



# Dark matter, neutrinos & nukes

based on arXiv:2104.13926

#### Quantum Universe Video Colloquium

09-28-2021, DESY and University of Hamburg

B. Cogswell, A. Goel, <u>P. Huber</u> Center for Neutrino Physics at Virginia Tech











## Early history of the neutrino

- Neutrino postulated in Pauli's famous letter in 1930
- Fermi theory of beta decay in 1933
- Bethe and Peierls estimate detection cross section in 1934 to be  $E_{\nu}^2 \, 10^{-43} \, \mathrm{cm}^2$
- and conclude that the neutrino is practically invisible

Physikalisches Institut der Eidg. Technischen Hochschule Zürich

Zürich, 4. Des. 1930 Oloriastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich huldvollst ensuhören bitte, Ihnen des näheren auseinandersetzen wird, bin ich angesichts der "falschen" Statistik der N- und Li-6 Kerne, sowie des kontinuierlichen beta-Spektrums auf einen versweifelten Ausweg verfallen um den "Wechselsats" (1) der Statistik und den Energiesats su retten. Mamlich die Möglichkeit, es könnten elektrisch neutrale Teilchen, die ich Neutronen nennen will, in den Kernen existieren, welche den Spin 1/2 haben und das Ausschliessungsprinzip befolgen und n von Lichtquanten ausserdem noch dadurch unterscheiden, dass sie wit Lichtgeschwindigkeit laufen. Die Masse der Neutronen arte von derselben Grossenordnung wie die Elaktronenmasse sein und jedenfalls nicht grösser als 0,01 Protonennasse -- Das kontinuierliche the Spektrum ware dann verständlich unter der Annahme, dass beim beta-Zerfall mit dem blektron jeweils noch ein Neutron emittiert irde derart, dass die Summe der Energien von Neutron und Elektron konstant ist.

## Early history of neutrino detection

- 1945 Pontecorvo proposes radiochemical detection
- Promptly classified by the Atomic Energy Commission
- Because of the potential for detection of nuclear submarines!



Report PD-141, National Research Council of Canada, Division of Atomic Energy, Chalk River

We will now discuss processes produced by "free neutrinos", i.e. processes produced by neutrinos after they have been emitted in a /3 disintegration. According to Fermi's B-ray theory, there is one process which free neutrinos must produce: the inverse of a  $\beta$  transformation, consisting of the concomitant absorption of a neutrino and emission of a A particle (positive or negative) by a neucleus. The inverse B-ray process is a typical effect produced by neutrinos, if neutrinos exist at all. The cross-section for such nuclear disintegration produced by neutrinos, however, is expected to be extremely small - according to Fermi's theory less than 10-42 sq. cm. It has been currently stated in the literature that it is impossible to observe such a process. The object of this note is to show that the experimental observation of an inverse /3 process is not out of the question, and to suggest a method which might make an experimental observation feasible.

#### CI-37+neutrino $\rightarrow$ Ar-37+electron

#### How to beat $10^{-42}$ cm<sup>2</sup>.

Can get  $10^{24}$  from Avogadro's number...

Where to get  $10^{18}$  neutrinos?

Need  $10^{18}$  beta decays...





On average 6 neutrinos per fission or  $10^{24}$  neutrinos from fissioning 1 kilogram of U-235

#### How neutrinos might be discovered

Cowan and Reines proposed this and got approved!

Their day job was to instrument atmospheric nuclear tests.



#### How neutrinos were really discovered





Delayed coincidence allowed

1956

them to use a reactor, instead

All reactor neutrinos to date have been detected based on the same principles.

## On to solar neutrinos...

- Helicity allows the CI-37 reaction only for left-handed states, i.e. can not happen at reactors.
- Sun is a neutrino source
- In the 1960s Ray Davis built a solar neutrino experiment at Homestake using Pontecorvo's Cl-37 reaction



Individual Ar-37 atoms are detected by their decay back into CI-37. The half-life is around 30 days, this is a crucial requirement!

### Passive detection of reactor neutrinos?

- Need inverse beta decay
- Need small Q-value and large matrix element (small log ft)
- Half-life needs to be more than weeks but less than years
- Looking through the whole chart of isotopes.

 $^{3}He+\bar{\nu}_{e}\rightarrow^{3}H+e^{+}$ 

- But: 12.3 years half-life
- He-3 is really rare and expensive



#### Charged current reaction will not work!



In the mean time, in the real world

GUANAJAY IRBM LAUNCH SITE

VEHICLE REVETMENTS

NKER

PROB NUCLEAR WARHEAD STORAGE SITE



SECURITY FENCE

STRUCTURE BEING

#### **Estimated Global Nuclear Warhead Inventories 1945-2021**



#### How to control nuclear weapons

• Only very few isotopes can maintain a fast neutron chain reaction

Isotope	235U	233U	239Pu
Half-life	700 Million years	160,000 years	24,000 years
Natural abundance	0.72%	0%	0%
Bare critical mass	52kg	15kg	10kg

# Only U-235 occurs naturally and needs to be enriched



U-233 and Pu-239 need to made in a nuclear reactor





#### NUCLEAR NONPROLIFERATION TREATY (NPT)

Bans the acquisition of nuclear weapons by non-weapon states, including the transfer of knowledge; commits weapon states to disarm; verified by IAEA safeguards

1970s



COMPREHENSIVE TEST BAN TREATY (CTBT)

Bans all nuclear explosions in all environments; verified by IMS and CTBTO; not yet entered into force, some states like the U.S. have not yet ratified it

1990s



PLUTONIUM MANAGEMENT AND DISPOSITION AGREEMENT (PMDA) Both Russia and the U.S. commit to dispose of 34 MT each of weapons-usable Pu-239 declared excess to their military needs; currently suspended



TREATY ON THE PROHIBITION OF NUCLEAR WEAPONS (TPNW) Prohibits development, testing, production, stockpiling, transfer, use, and threat of use of nuclear weapons

2021

2010s

Slide courtesy B. K. Cogswell <sup>12</sup>



#### Neutrino reactor monitoring



First proposed by Borovoi & Mikaelyan in 1978 Soviet Atomic Energy 44 (6), 589-592.

Proof of concept at the Rovno power plant in the 1980s Based on inverse beta decay,



### Uses of Neutrinos

Chartered in 2019 by the Office for Defense Nuclear Non-proliferation within the US DOE NNSA

The charge was to engage end-users and to explore via this engagement what area are plausible use cases for neutrinos. Two areas were identified as particular strong cases:



*Advanced Reactors:* Advanced reactors present novel safeguards challenges which represent possible use cases for neutrino monitoring.

*Future Nuclear Deals:* There is interest in the policy community in neutrino detection as a possible element of future nuclear deals involving cooperative reactor monitoring or verifying the absence of reactor operations.

Full report at https://nutools.ornl.gov



#### Neutrino detectors as of today



In 2018 finally neutrino detectors could function **at the surface** and actually would be able to do the above

PROSPECT liquid scintillator can and does leak





CHANDLER, solid scintillator,

Million dollar, ton-scale devices – not a great fit for the IAEA

#### CEvNS

Coherent Elastic Neutrino Nucleon Scattering

$$\frac{d\sigma}{dT} = \frac{G_F^2}{4\pi} N^2 M_N \left(1 - \frac{M_N T}{2E_\nu^2}\right)$$

T recoil energy, N neutron number

Large cross section but only at very low recoil energy



First observed by the COHERENT collaboration in 2017 at the SNS at ORNL

Detection at reactors still outstanding

Bowen, PH, Phys. Rev. D 102 (2020) 053008.

# The concept

- Nuclear recoils damage crystal lattice permanently, either by forming tracks or vacancies
- This allows off-site readout, hence detector is <u>passive</u>
- <u>Intrinsic rejection</u> of ionizing backgrounds, leaves only neutrons
- This has been explored for dark matter detection but not for reactor neutrinos



Essig, Marden, Slone, Volansky Phys.Rev.D 95 (2017) 5, 056011. Rajendran *et al.*, Phys.Rev.D 96 (2017) 3, 035009. Budnik, Chesnovsky, Slone, Volansky, Phys.Lett.B 782 (2018) 242-250. Baum, Drukier *et al.*, Phys.Lett.B 803 (2020) 135325, Phys.Rev.D 99 (2019) 4, 043014, Phys.Rev.D 99 (2019) 4, 043541

#### Solid state nuclear track detectors

- Crystal damage tracks from multi-MeV events
- Made visible by etching
- Mica and special plastics
- Has been used to put limits on magnetic monopoles using ancient mica Price, Phys. Rev. Lett. 56 (1986) 1226-1229.



Vater, et al., NIM 174 (1977) 271-278

## Tracks vs vacancies

- Damage tracks result from distortion of crystal lattice in response to ionization
- Vacancies are point defects in the lattice
- Both are persistent at room temperature in many materials



Example:

100g NaI 20m from a 3GWth reactor, 90 days Sea level cosmogenic neutron flux Simulation with TRIM

# Vacancy formation

- Need big enough separation to prevent recombination
- Main mechanism for radiation damage in e.g. reactor vessels
- We use the TRIM package in full cascade mode

Ziegler, Ziegler, Biersack, NIM B 268 (2010) 1818-23.





## Threshold displacement energy

- Crucial input parameter to TRIM
- Difficult to measure
- Data mostly for metals
- Empirically found to scale with the melting point
- Use this heuristic to obtain values for materials of interest



Konobeyev, *et al.*, Nucl. Energy & Technology, 3 (2017) 169-175

## Test case CaF2

- CaF2 has been studied in detail
- Goal: neutron dosimetry at very high fluences (like inside a reactor)
- Molecular dynamics simulation with LAMPPS
- Strong anisotropy in TDE
- Value for average TDE and resulting recoil threshold within 20-30% of TRIM and melting point heuristics



3 ps

F Interstitial

0.21 ps

F PKA

Morris, Cowen, Teysseyre, Hecht, Comp. Mat. Science, 172 (2020) 109293,

# Color centers

- In ionic crystals (e.g. NaCl), anion vacancies trap electrons
- Quantum mechanics in a square well  $\rightarrow$  distinct energy levels
- Used to artificially change colors of gem stones
- Individual color centers can be seen in visible light by fluorescence spectroscopy (NV in diamond, SiC)
- Observed in a wide class of materials







From: R. Tilley, Encyclopedia of Color Science and Technology, Springer 2013



### Does it work?





Mosbacher et al, arXiv:1902.10668 irradiated a number of materials with neutrons and gammas to study bulk color center formation.

Similar results for Al2O3 and LiF2

Nothing here with 500x higher dose of gammas

# Readout schemes







Volume imaging, but no location of defect.

Problem for preexisting defects!



#### How does this work?



Crystal is scanned prior to deployment to record the location of pre-existing defects (green dots).

Crystal is scanned after deployment and any new defects are candidates for neutrino events (red dots).

Finding the green dots in their correct location verifies that the crystal is indeed the **same** crystal and that no attempt at **erasing** any defects by heating was made.



Tamper proof!

## Nuclear recoil threshold

- CEvNS/DM detection rates are a steep function of the nuclear recoil threshold (NRT)
- NRT is generally higher than ionization or scintillation threshold due to quenching
- NRT is a key performance parameter: sub 100eV for both tracks and vacancies.
- NRT<100eV only achieved by mK cryogenic bolometers.



#### Material selection

	A	$m_A$	density	melting	TDE	$E_{1/2}$	$\operatorname{CEvNS}$	selection	usable
				point			events	efficiency	events
material	[u]	[%]	$[\mathrm{gcm^{-3}}]$	[K]	[eV]	[eV]		[%]	
LiF	19.0	73.2	2.64	1120	27	80	5600	75	4200
$BaF_2$	137.3	78.3	4.88	1625	35	105	48800	20	9600
NaI	126.9	85.0	3.67	935	24	65	46900	32	15100
CsI	132.9	$\sim \! 100$	4.51	900	23	55	55700	37	20600
$CaWO_4$	183.8	63.9	6.06	1895	41	110	55400	8	4600
$\mathrm{Bi}_{12}\mathrm{GeO}_{20}$	209.0	86.4	9.22	1175	28	85	83500	17	14000

- Melting point well above 300K
- Electrical insulator
- Permit color centers
- Color centers selectively formed by nuclear recoil

- Optical quality crystals
- High atomic mass (for CEvNS)
- Low threshold damage energy

•

#### Dark matter detection

LiF2 great target for spin-dependent interactions

Spin-dependent interactions: hard to compete with Ge detectors

Neutron backgrounds negligible well into the 10 kg year exposure range



# **CEvNS** detection

- CEvNS not yet observed at reactors, very active field
- 10g crystals could deliver 5 sigma detection
- Neutron background:
- 1mwe passive shield would be really helpful
- Active shielding not required
- Passive detector, no cryogenics small footprint



# CEvNS for safeguards

- Verification of reactor shutdown
- 90 day, 1SQ criterion
- Several candidate materials
- All commercially available as radio-clean optical quality crystals
- 1 meter water equivalent passive shield needed



# **Mission Relevance**

- Verification of reactor shutdown
- Verification of end of plutonium production
- For smaller reactors: limit on reactor power & duration of operation
- $\rightarrow$  directly supports non-proliferation



#### Nuclear disarmament verification

How to prove two objects are the same without revealing any other information?



Glaser, Borak, Goldstone, Nature 510 (2014) 497-502.

So-called zero knowledge proof – ideal for verifying that an object belongs to the same equivalence class as a template

Application to a warhead could look like this



#### **Detector needs**

Experimental proof-of-concept



Philippe, Glaser, Goldstone, d'Errico, Nature Commun. 7 (2016) 12890.

Uses super-heated emulsion detectors (think PICO)

Detector needs to be

pre-loadable pre-load must be persistent pre-load must be verifiable by host insensitive to gammas non-electronic un-spoofable

Color center-based detectors check all these boxes.

MeV neutrons are much easier to detect than reactor CEvNS.

### **Expected Impact**

- Non-spoofable: neutrino signal unique, distinction to neutrons by different scaling with atomic mass, pre-existing point defects fingerprint a given detector.
- Passive detectors of ~100g likely can be incorporated into safeguards operations without disruption.
- Detector materials are commercially available.
- Light sheet microscopes are commercially available.









# Summary

- Passive CEvNS detection based on color centers seems theoretically feasible based on detailed simulations
- Resulting detector sizes & deployment footprints appear promising:
- No cryogenics, detectors less 1kg
- Possible application to dark matter and neutron detection

#### Thanks to my collaborators:





Dr. Bernadette K. Cogswell

Apurva Goel