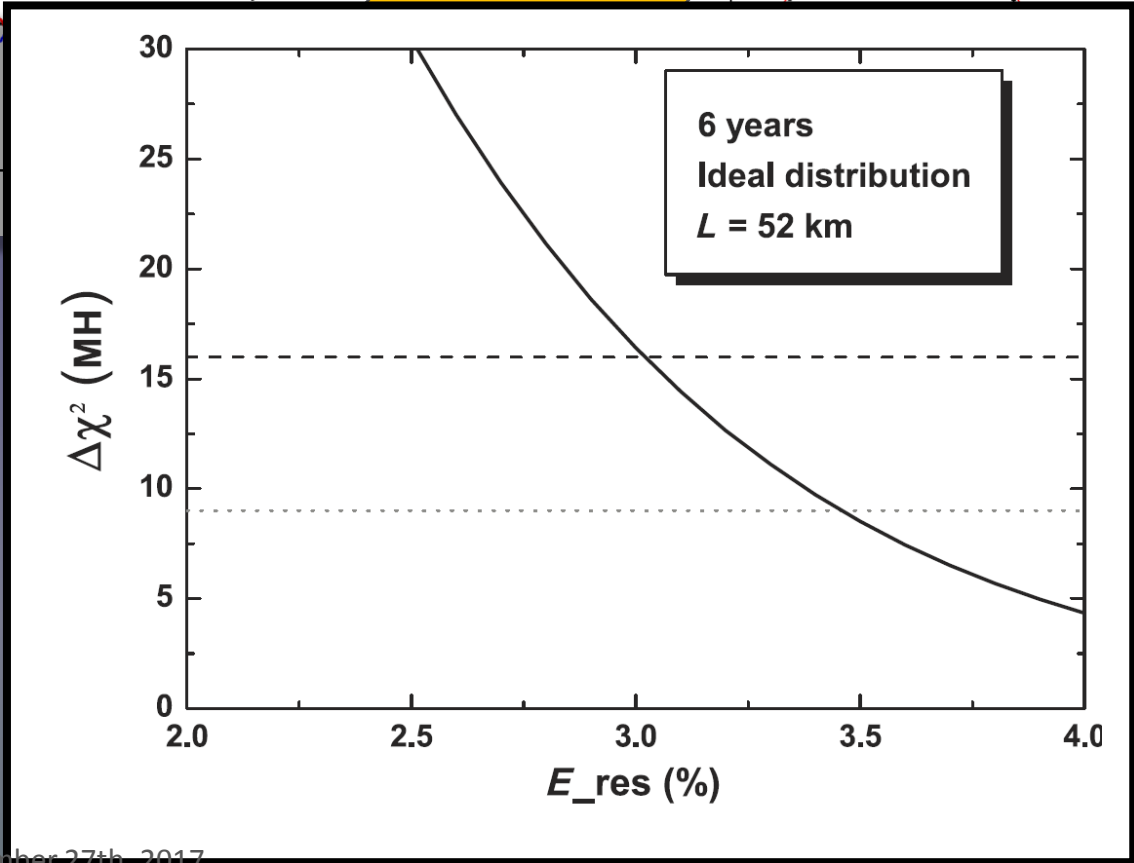
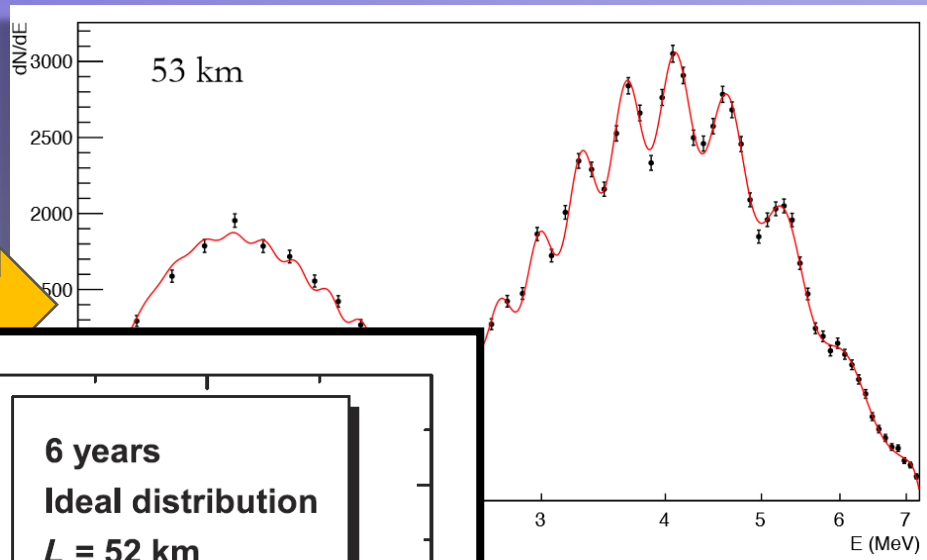
Liquid Scintillator

Ultra-high purity
(BOREXINO technology)
20.000 t fiducial volume
acrylic sphere ($\varnothing 35.5$ m)
2m water buffer
20.000 PMTs (20")
embedded in a water
Čerenkov veto
Muon tracker on top

THE CHALLENGE



Resolution



LENA: 7%/√E
KamLand: 5%/√E
JUNO: 3%/√E

THE CHALLENGE

Excellent Energy Resolution (3% @ 1 MeV)

Photonstatistics

- high lightyield
- good transparency
- PMT-coverage
- PMT-DE

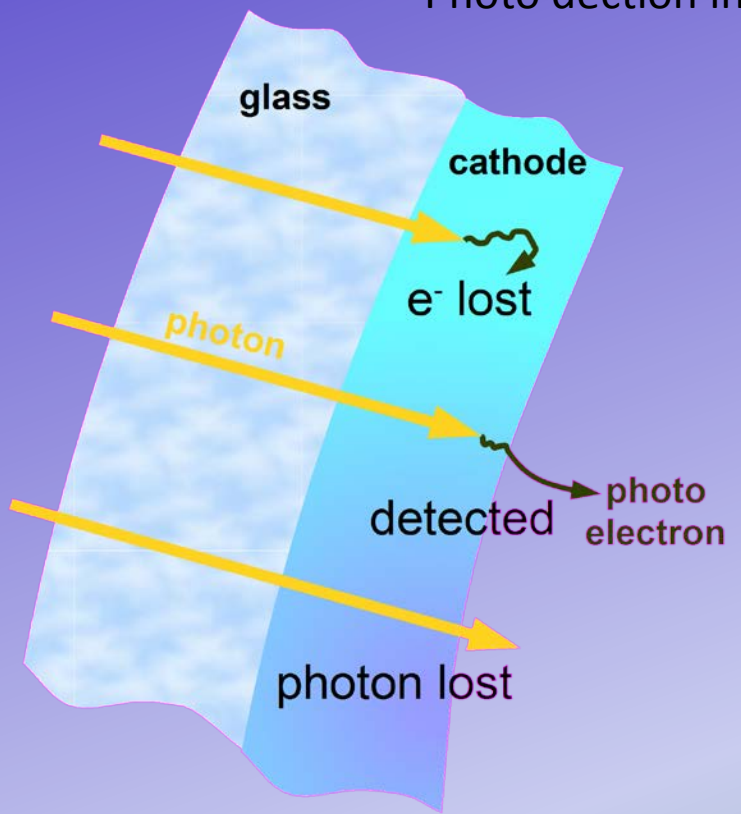
&

Calibration

- $\alpha/\beta/\gamma$ sources
(in all positions)
- light pulsers
(in all positions)
- UV-laser
(in many positions)
- e^+ beam
(along axis)

The MCP-PMT

Photo dection in the cathode

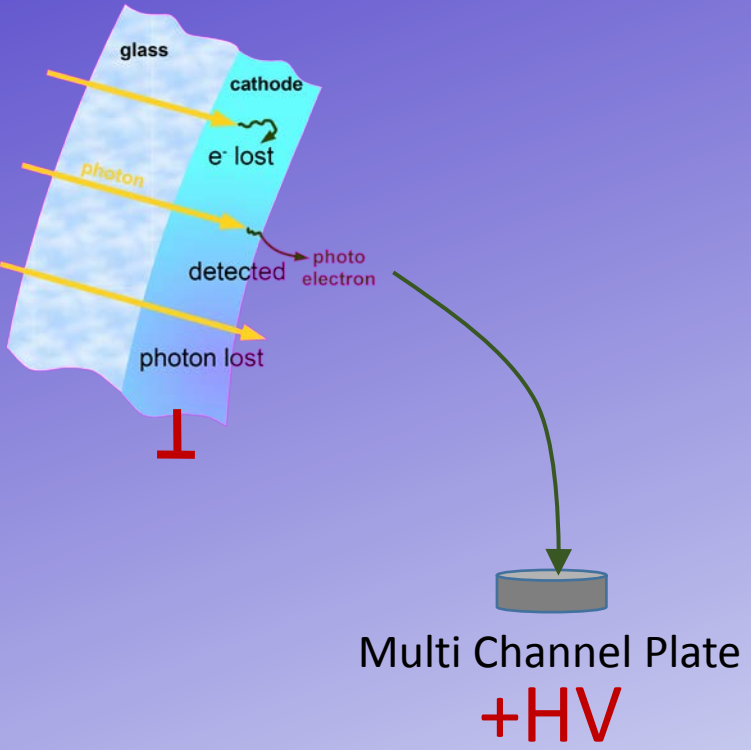


Hamamatsu R12860 (20" PMT)

Detection efficiency =
quantum efficiency
x collection efficiency
x area coverage

Typ. 27%
Spec. > 24%

The MCP-PMT



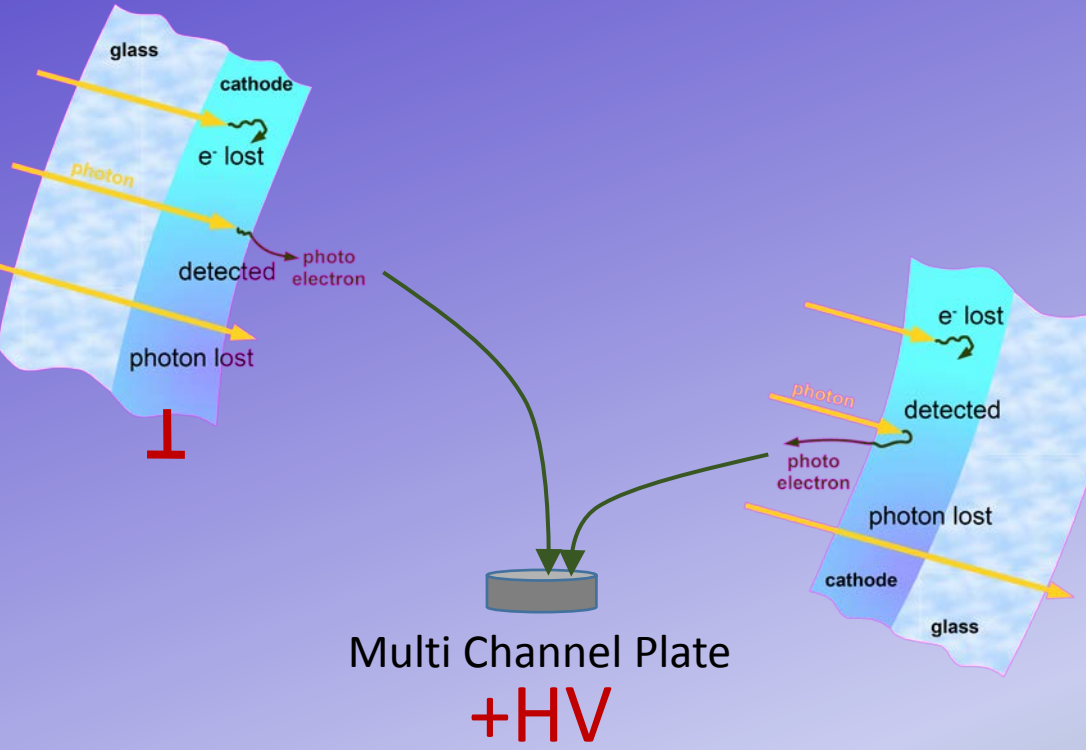
Detection efficiency =
quantum efficiency
x collection efficiency
x area coverage

Typ. 27%
Spec. > 24%



MCP-PMT 8" prototype

The MCP-PMT



Detection efficiency =
quantum efficiency
x collection efficiency
x area coverage

Typ. 27%
Spec. > 24%



MCP-PMT 8" prototype

The Site

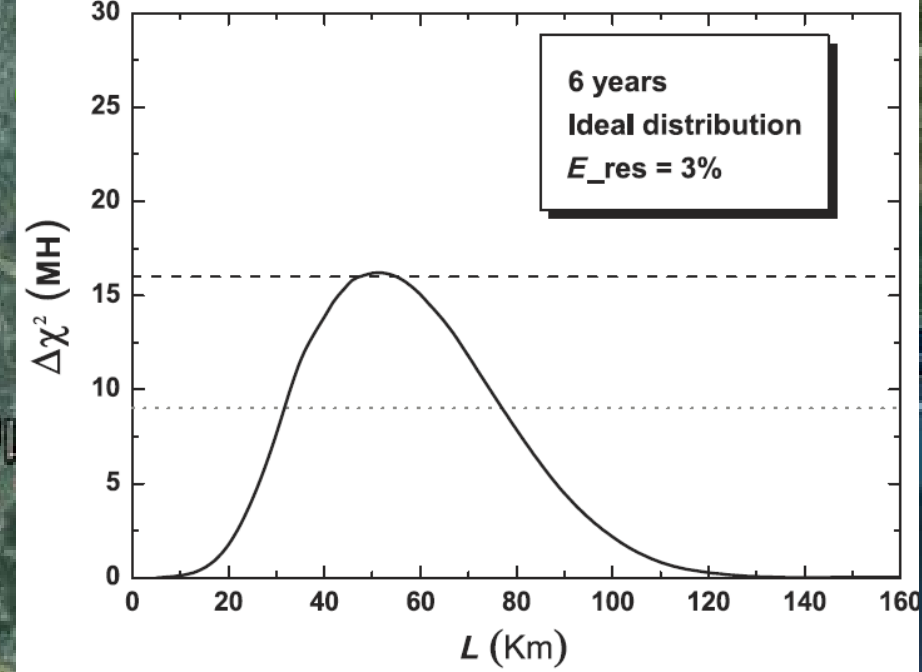
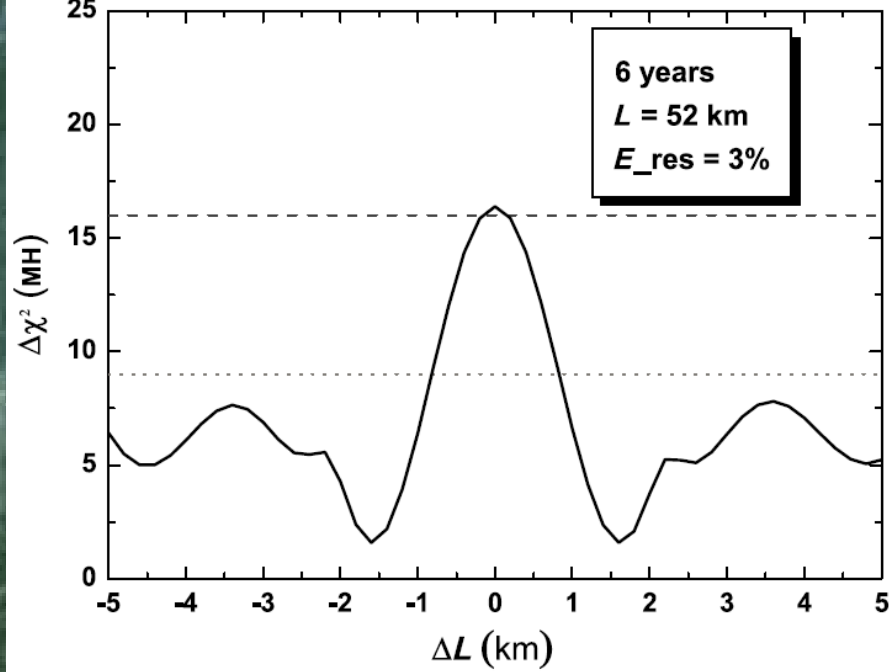
Guangdong province
Jiangmen prefecture
Kaiping city



China

- International Boundary
- Province Boundary
- Road
- River
- ★ National Capital
- ⊙ Province Capital
- City or Town

0 250 500 KM
0 250 500 Miles





n



ivaco



Hongkong



THE SURFACE LAB

江门中微子实验站配套基建工程整体鸟瞰图



Future: 2018

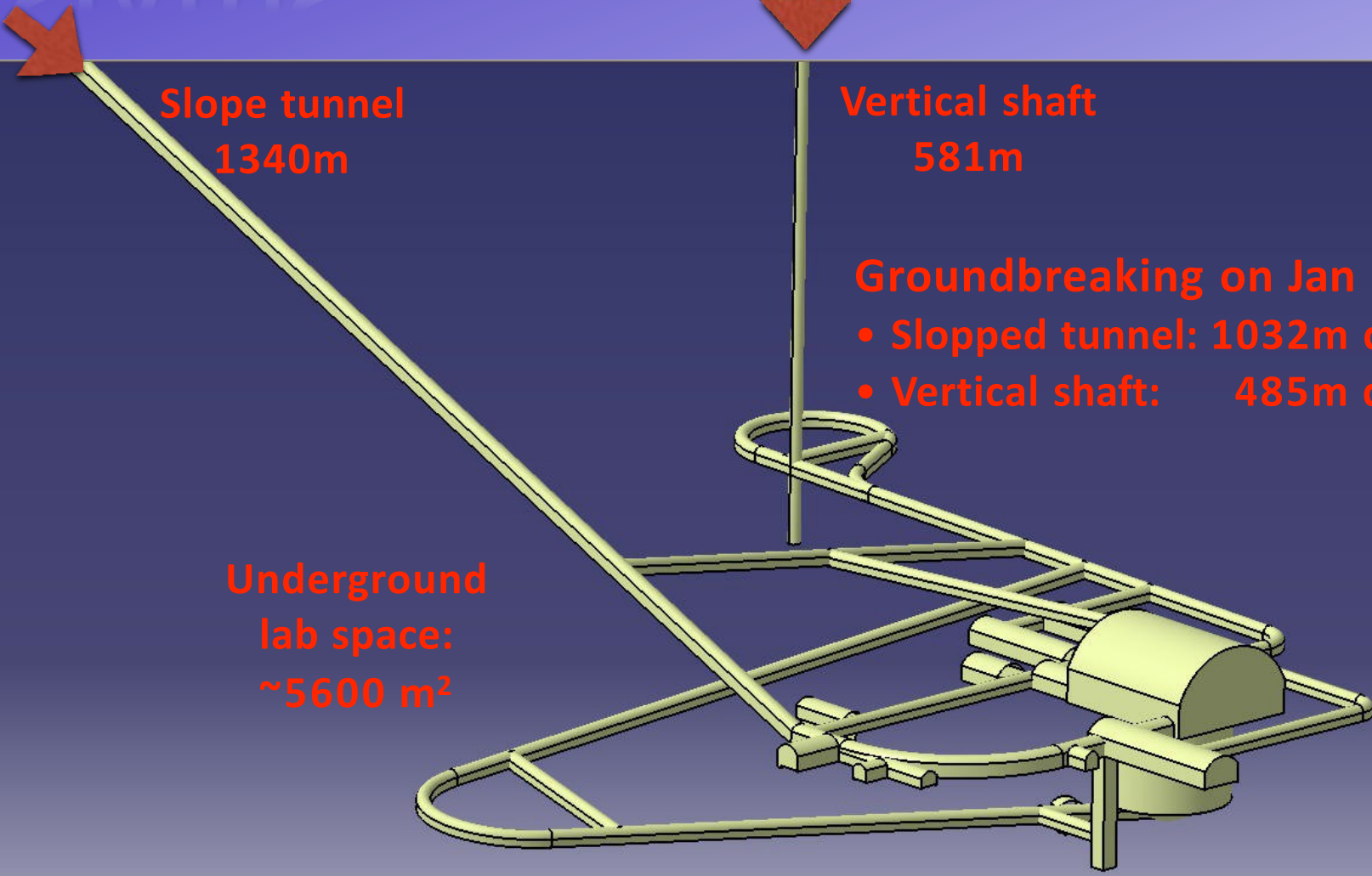
黄河勘测规划设计有限公司

January 2010

January

STATUS

Civil engineering: Completion July 2018



Slope tunnel
1340m

Vertical shaft
581m

Groundbreaking on Jan 10, 2015

- Sloped tunnel: 1032m done
- Vertical shaft: 485m done

Underground
lab space:
~5600 m²



THE JUNO PROJECT



start datataking in 2020

mass hierarchy to 3...4 sigma in 6 years

Physics of JUNO

Mass Hierarchy

MC-studies:

>3 sigma in 4 years

(3% resolution @ 1 MeV)

ν -oscillations with reactor neutrinos:

Mass hierarchy

Precision Measurements

Others

Super Nova

- Direct observation
- Diffuse Super Nova background

Solar Neutrinos

- Oscillation parameters
- Metallicity

Atmospheric Neutrinos

- Oscillations
- Mass hierarchy ?

Geo Neutrinos

- Models of the earth's interior
- Heat production \rightarrow climate

Nucleon Decay

- i.e. $p \rightarrow K^+ \nu$

Dark Matter

- $\chi \rightarrow \nu\nu$

Sterile Neutrinos

- With radioactive sources

Physics of LS-Detectors

Others

LENA @ Phyäsalmi

DETECTOR LAYOUT

Cavern

height: 115 m, diameter: 50 m
shielding from cosmic rays: ~4,000 m.w

Muon Vet

plastic scint
Water Chere
1,500 photo
100 kt of wa
reduction of
neutron bac

Steel Cyl

height: 100 m
70 kt of orga
13,500 phot

Buffer

thickness: 2
non-scintillating organic liquid
shielding external radioactivity

Nylon Vessel

parting buffer liquid
from liquid scintillator

Target Volume

height: 100 m, diameter: 26 m
50 kt of liquid scintillator

LENA

50 kt LS

1400 m overburden

> 200 km to next reactor

7% resolution @ 1 MeV

Others

JUNO @ Jiangmen

JUNO

20 kt LS

700 m overburden

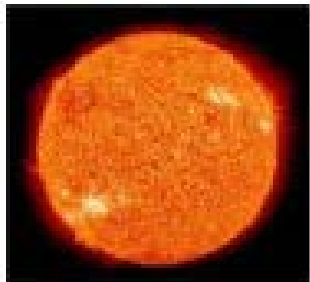
35 GW at 55 km

3% resolution @ 1 MeV



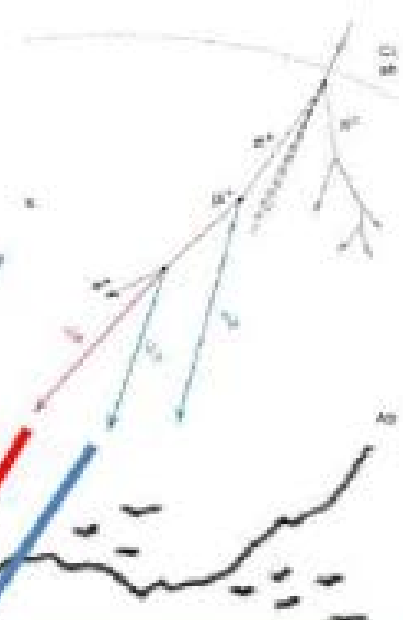
Neutrino Rates

Supernova ν
 $\sim 5k$ in 10s for 10kpc



Solar ν
(10-1000)/day

Atmospheric ν
several/day



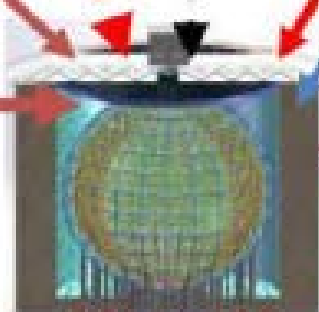
Cosmic muons
 $\sim 250k$ /day

0.003 Hz/m², 215 GeV
10% multiple-muon

36 GW, 53 km

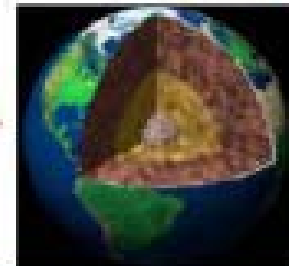


reactor ν , ~ 60 /day



20k ton LS

Geo-neutrinos
1-2/day



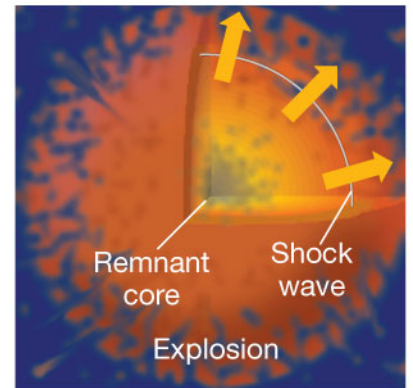
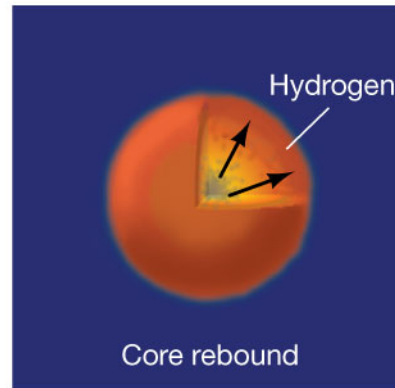
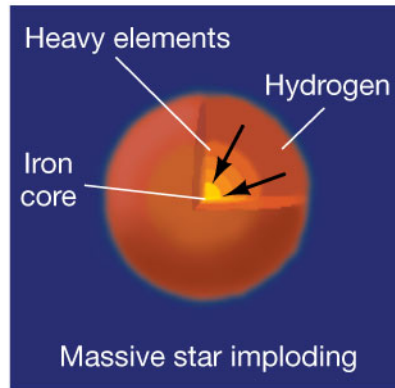
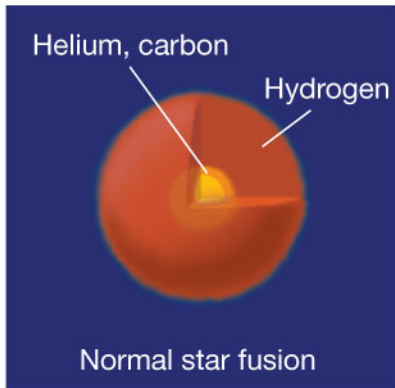
Super Nova

What will we detect?

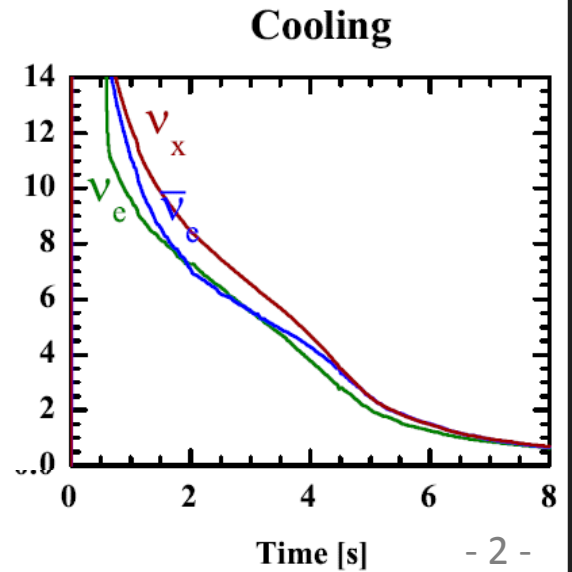
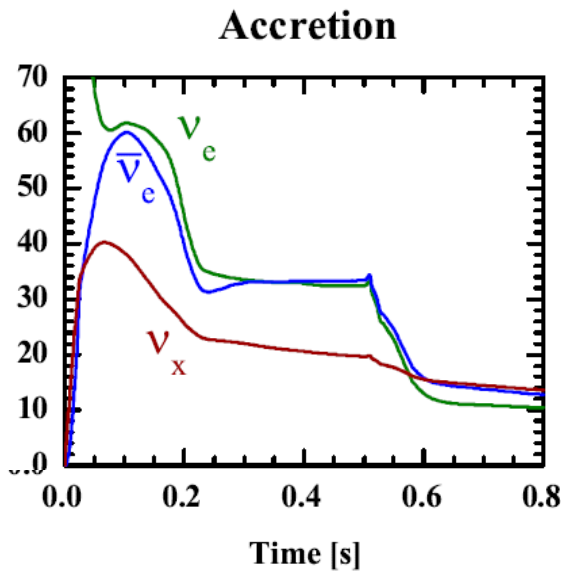
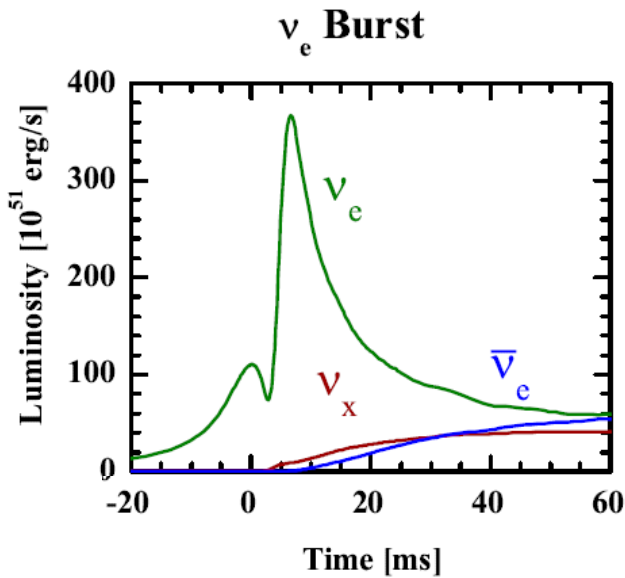
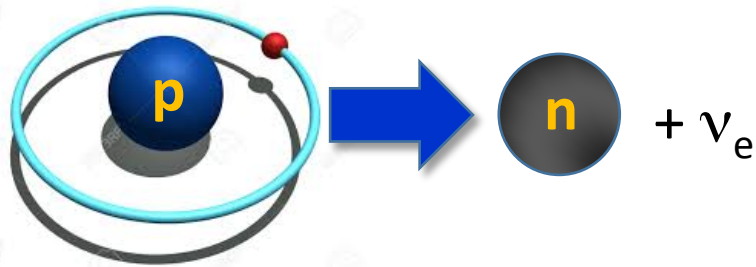
What can we learn about super novae?



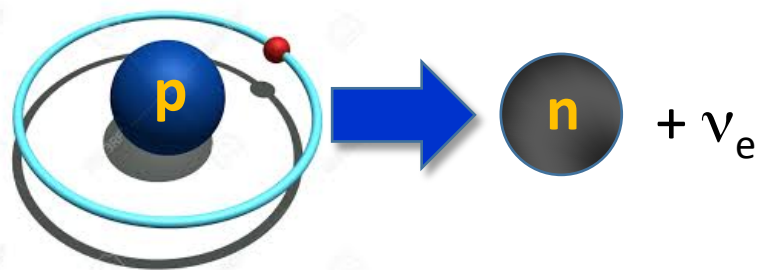
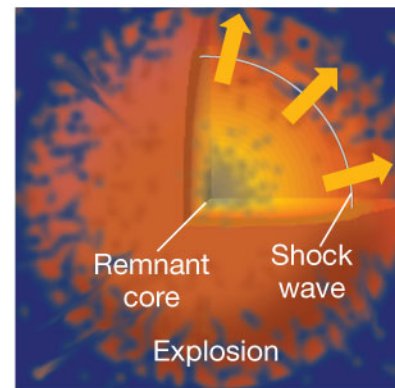
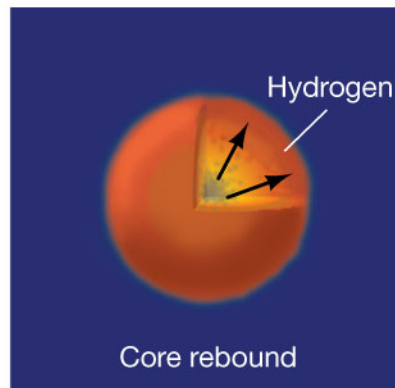
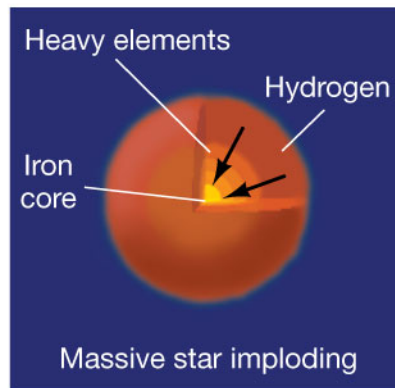
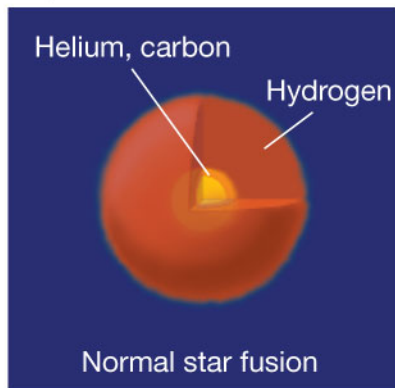
(b) Type II Supernova



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(b) Type II Supernova



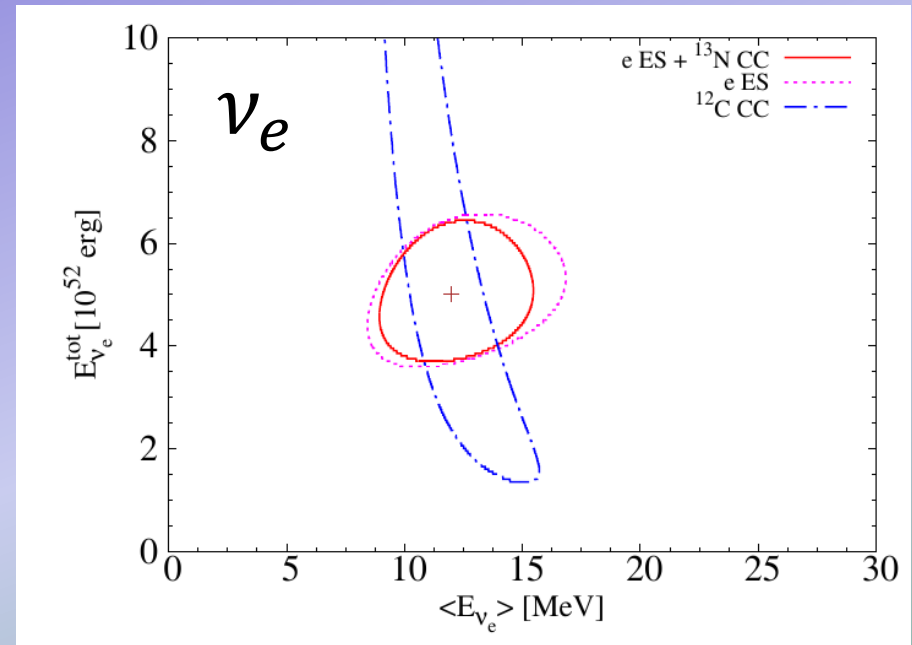
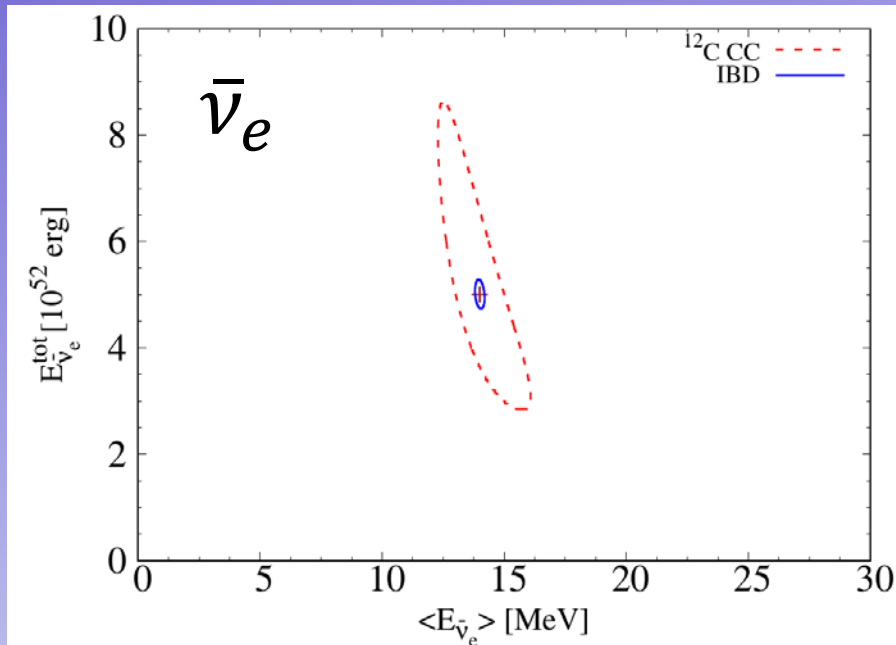
10
kpc

Channel	Type	Events for different $\langle E_\nu \rangle$ values		
		12 MeV	14 MeV	16 MeV
$\bar{\nu}_e + p \rightarrow e^+ + n$	CC	4.3×10^3	5.0×10^3	5.7×10^3
$\nu + p \rightarrow \nu + p$	NC	0.6×10^3	1.2×10^3	2.0×10^3
$\nu + e \rightarrow \nu + e$	ES	3.6×10^2	3.6×10^2	3.6×10^2
$\nu + {}^{12}\text{C} \rightarrow \nu + {}^{12}\text{C}^*$	NC	1.7×10^2	3.2×10^2	5.2×10^2
$\nu_e + {}^{12}\text{C} \rightarrow e^- + {}^{12}\text{N}$	CC	0.5×10^2	0.9×10^2	1.6×10^2
$\bar{\nu}_e + {}^{12}\text{C} \rightarrow e^+ + {}^{12}\text{B}$	CC	0.6×10^2	1.1×10^2	1.6×10^2

Neutrino Spectrum

Type IIa; standard parameters; 10 kpc

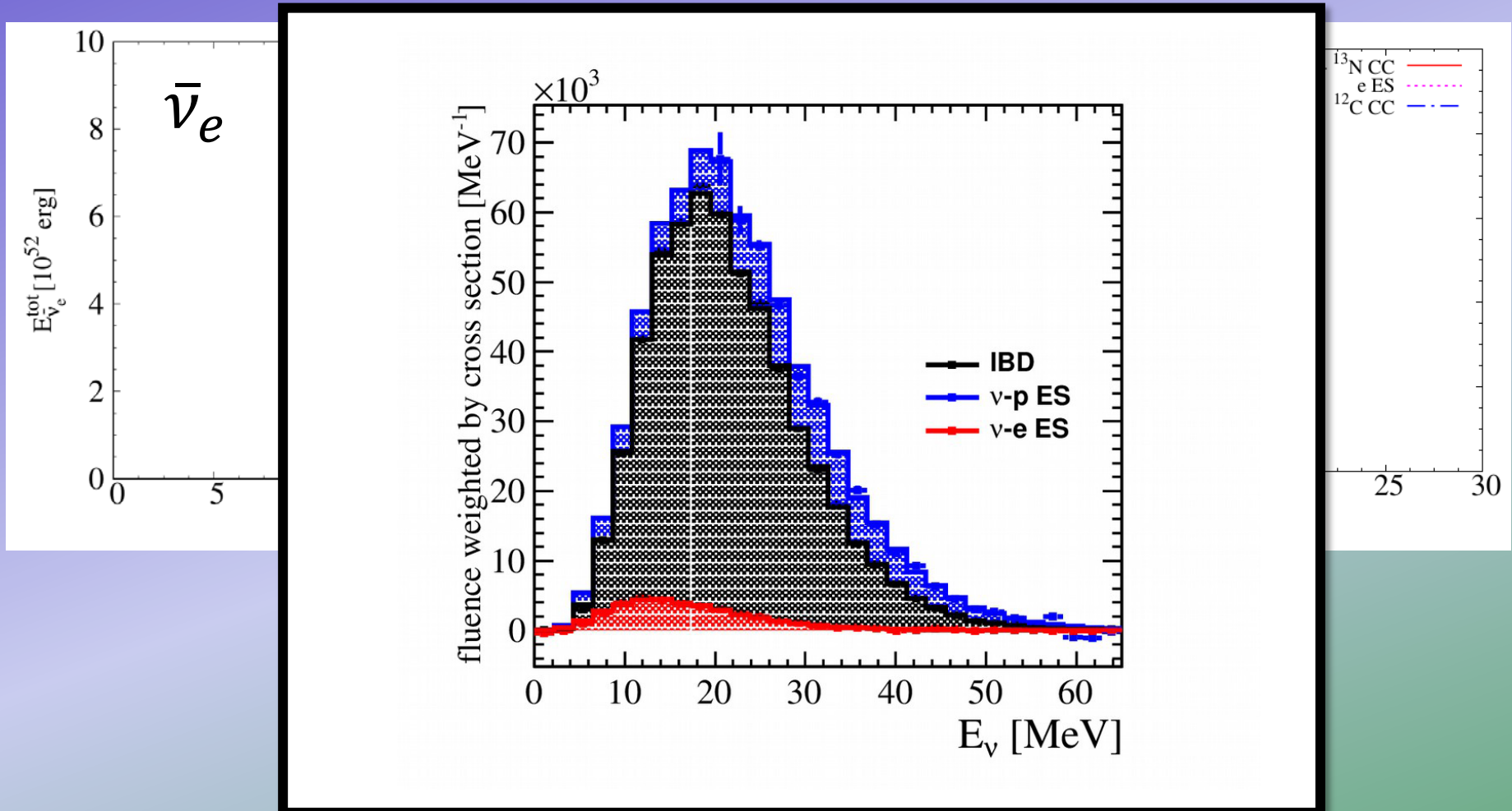
J.-S. Lu et al., Phys.Rev. D94 023006



Neutrino Spectrum

Type IIa; standard parameters; 10 kpc

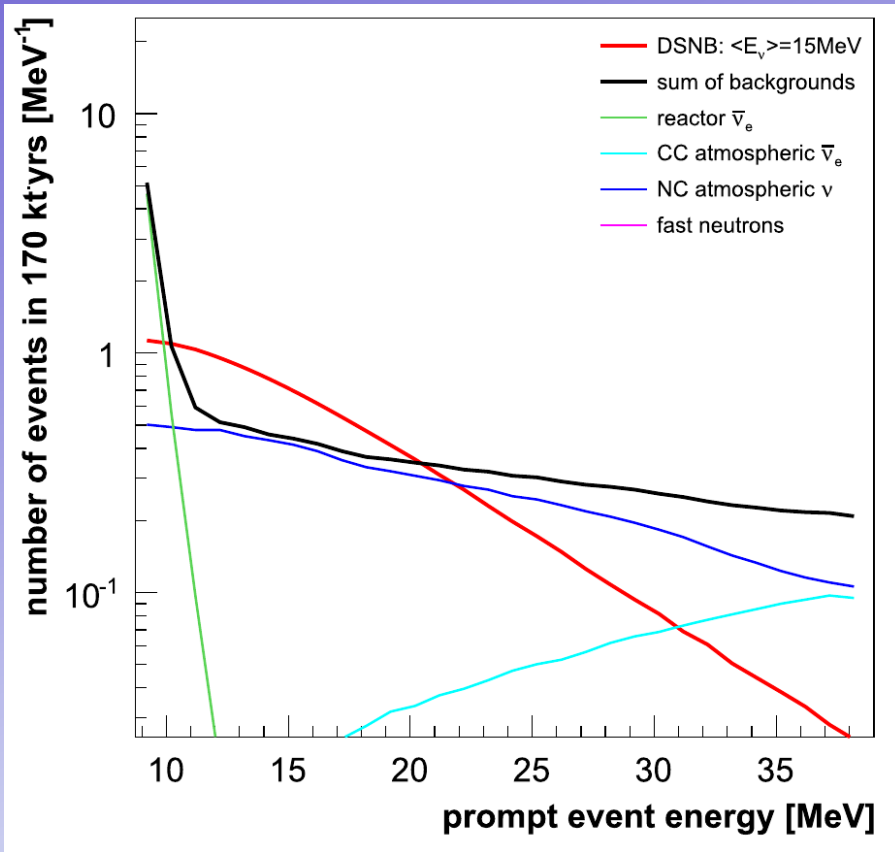
J.-S. Lu et al., Phys.Rev. D94 023006



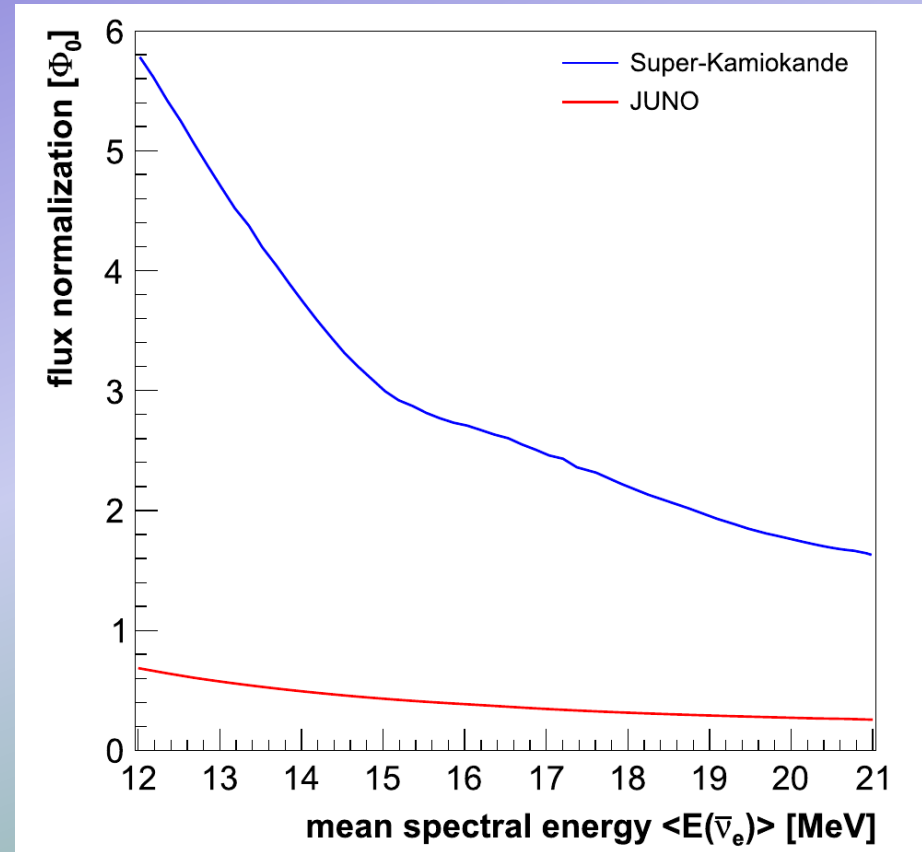
Diffuse Supernova Neutrino Background

Averaged neutrino signal from all supernovae in the universe

spectrum

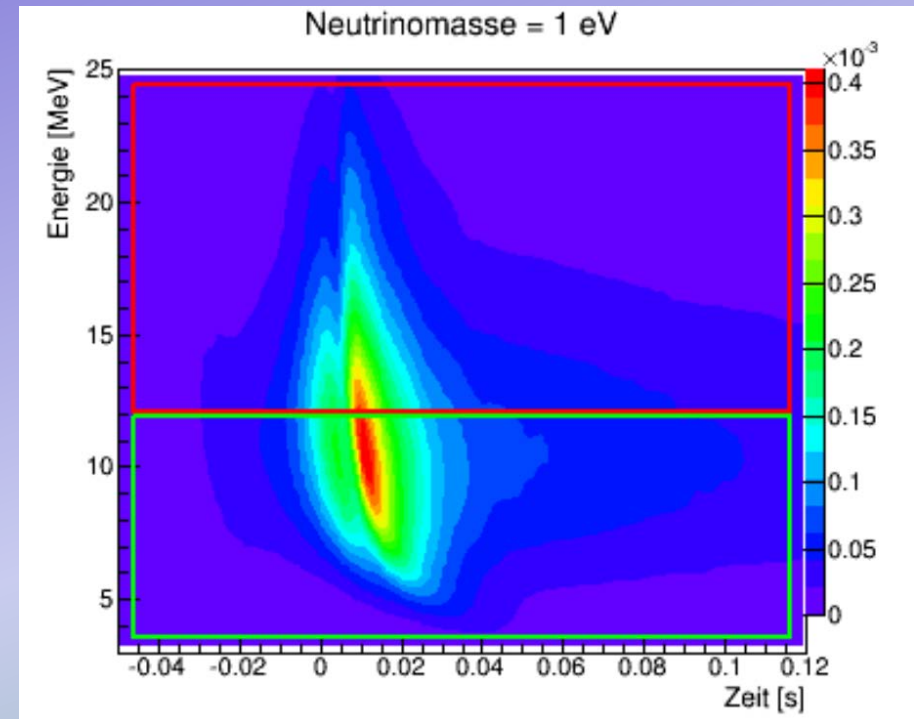
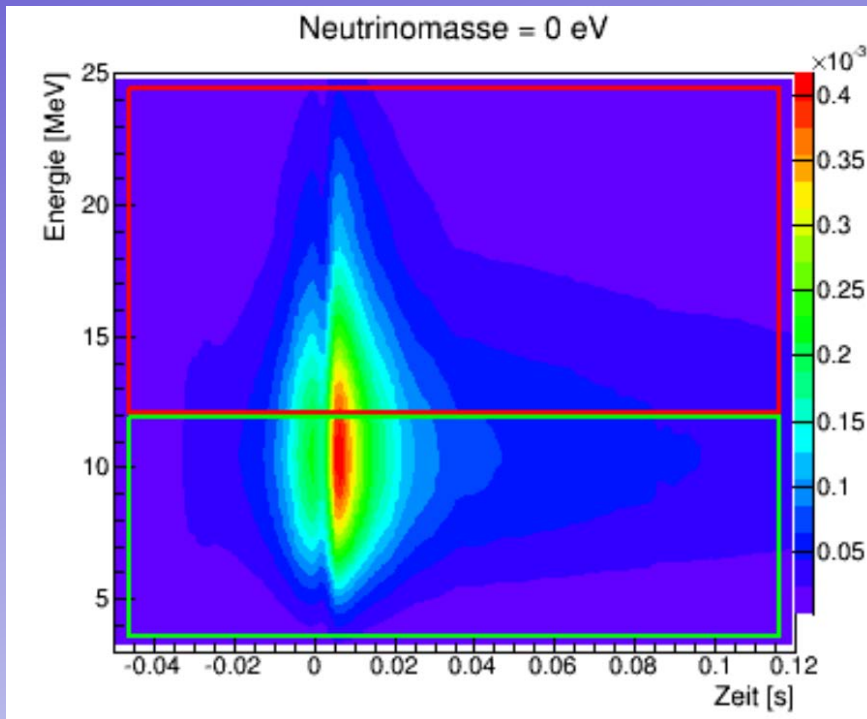


flux



Neutrino Mass from Time-of-Flight

Type Ia; standard parameters; 10 kpc



- Sensitivity ~ 1 eV
- Independent of distance



Conclusions



Conclusions

- Neutrino Physics is a very active field

Conclusions

T2K running with anti-neutrinos

$2\beta 0\nu$: GERDA with strong German contribution

IceCube: cosmic ν and more

Nova just started

Waiting for first results from KATRIN

Japan: HyperK?

Many activities on sterile ν : SOX, Soli δ , ...

US long baseline approaching approval

Will PINGU/ORCA come?

LENA: not forgotten!

Conclusions

- Neutrino Physics is a very active field
- Oscillation parameters reaching % region
- The precision on θ_{13} is still improving
- Mass hierarchy is the next step
- JUNO has been approved.
Construction started Jan. 2015

Colloquium – DESY Zeuthen



Thanks

	Normal Ordering (best fit)		Inverted Ordering ($\Delta\chi^2 = 0.83$)		Any Ordering
	bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range	3σ range
$\sin^2 \theta_{12}$	$0.306^{+0.012}_{-0.012}$	$0.271 \rightarrow 0.345$	$0.306^{+0.012}_{-0.012}$	$0.271 \rightarrow 0.345$	$0.271 \rightarrow 0.345$
$\theta_{12}/^\circ$	$33.56^{+0.77}_{-0.75}$	$31.38 \rightarrow 35.99$	$33.56^{+0.77}_{-0.75}$	$31.38 \rightarrow 35.99$	$31.38 \rightarrow 35.99$
$\sin^2 \theta_{23}$	$0.441^{+0.027}_{-0.021}$	$0.385 \rightarrow 0.635$	$0.587^{+0.020}_{-0.024}$	$0.393 \rightarrow 0.640$	$0.385 \rightarrow 0.638$
$\theta_{23}/^\circ$	$41.6^{+1.5}_{-1.2}$	$38.4 \rightarrow 52.8$	$50.0^{+1.1}_{-1.4}$	$38.8 \rightarrow 53.1$	$38.4 \rightarrow 53.0$
$\sin^2 \theta_{13}$	$0.02166^{+0.00075}_{-0.00075}$	$0.01934 \rightarrow 0.02392$	$0.02179^{+0.00076}_{-0.00076}$	$0.01953 \rightarrow 0.02408$	$0.01934 \rightarrow 0.02397$
$\theta_{13}/^\circ$	$8.46^{+0.15}_{-0.15}$	$7.99 \rightarrow 8.90$	$8.49^{+0.15}_{-0.15}$	$8.03 \rightarrow 8.93$	$7.99 \rightarrow 8.91$
$\delta_{\text{CP}}/^\circ$	261^{+51}_{-59}	$0 \rightarrow 360$	277^{+40}_{-46}	$145 \rightarrow 391$	$0 \rightarrow 360$
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.50^{+0.19}_{-0.17}$	$7.03 \rightarrow 8.09$	$7.50^{+0.19}_{-0.17}$	$7.03 \rightarrow 8.09$	$7.03 \rightarrow 8.09$
$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.524^{+0.039}_{-0.040}$	$+2.407 \rightarrow +2.643$	$-2.514^{+0.038}_{-0.041}$	$-2.635 \rightarrow -2.399$	$\left[+2.407 \rightarrow +2.643 \right]$ $\left[-2.629 \rightarrow -2.405 \right]$

Neutrinos

www.nu-fit.org

NuFIT 3.0 (2016)

$$U|_{3\sigma} = \begin{pmatrix} 0.800 \rightarrow 0.844 & 0.515 \rightarrow 0.581 & 0.139 \rightarrow 0.155 \\ 0.229 \rightarrow 0.516 & 0.438 \rightarrow 0.699 & 0.614 \rightarrow 0.790 \\ 0.249 \rightarrow 0.528 & 0.462 \rightarrow 0.715 & 0.595 \rightarrow 0.776 \end{pmatrix}$$

Quarks

PDG 2017

$$V_{\text{CKM}} = \begin{pmatrix} 0.97434^{+0.00011}_{-0.00012} & 0.22506 \pm 0.00050 & 0.00357 \pm 0.00015 \\ 0.22492 \pm 0.00050 & 0.97351 \pm 0.00013 & 0.0411 \pm 0.0013 \\ 0.00875^{+0.00032}_{-0.00033} & 0.0403 \pm 0.0013 & 0.99915 \pm 0.00005 \end{pmatrix}$$